

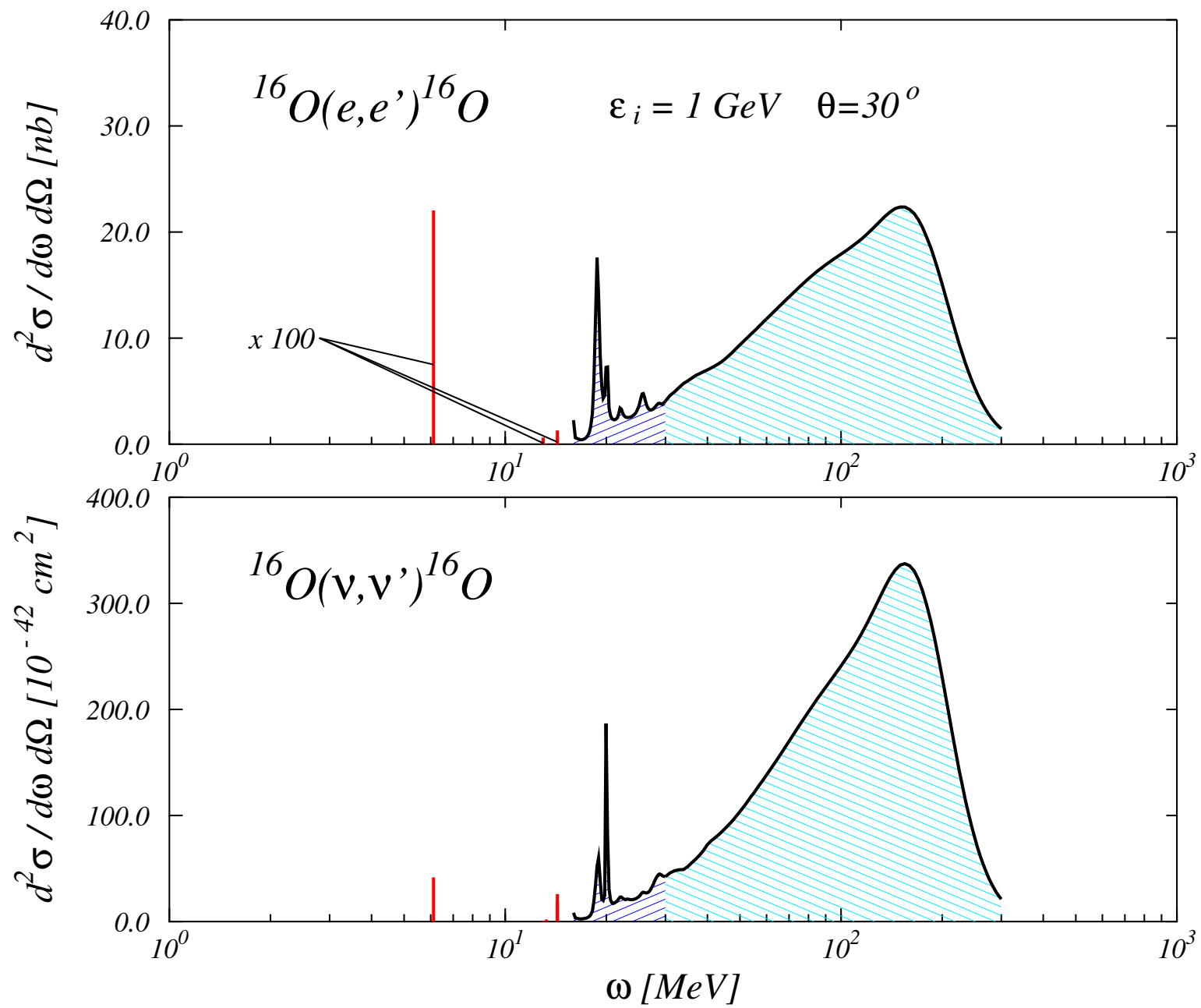
Okayama, Sept. 26-29, 2005

NuInt05

Nuclear Models and Neutrino Cross Sections

Giampaolo Co'

University and INFN, Lecce, ITALY



Random Phase Approximation (RPA)

$$|\Psi_n\rangle = Q_n^\dagger |\Psi_o\rangle \quad Q_n |\Psi_o\rangle = 0$$

$$Q_n^\dagger = \sum_{ph} \textcolor{red}{X}_{ph}^n a_p^\dagger a_h + \sum_{ph} \textcolor{red}{Y}_{ph}^n a_h^\dagger a_p$$

$$W = \sum_{\alpha\beta} w a_\alpha^\dagger a_\beta$$

$$\langle \Psi_n | W | \Psi_o \rangle = \sum_{ph} \textcolor{red}{X}_{ph}^n \langle p | w | h \rangle - \sum_{ph} \textcolor{red}{Y}_{ph}^n \langle h | w | p \rangle$$

Equations of motion

$$\begin{aligned} & \langle \Psi_o | [\delta Q_n^+, [H^{eff}, Q_n^+]] | \Psi_o \rangle \\ &= (E_n - E_o) \langle \Psi_o | [\delta Q_n^+, Q_n^+] | \Psi_o \rangle \end{aligned}$$

RPA secular equations

$$\begin{pmatrix} A & B \\ B^* & A^* \end{pmatrix} \begin{pmatrix} X \\ Y \end{pmatrix} = (E_n - E_o) \begin{pmatrix} X \\ -Y \end{pmatrix}$$

$$A_{ph,p'h'} = (\epsilon_p - \epsilon_h) \delta_{ph,p'h'} + \langle ph' | V^{eff} | hp' \rangle - \langle ph' | V^{eff} | p'h \rangle$$

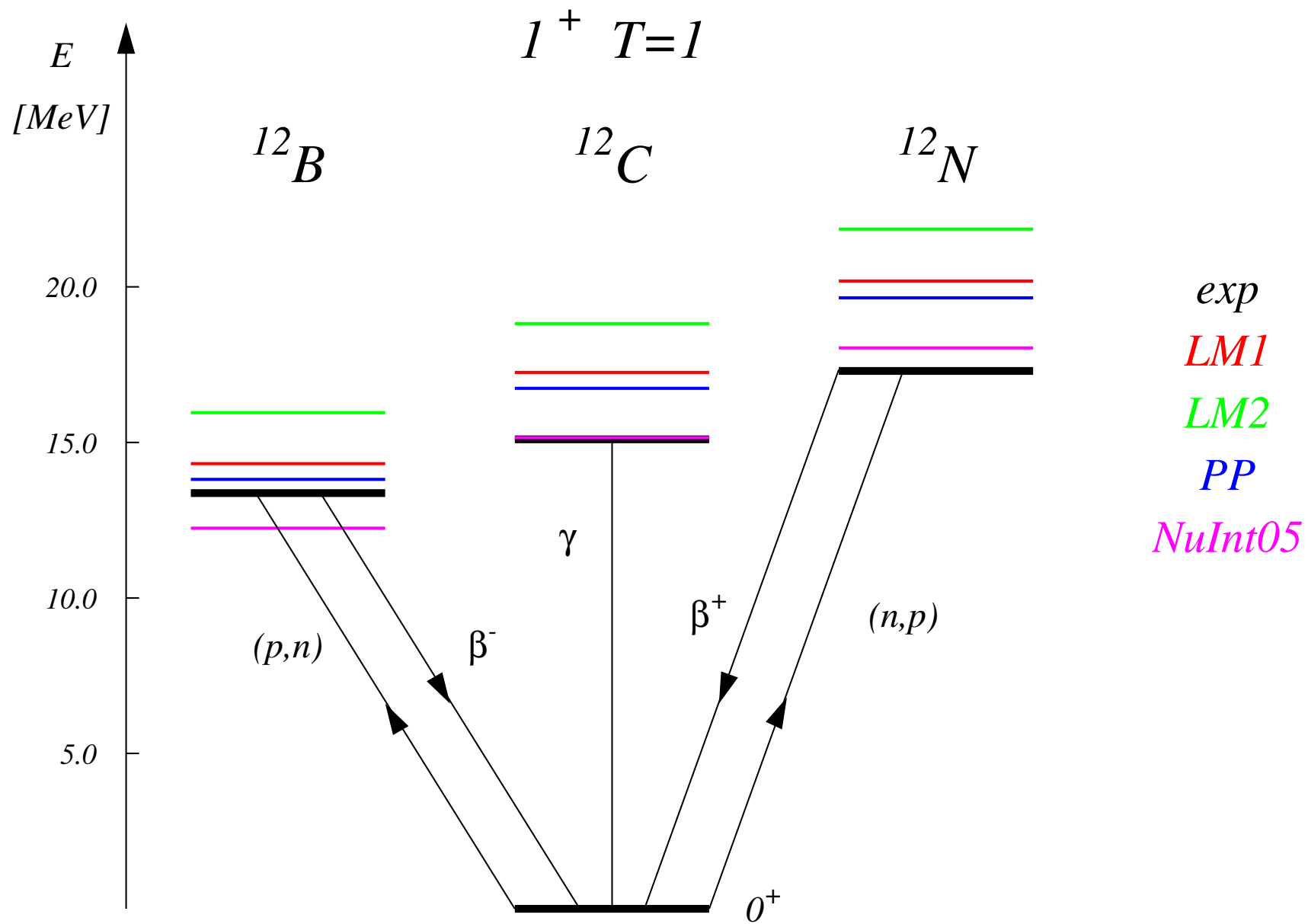
$$B_{ph,p'h'} = \langle pp' | V^{eff} | hh' \rangle - \langle pp' | V^{eff} | h'h \rangle$$

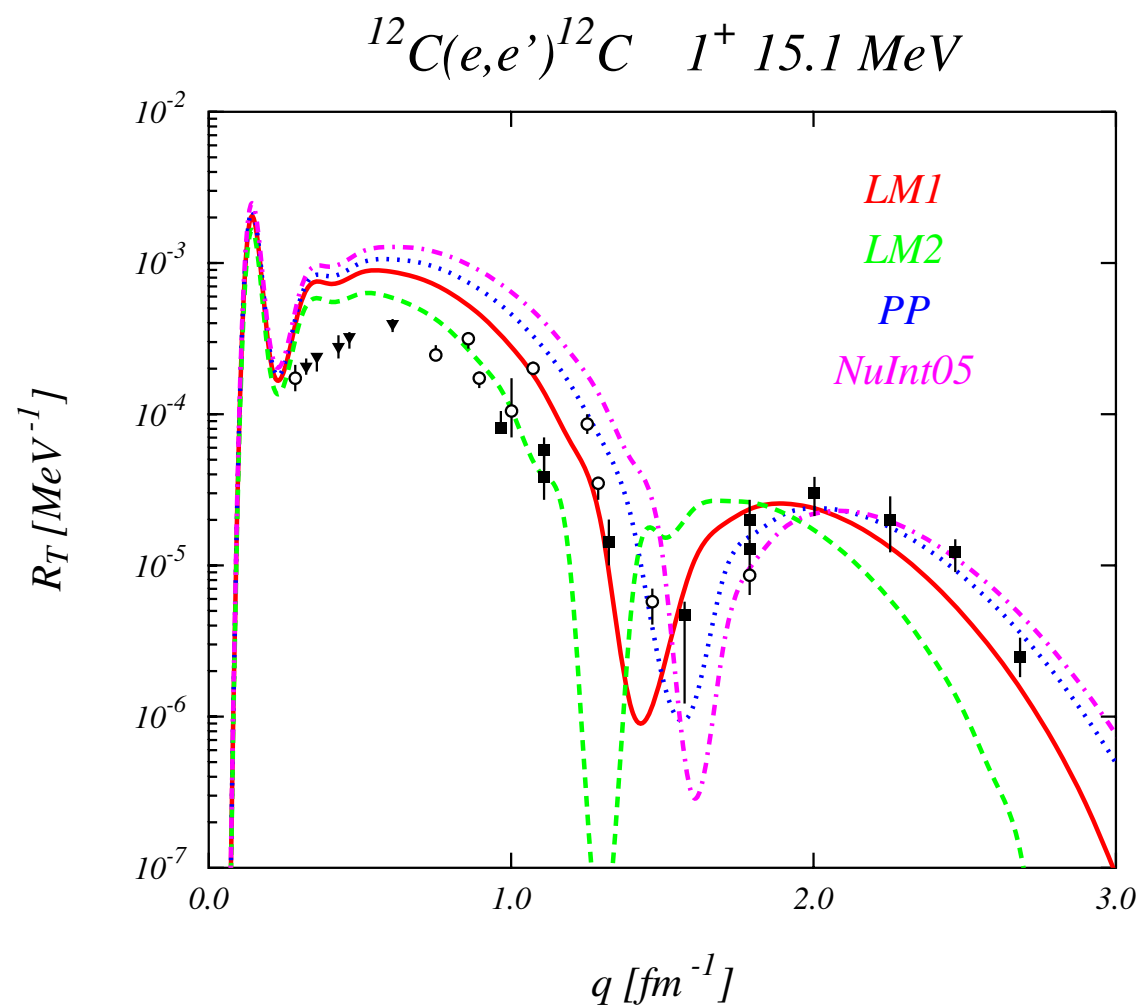
		^{12}C		^{16}O	
	MeV	WS	exp	WS	exp
p	1s1/2	-30.90		-29.90	
	1p3/2	-15.76	-15.96	-16.92	-18.44
	1p1/2	-12.95	-1.94	-12.74	-12.11
	1d5/2	2.0		-3.76	-0.60
	2s1/2			-0.89	-0.10
n	1s1/2	-34.04		-33.98	
	1p3/2	-18.92	-18.72	-20.52	-21.81
	1p1/2	-15.06	-4.96	-16.63	-15.65
	1d5/2	-3.88	-1.10	-6.84	-4.14
	2s1/2	-1.96	-1.86	-3.90	-3.27
	1d3/2	1.6		-3.90	-3.27

$$\begin{aligned}
V^{eff}(k, \rho) = & \mathcal{K} [F(k, \rho) + F'(k, \rho) \vec{\tau}(1) \cdot \vec{\tau}(2) \\
& + G(k) \vec{\sigma}(1) \cdot \vec{\sigma}(2) \\
& + G'(k) \vec{\sigma}(1) \cdot \vec{\sigma}(2) \vec{\tau}(1) \cdot \vec{\tau}(2)]
\end{aligned}$$

$$F(k, \rho) = F^{ext}(k, \rho) + \left[F^{int}(k, \rho) - F^{ext}(k, \rho) \right] \frac{\rho(r)}{\rho(0)}$$

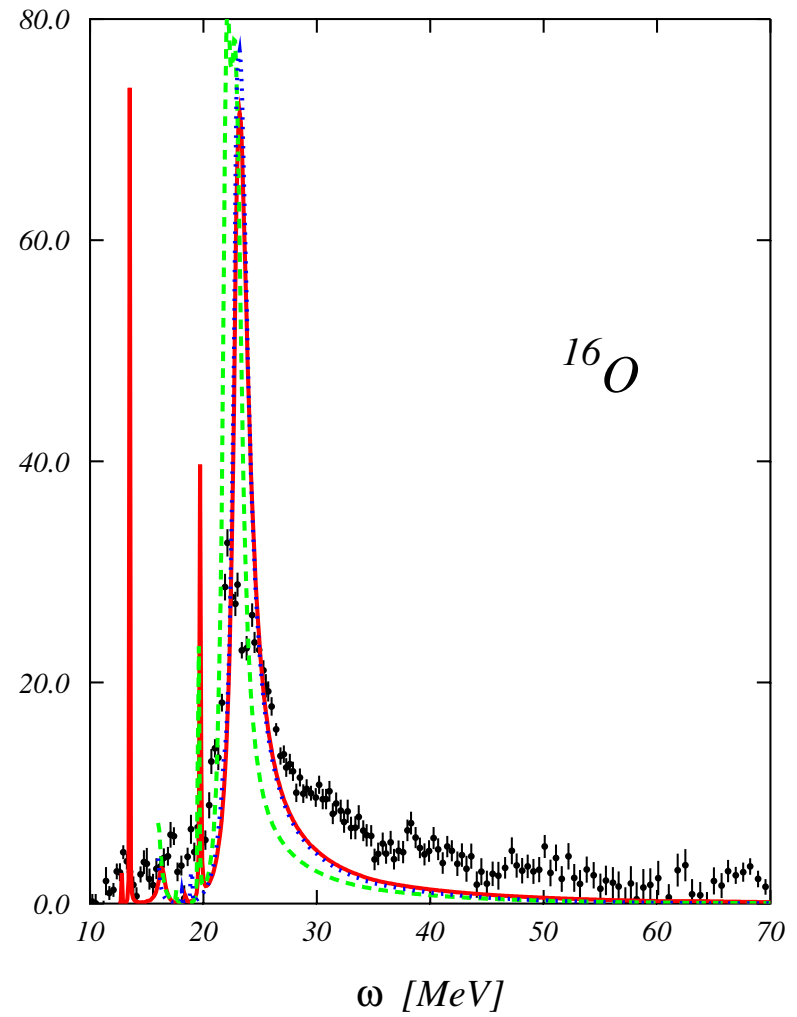
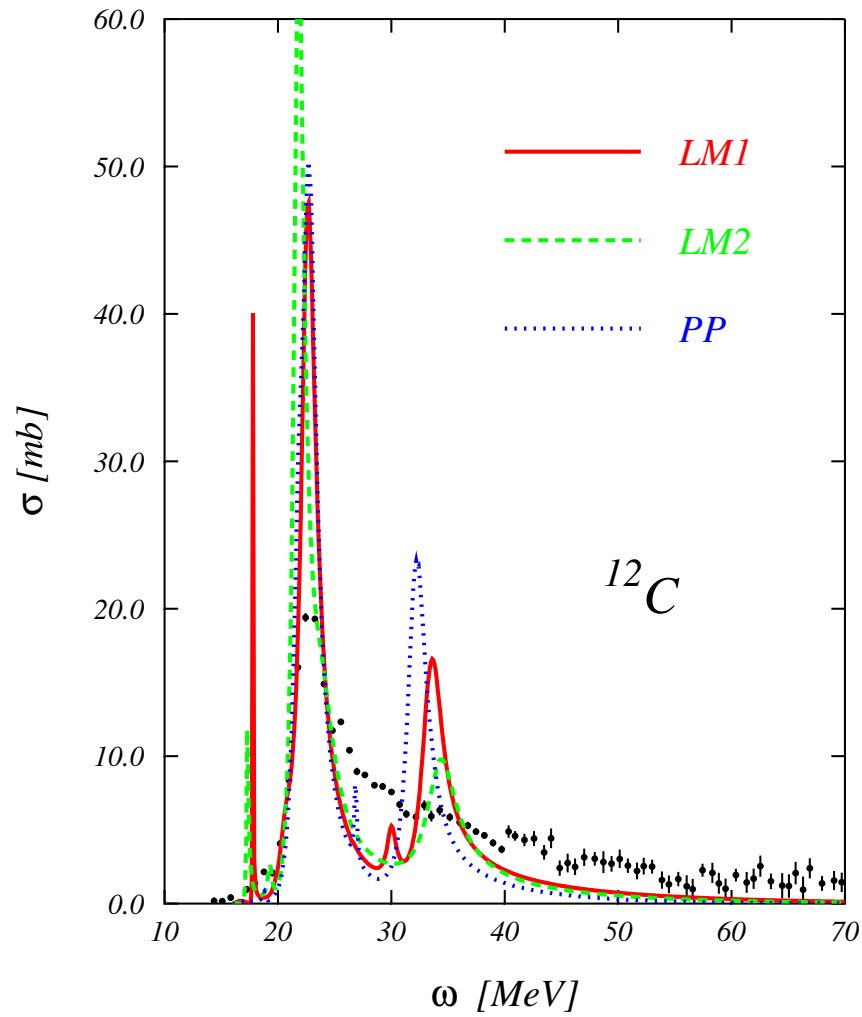
MeV fm ³	F^{ext}	F^{int}	F'^{ext}	F'^{int}	G	G'	$\mathcal{K}$ (¹² C)	$\mathcal{K}$ (¹⁶ O)
LM1	-740.0	60.4	453.0	453.0	166.1	211.4	0.8	0.865
LM2	-630.0	0.0	450.0	195.0	0.0	300.0	1.0	0.865
PP	-1007.0	-54.7	677.8	201.3	60.0	241.0	1.2	1.030
NuInt05	-740.0	60.4	453.0	302.0	-30.2	51.34	1.0	-





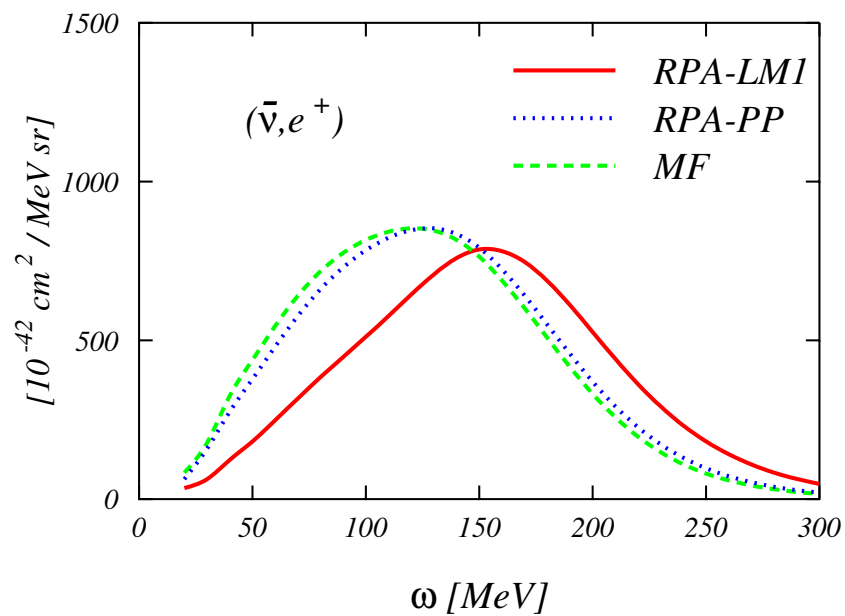
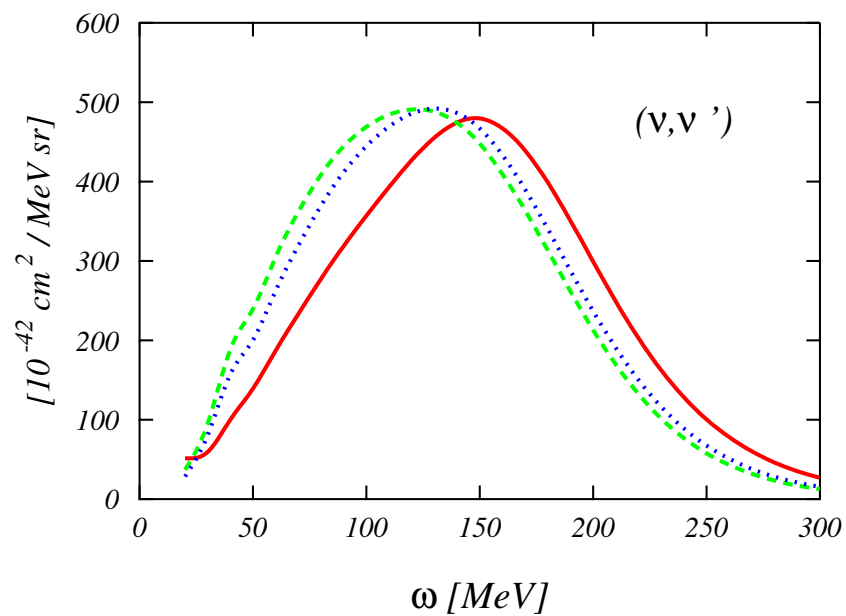
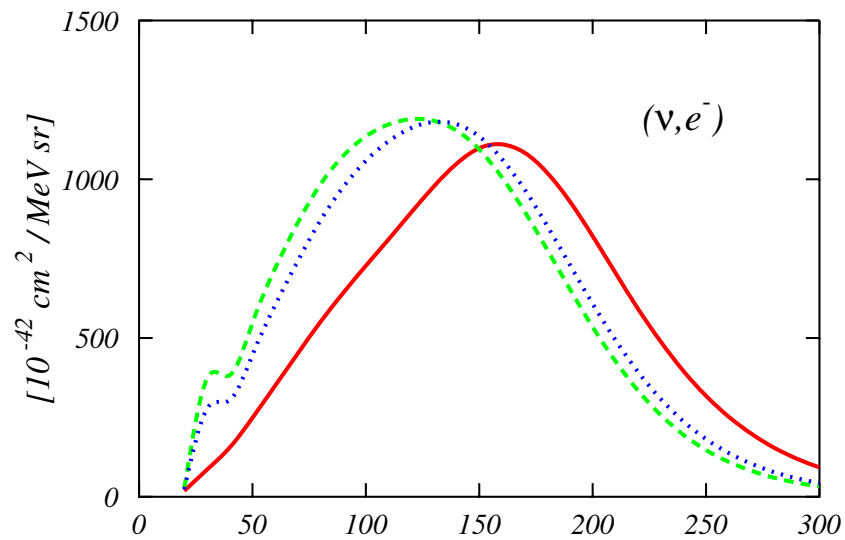
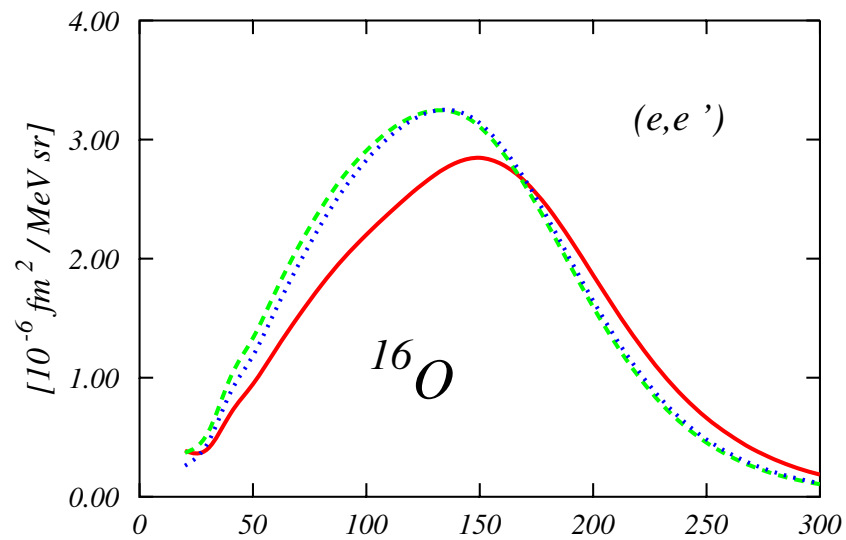
DATA from T.W. Donnelly, J.D. Walecka, I Sick, E.B. Hughes
Phys. Rev. Lett. 21 (1968) 1197.

PHOTOABSORPTION



DATA: J. Ahrens et al., Nucl. Phys. A 251 (1975) 479.

$$\varepsilon_i = 1 \text{ GeV} \quad \theta = 30^\circ$$



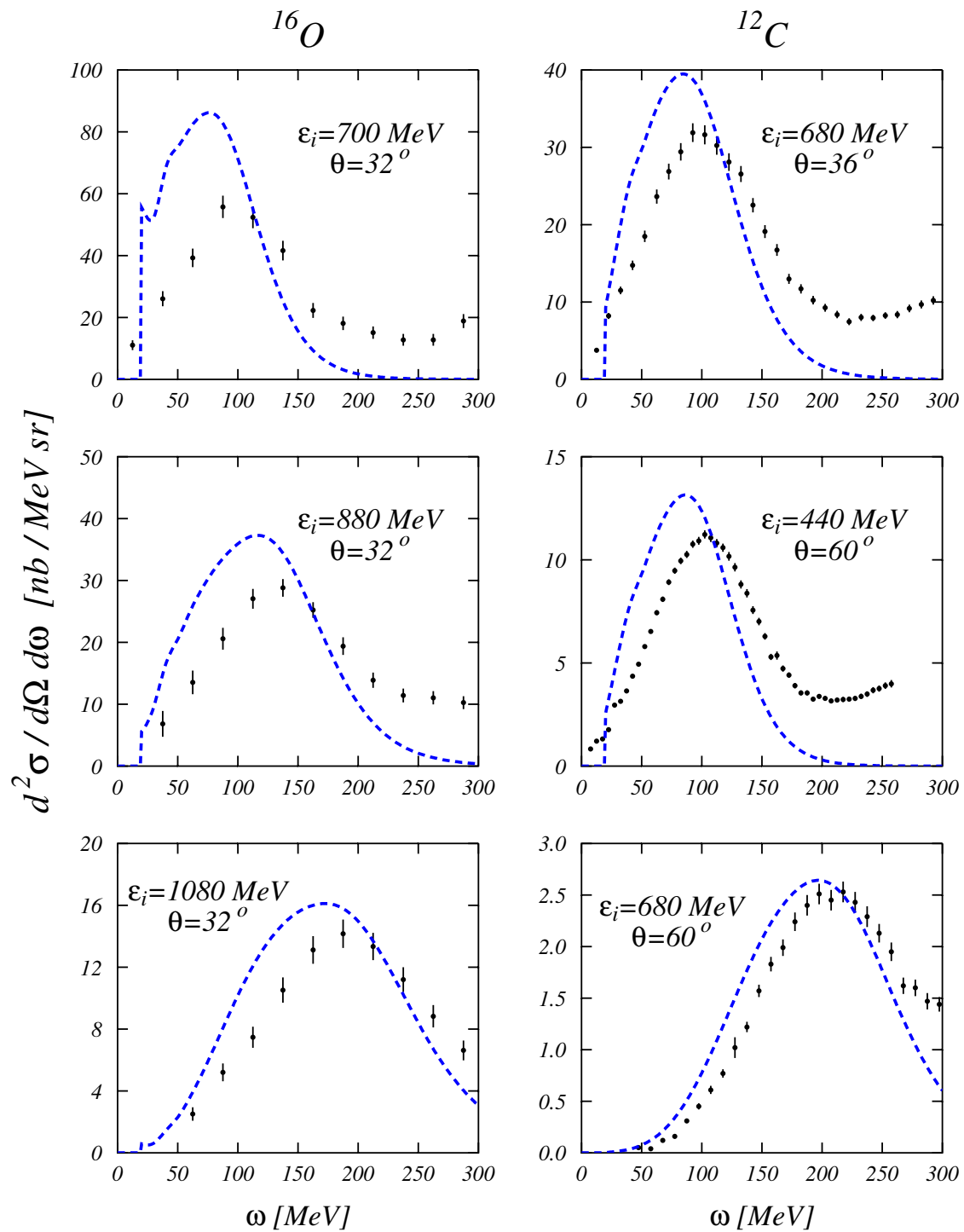
DATA:

M.Anghinolfi *et al.*,

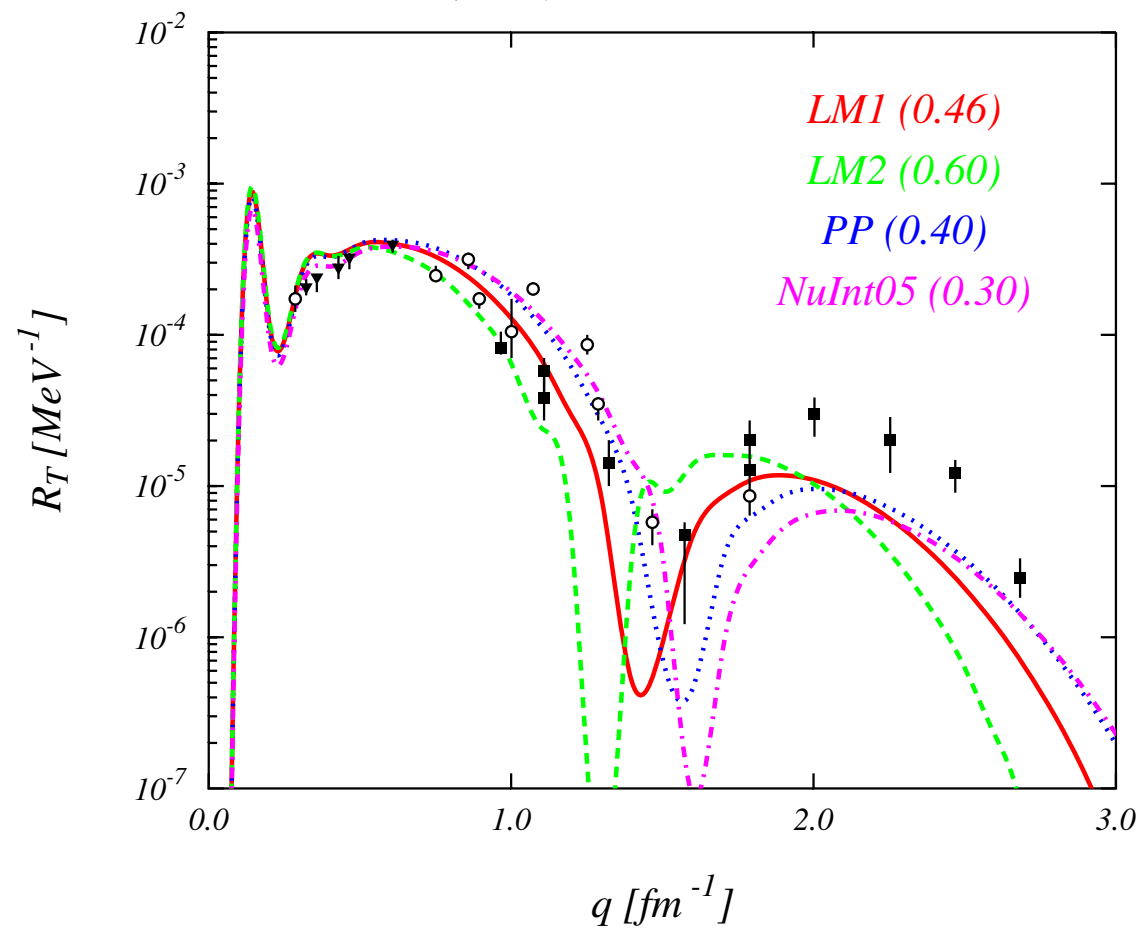
Nucl. Phys. A 602 (1996) 405.

P.Barreau *et al.*,

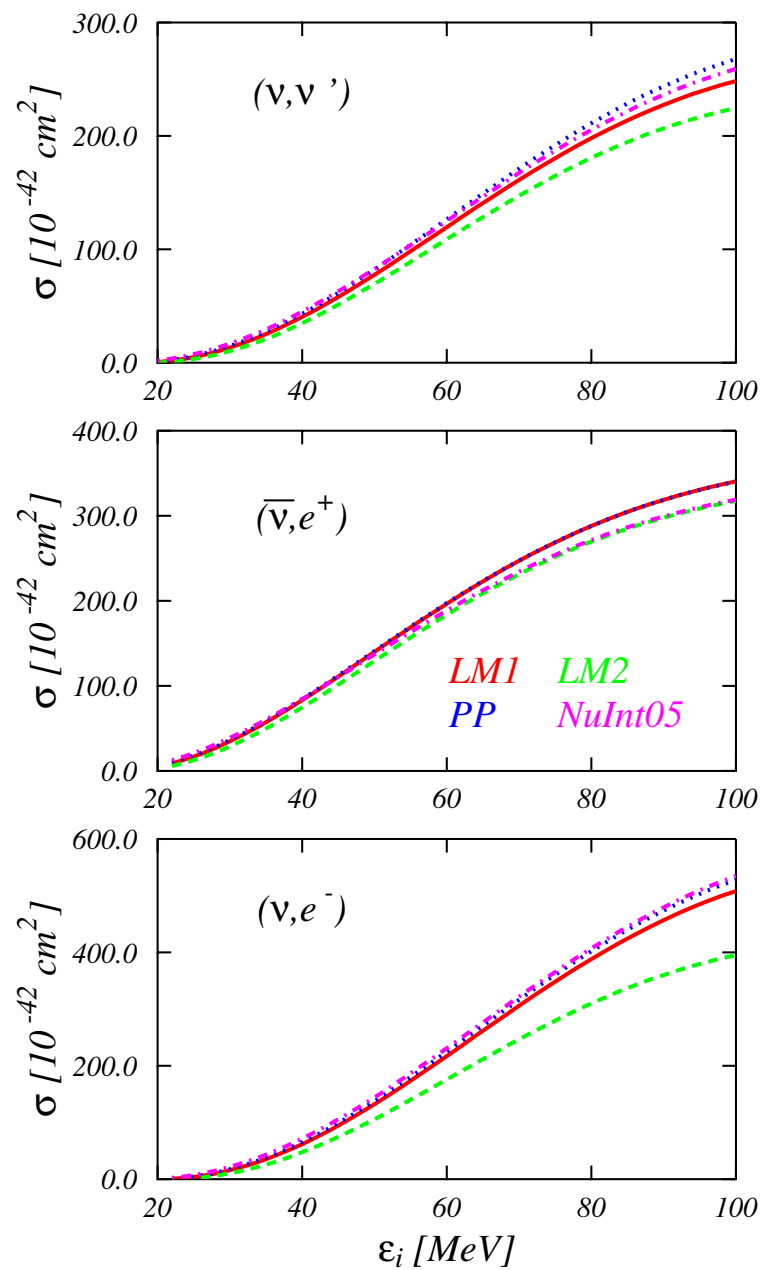
Nucl. Phys. A 402 (1983) 515.



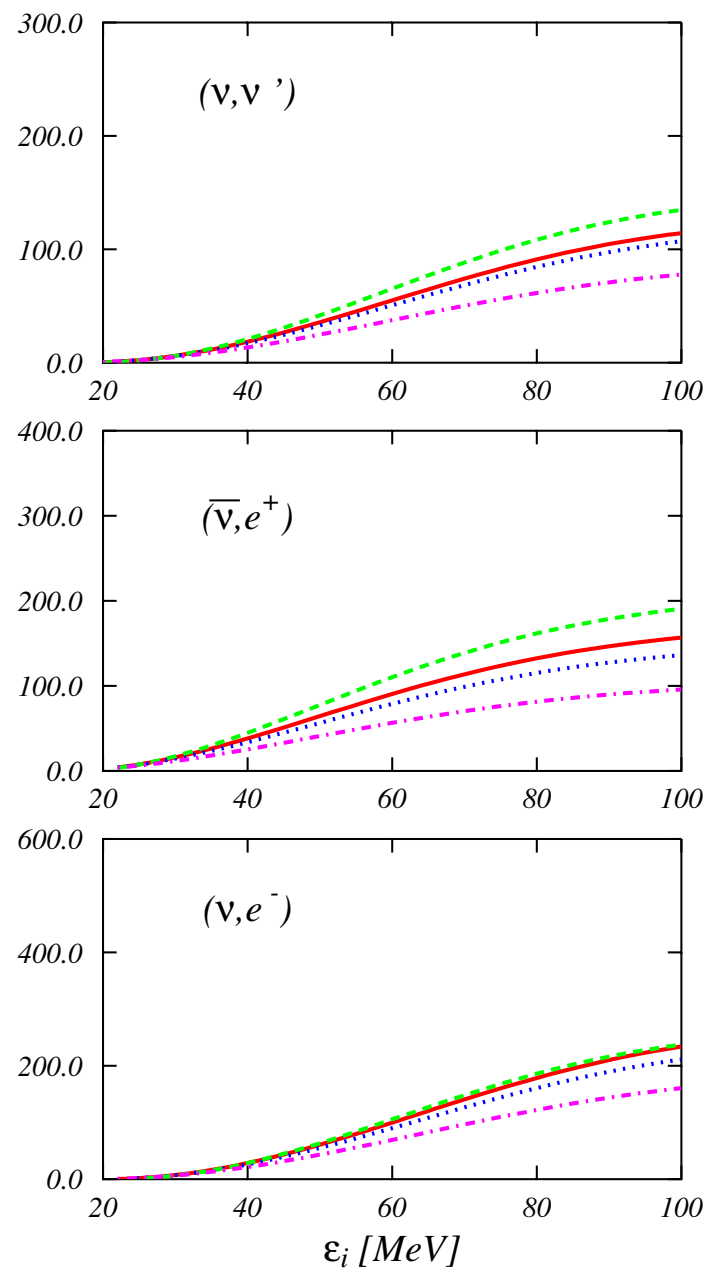
$^{12}\text{C}(e,e')^{12}\text{C} \quad l^+ \quad 15.1 \text{ MeV}$



RPA



quenched RPA



Beyond Continuum RPA (FSI)

$$R^{FSI}(q, \omega) = \int_0^\infty dE R^{RPA}(q, E) [h(E, \omega) + h(E, -\omega)]$$

$$h(E, \omega) = \frac{1}{2\pi} \frac{\Gamma(\omega)}{[E - \omega - \Delta(\omega)]^2 + [\Gamma(\omega)/2]^2}$$

$$\Delta(\omega) = \frac{1}{2\pi} P \int_{-\infty}^{\infty} d\omega' \frac{\Gamma(\omega')}{\omega' - \omega}$$

$$\Gamma(\omega) = \frac{1}{\omega} \int_0^\omega d\epsilon [\gamma(\epsilon + \omega) + \gamma(\epsilon - \omega)]$$

FSI parameters

$$\gamma(\epsilon) = A \left(\frac{\epsilon^2}{\epsilon^2 + B^2} \right) \left(\frac{C^2}{\epsilon^2 + C^2} \right)$$

ϵ [MeV]	A [MeV]	B [MeV]	C [MeV]
≤ 10.0	6.0	60.0	110.0
≥ 40.0	11.0	20.0	110.0

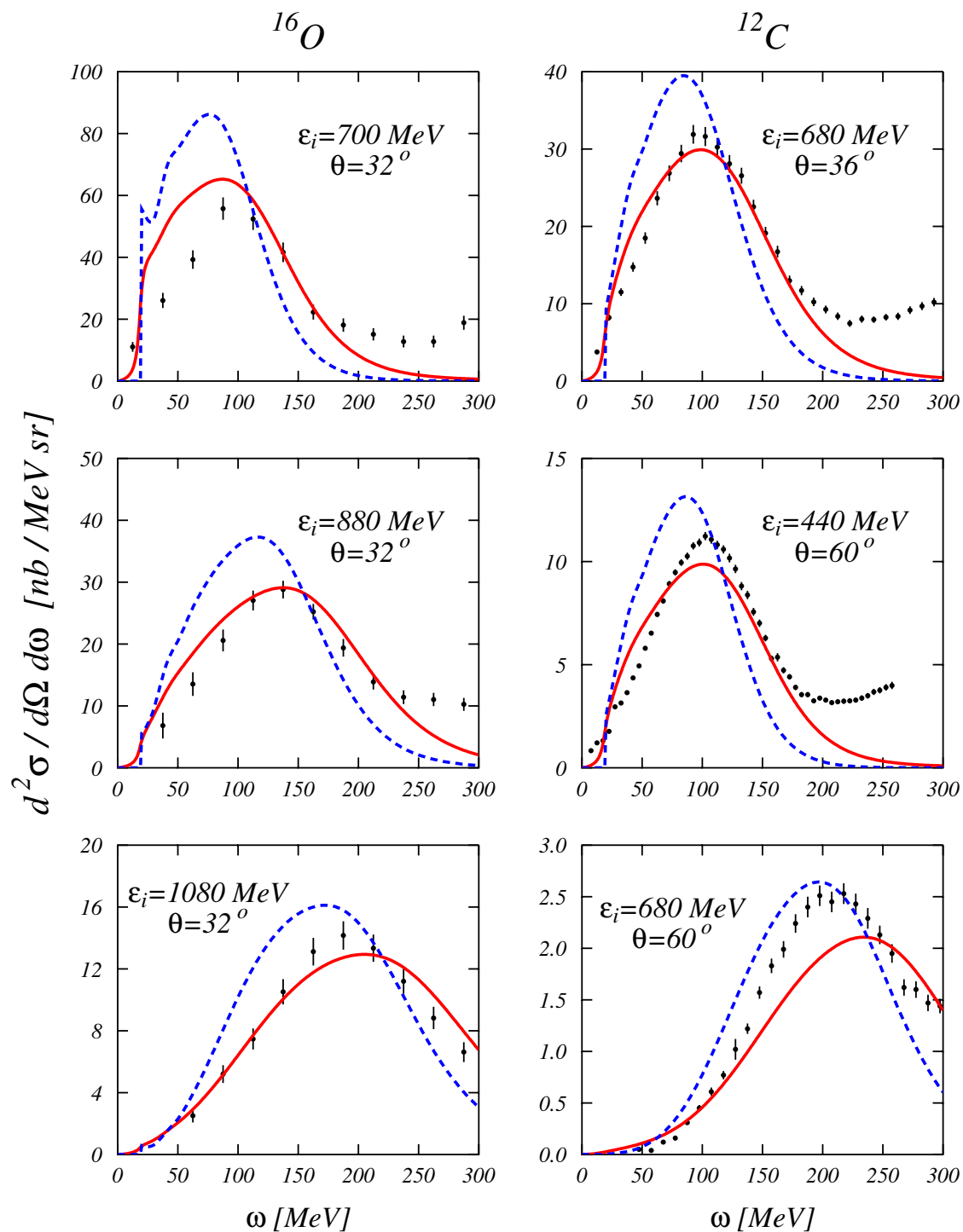
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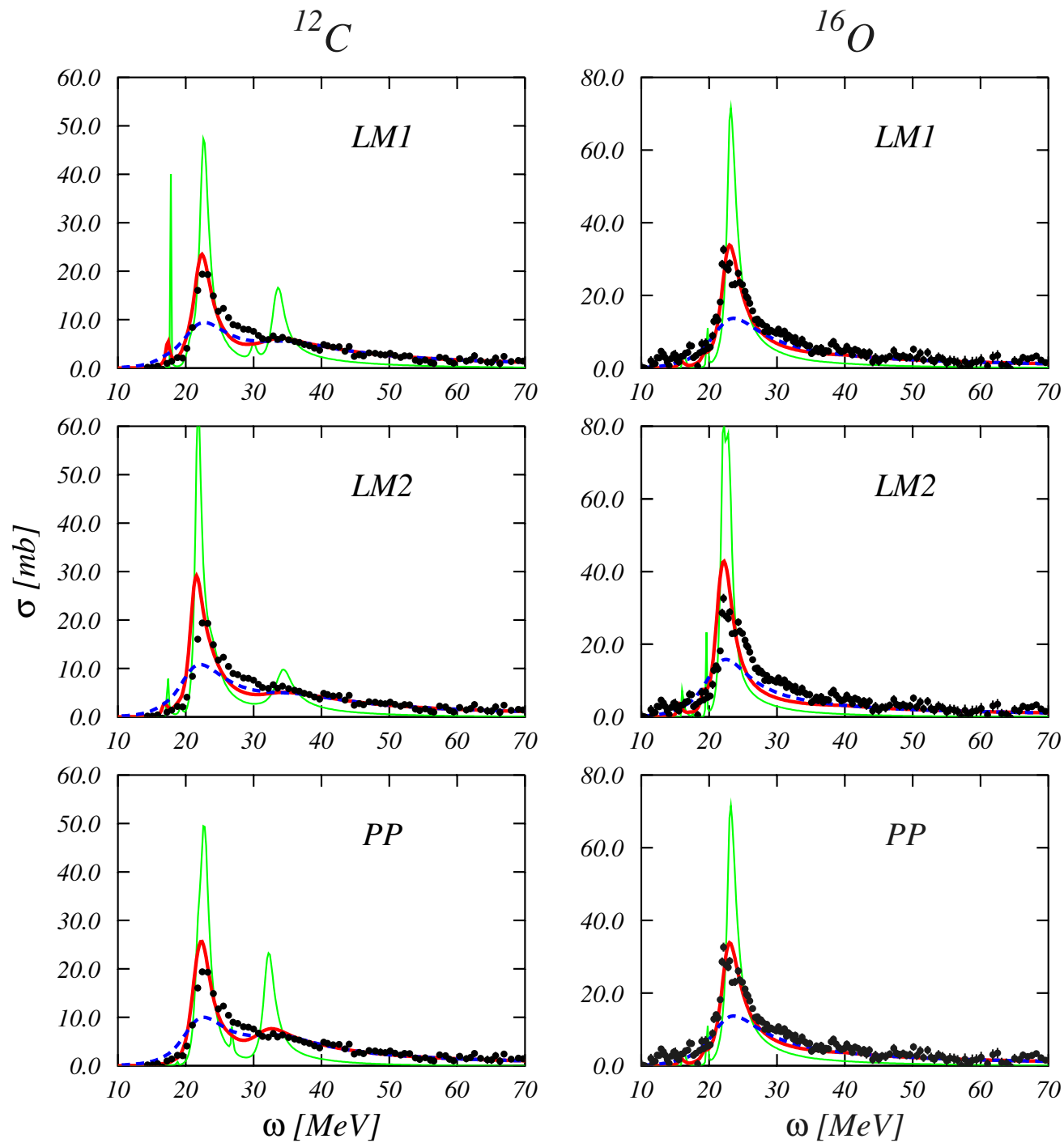
M.Anghinolfi *et al.*,

Nucl. Phys. A 602 (1996) 405.

P.Barreau *et al.*,

Nucl. Phys. A 402 (1983) 515.





$$^{12}\text{C}(e,e')^{12}\text{C}$$

DATA:

(a,b,c) quoted in

E. Kolbe

Nucl. Phys. A

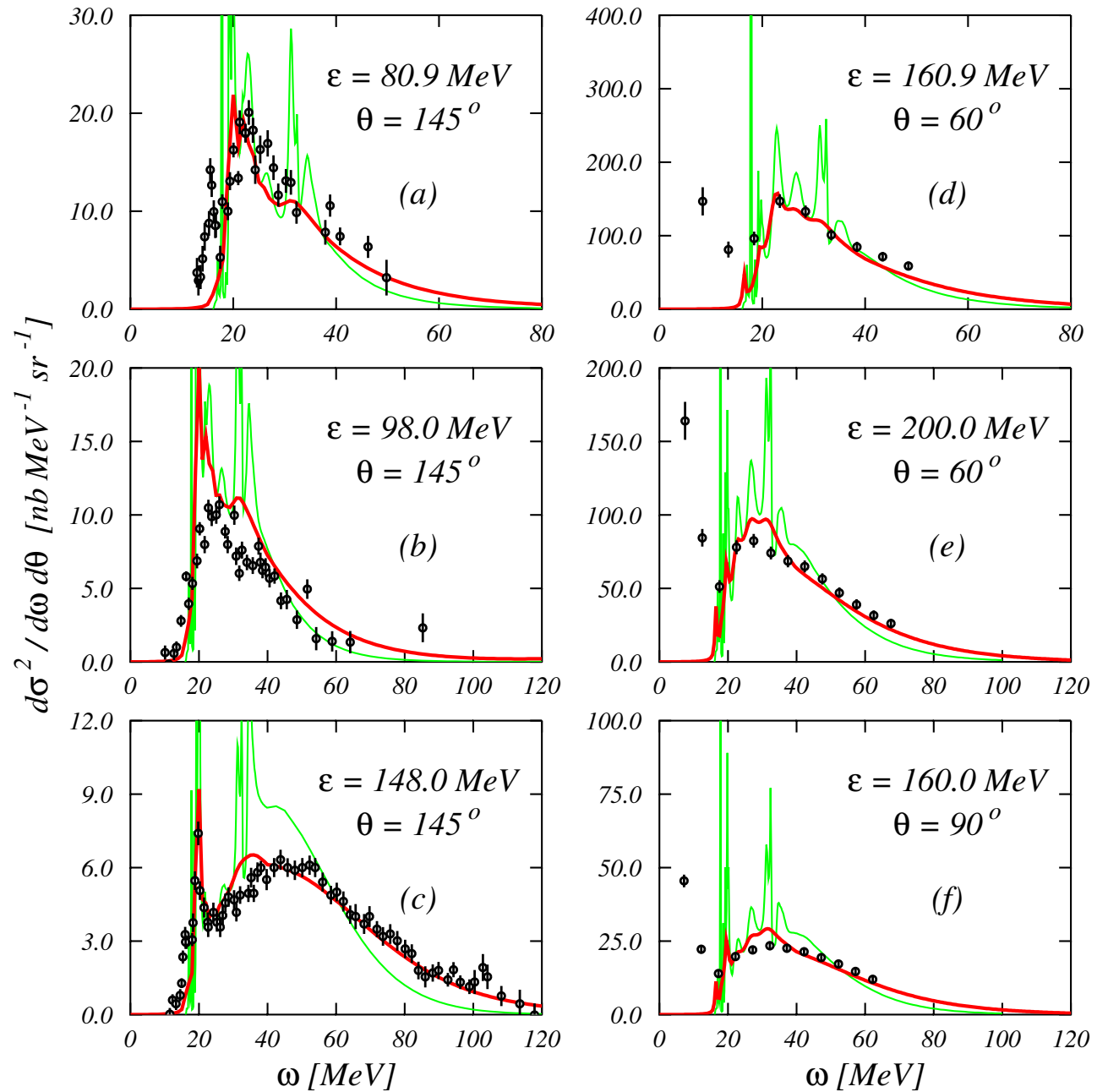
602 (1996) 405.

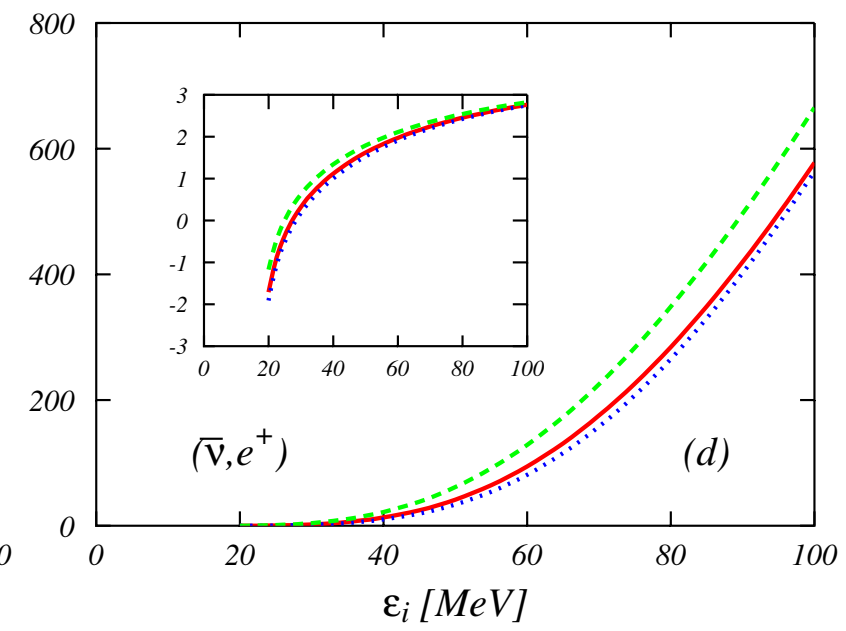
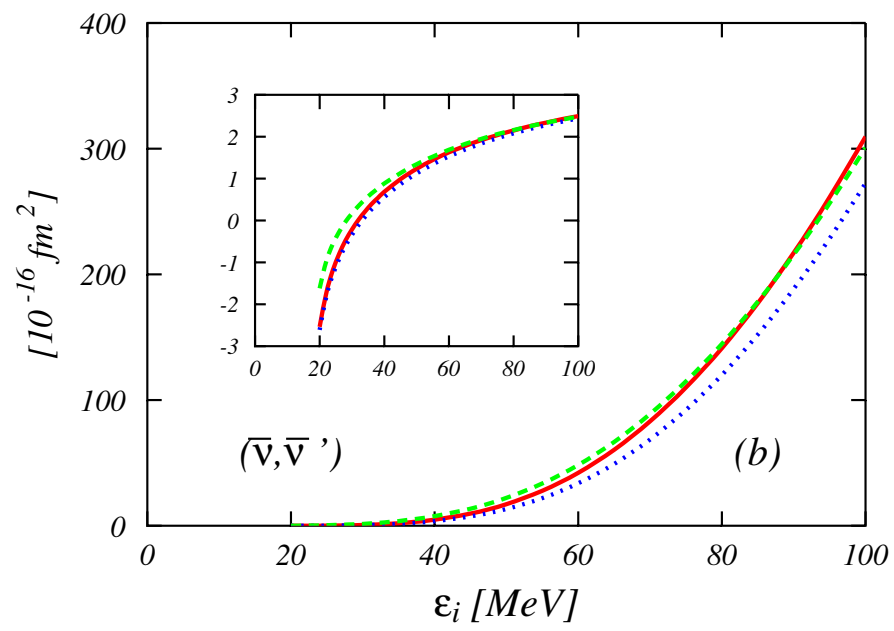
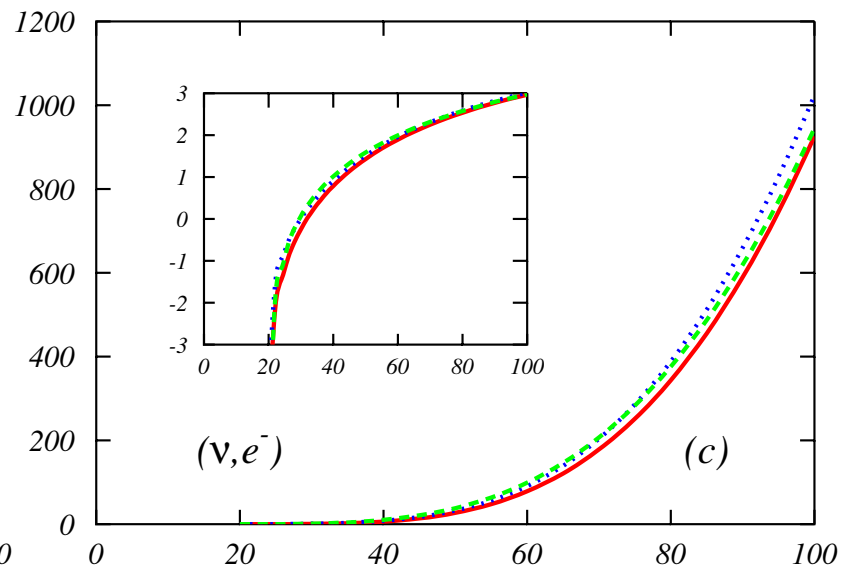
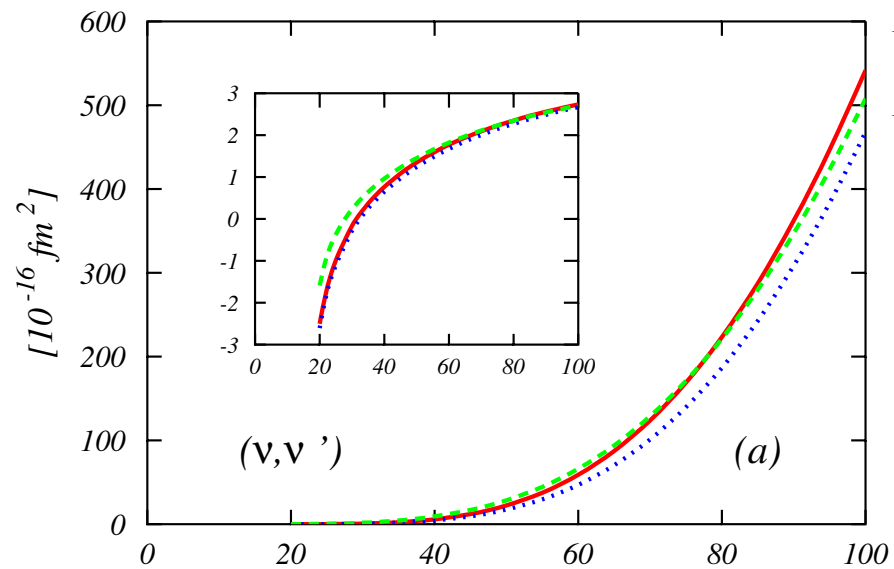
(d,e,f)

P.Barreau *et al.*,

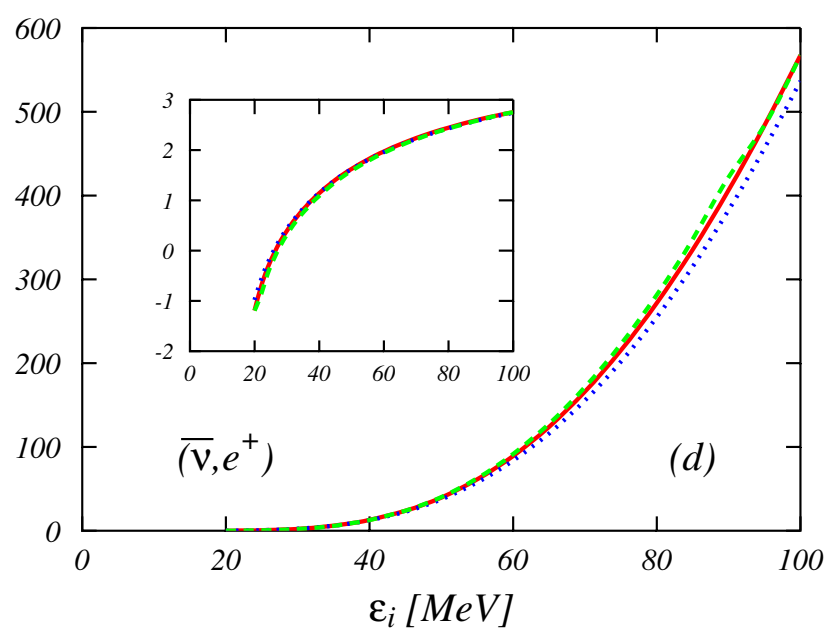
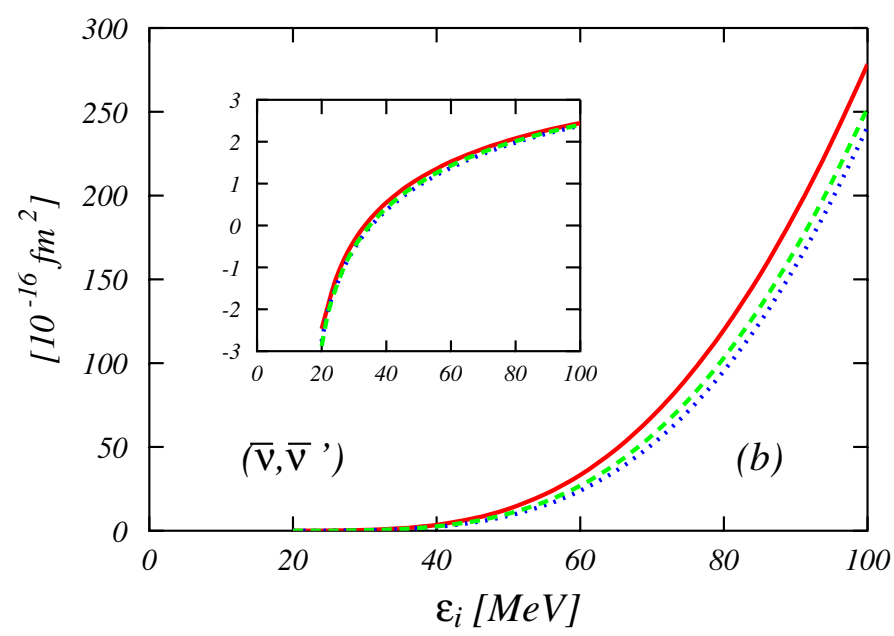
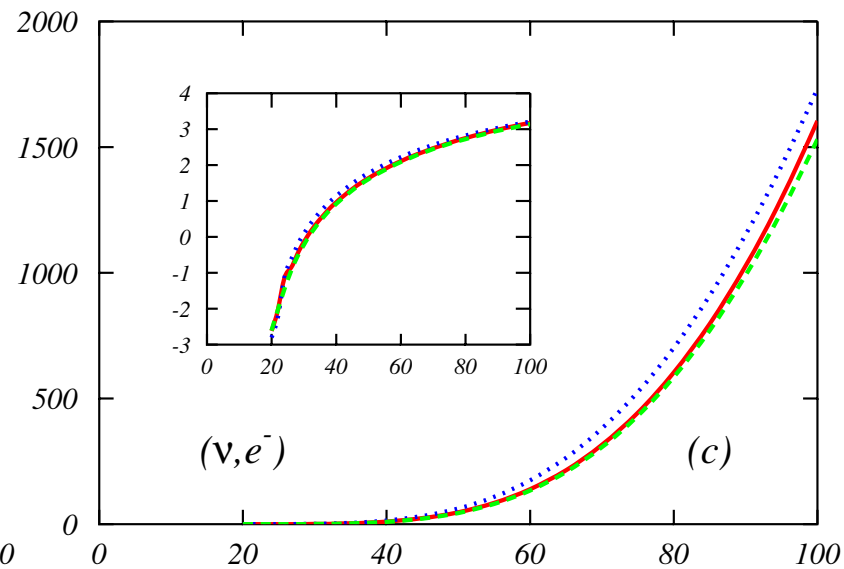
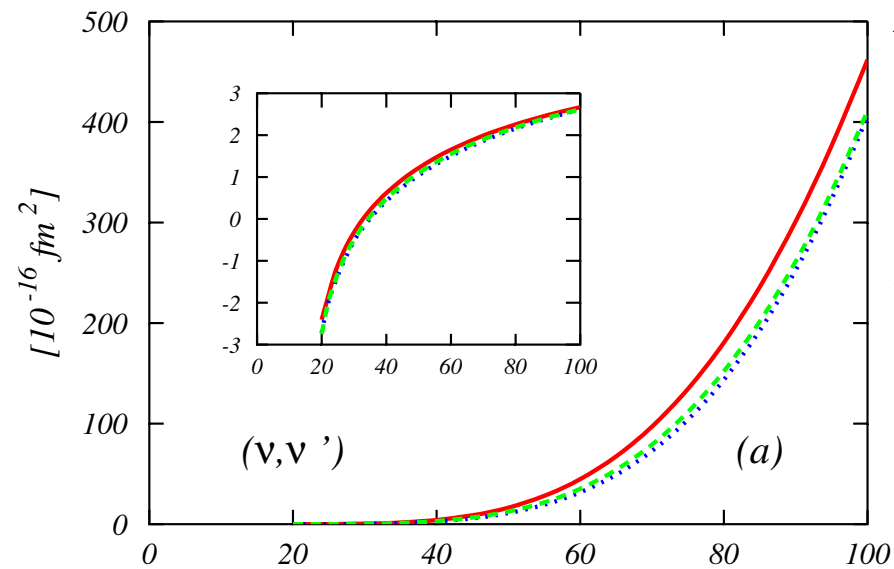
Nucl. Phys. A

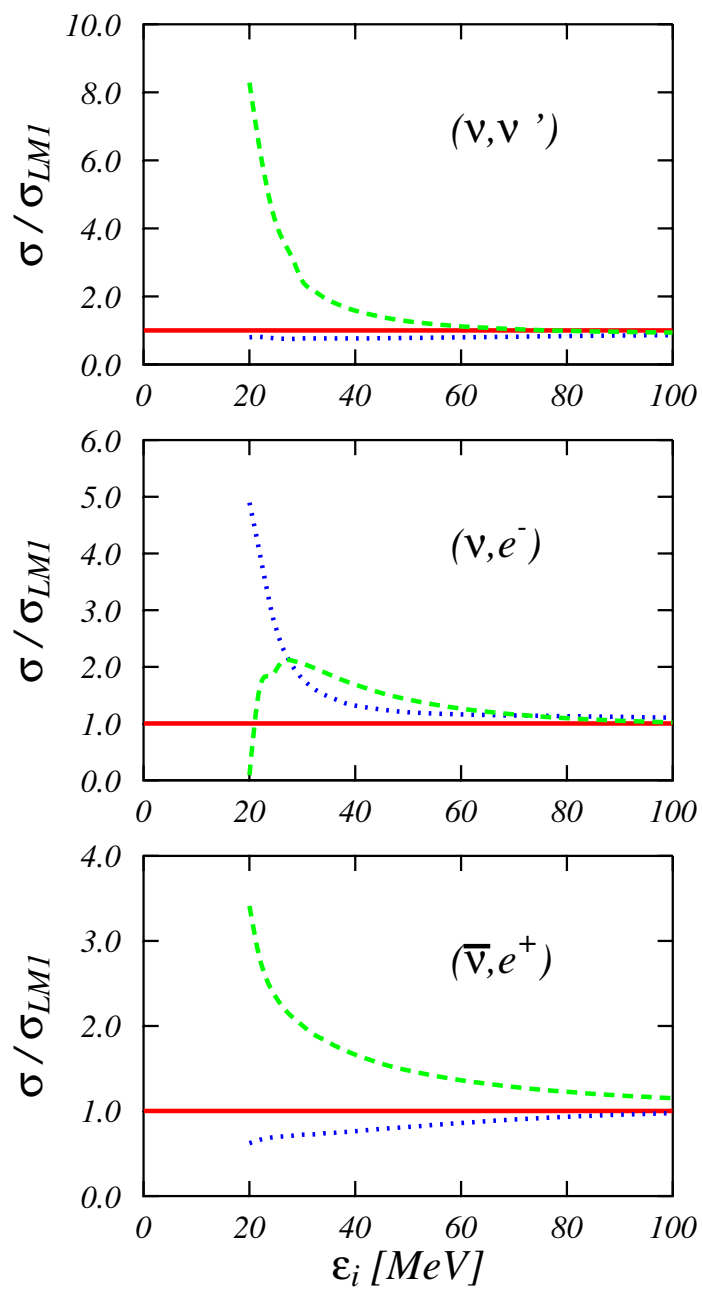
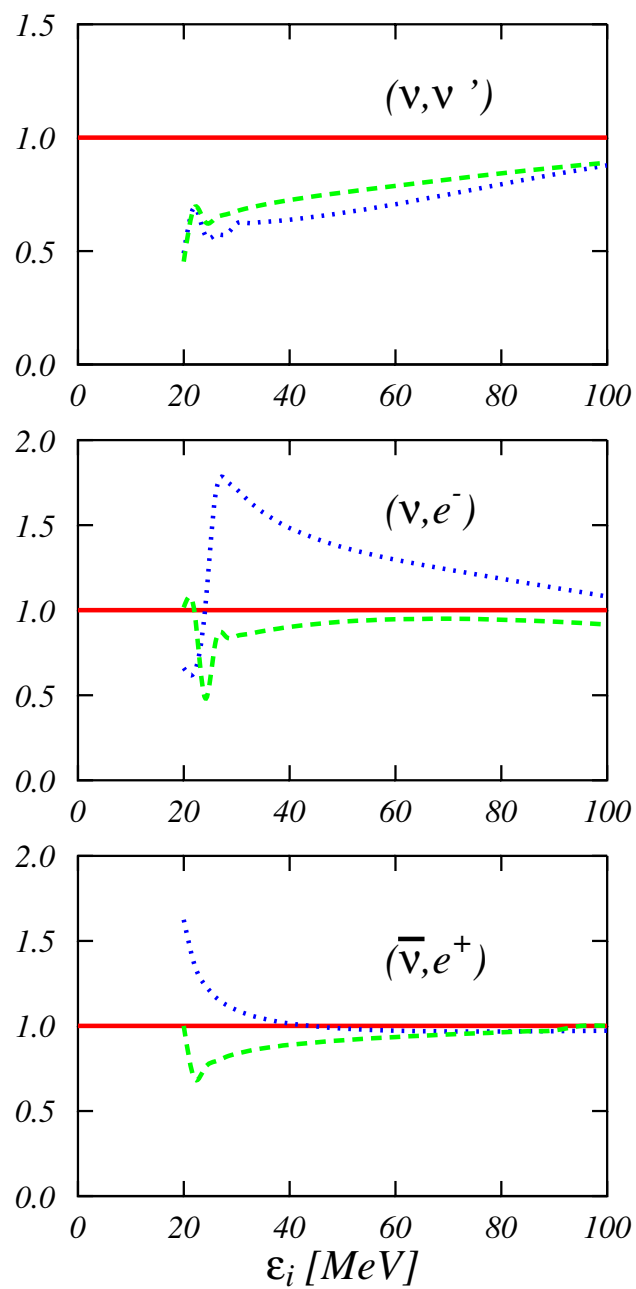
402 (1983) 515.



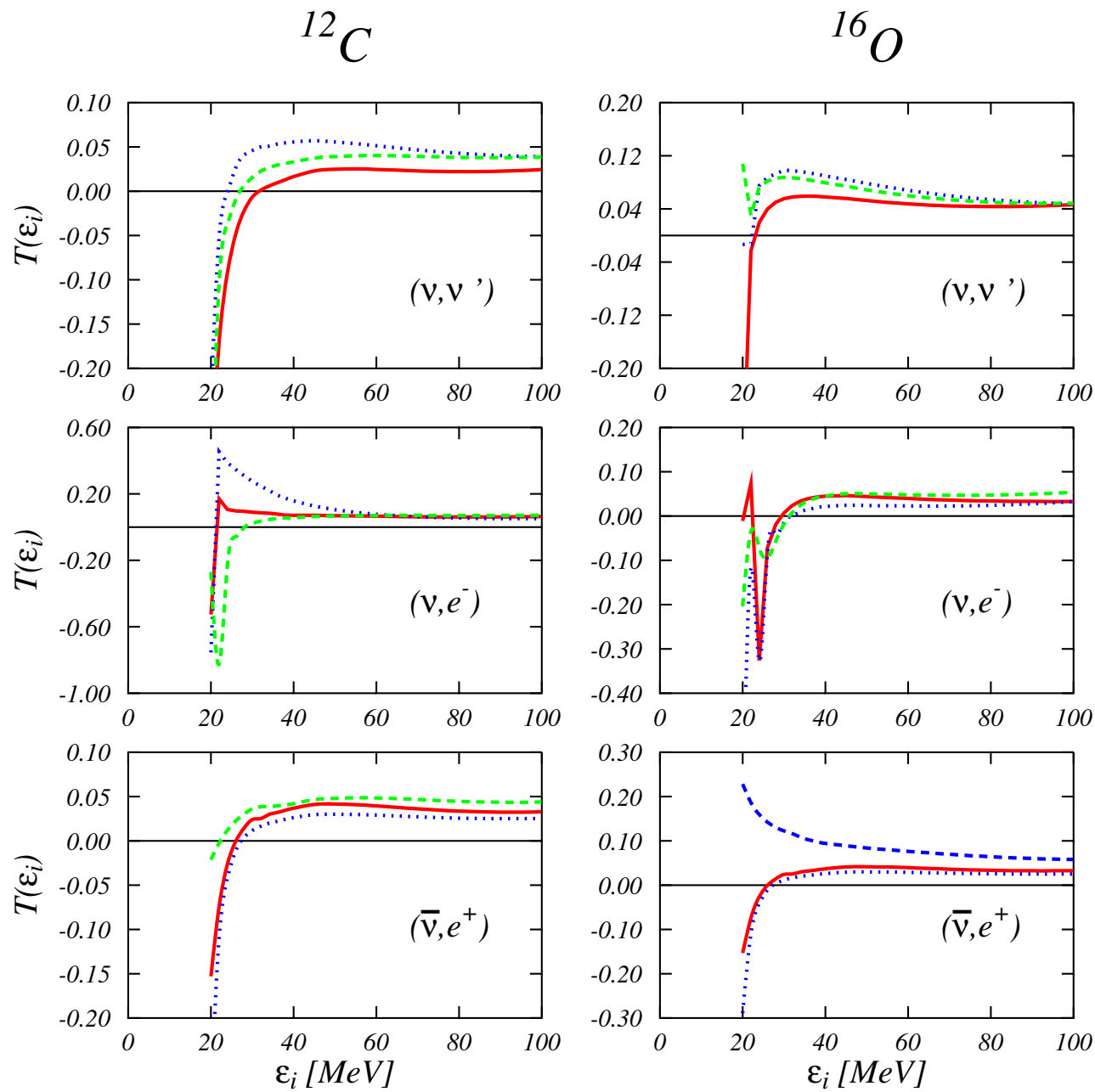
^{12}C 

^{16}O



^{12}C  ^{16}O 

$$T(\epsilon_i) = \frac{\sigma^{RPA} - \sigma^{FSI}}{\sigma^{RPA} + \sigma^{FSI}}$$



Supernova neutrino fluence

$$f(\epsilon_i) = \mathcal{C} \frac{\epsilon_i^2}{1 + \exp(\epsilon_i/T - \alpha)}$$

	T [MeV]	α	$\langle \epsilon_i \rangle$ [MeV]
I	3.49	0.0	11.0
II	3.80	0.0	12.0
III	2.46	4.0	11.0

(v, v')

