

Evidenze sperimentali di transizione di fase liquido-gas nei nuclei

M. Bruno, F. Cannata, **M. D'Agostino**,
E. Geraci, P. Marini, J. De Sanctis, G. Vannini



Universita' Bologna



INFN-Bologna

NUCL-EX Collaboration:

INFN e Universita' Bologna, Firenze, Milano, Napoli,
Trieste

INFN - Laboratori Nazionali di Legnaro

LPC e GANIL - Caen (Francia)

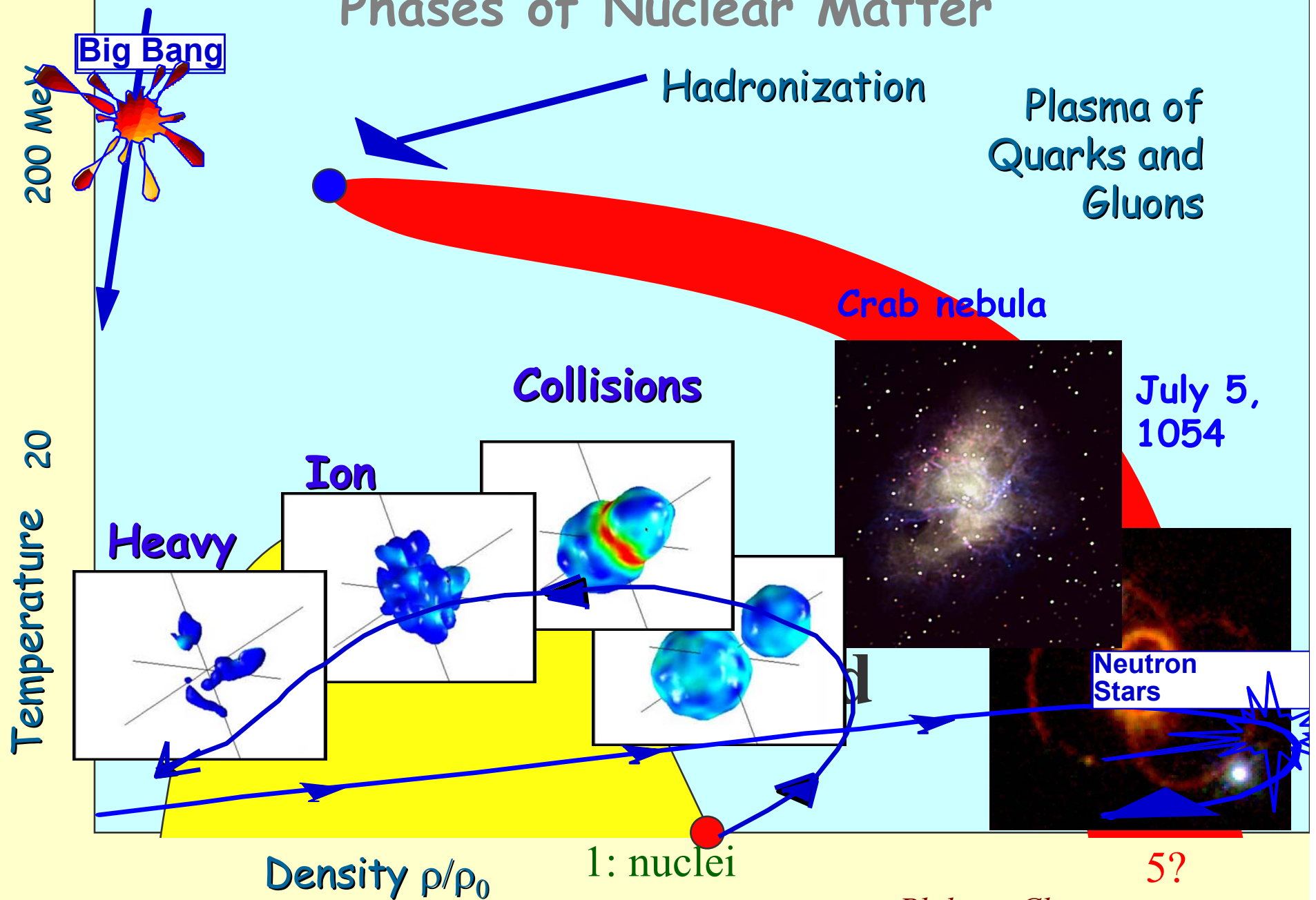
IPN - Orsay (Francia)

Schema

- Che osservabili prevede la teoria nelle transizioni di fase
- Cosa si misura e come
- Come si analizzano i dati: osservabili, sorting
- Confronto fra dati e teoria
- Cosa ancora si deve misurare

Non necessariamente in questo ordine !!!

Phases of Nuclear Matter



Philippe Chomaz artistic view

Transizioni di fase:

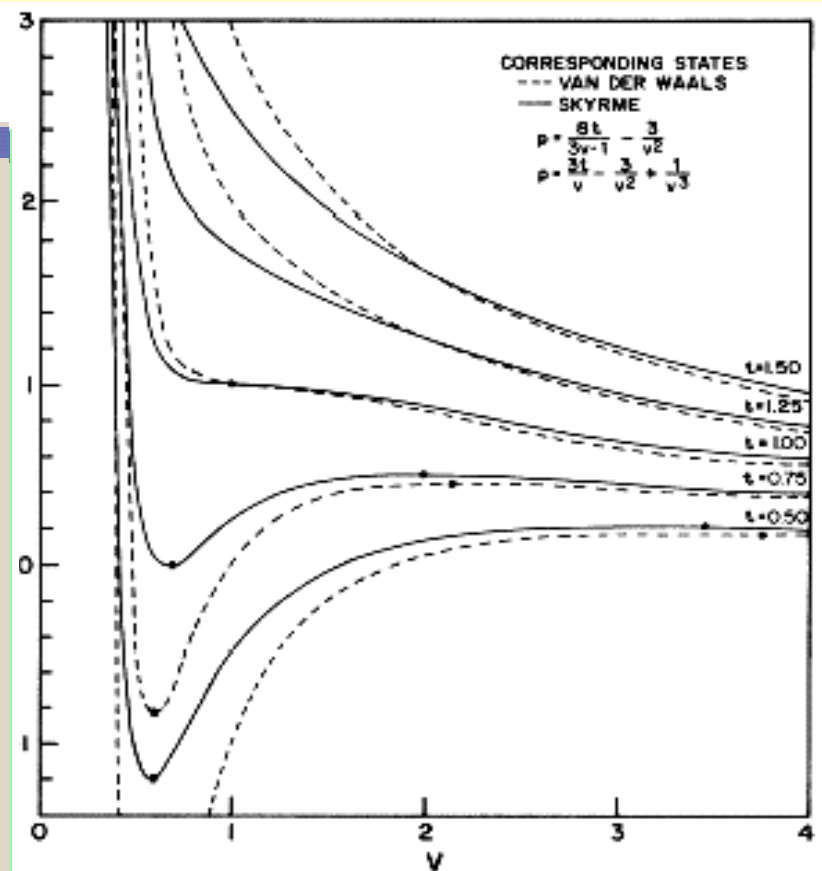
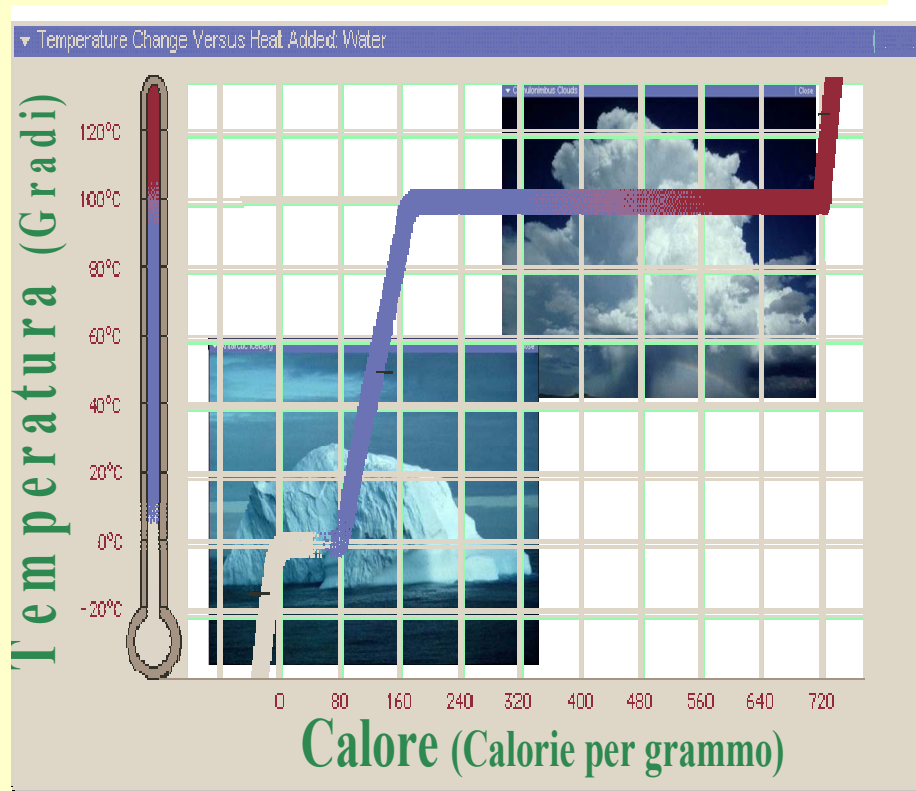
Keywords

	QG Plasma	Liquid-Gas
Soppressione di canali	J/Ψ	Risonanza gigante di dipolo
Fenomeno critico	deconfinamento	multiframmentazione
Tempi di equilibrio e di rilassamento	$t_{eq} \approx 1 \text{ fm}/c$	$t_{eq} \approx 100 \text{ fm}/c$
Parametri critici	Temperatura critica ($T_c \approx 170 \text{ MeV}$) Esponenti critici	Temperatura critica ($T_c \approx 5 \text{ MeV}$) Esponenti critici
Fluttuazioni	temperatura e molteplicità	energia (capacità termica negativa)
Ordine della transizione	Primo o secondo?	Primo o secondo?

Forze nucleari:

Simili a forze di Van der Waals
repulsive a piccole distanze
attrattive a grandi distanze

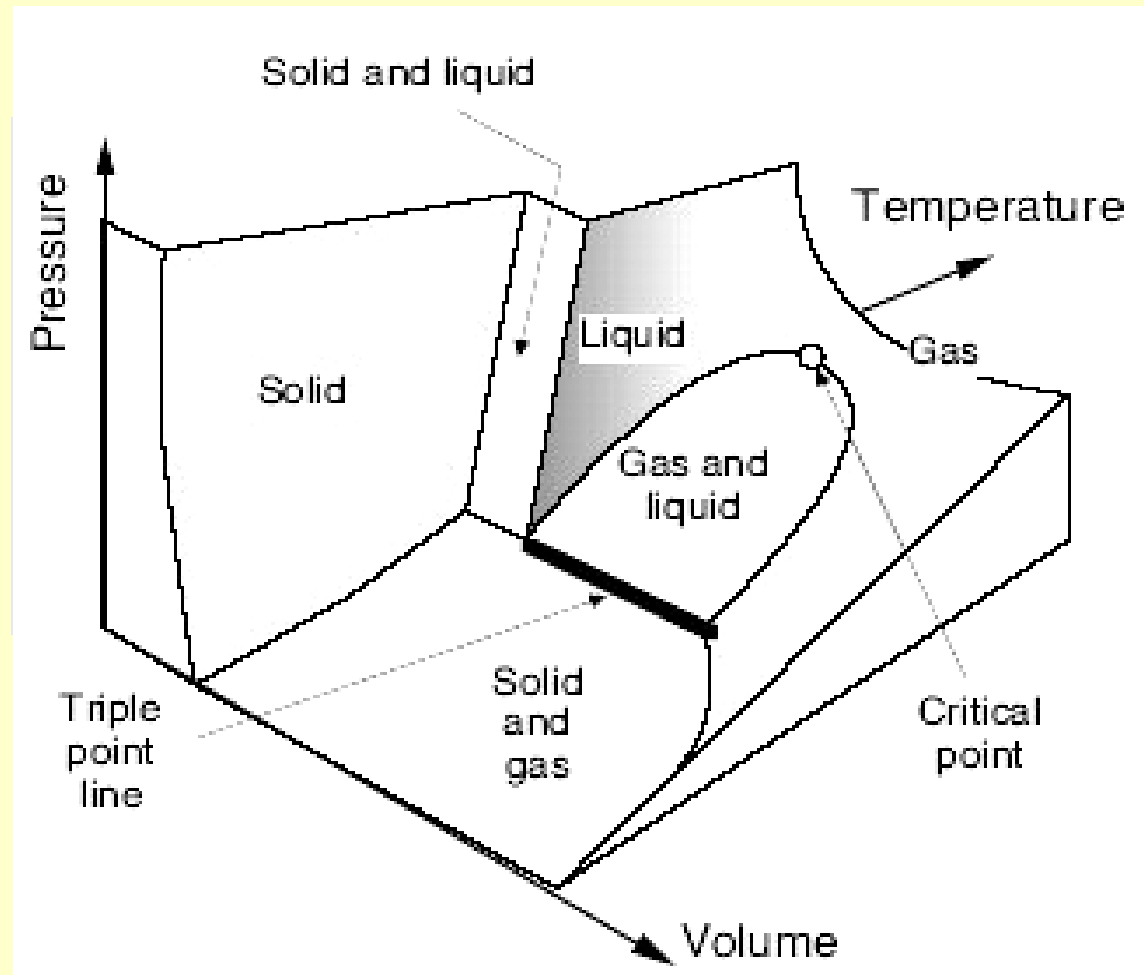
H.Jaqaman et al. PRC27(1983)2782



Equazione di stato della materia nucleare

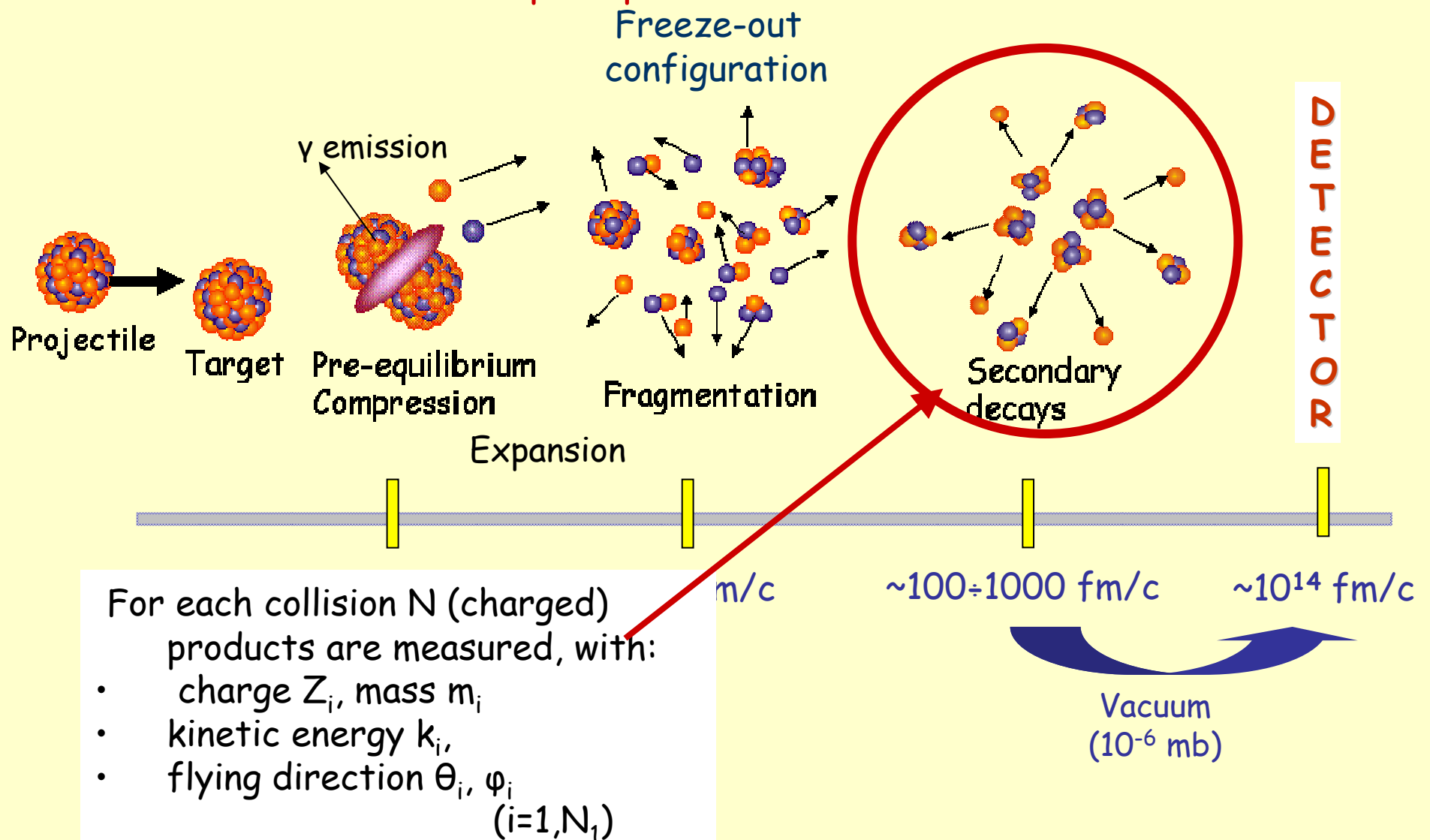
Sono possibili
transizioni di fase?

Il nucleo a basse energie di eccitazione si comporta come un liquido (formula di massa di Weizsäcker) ad alta energie di eccitazione come un gas (modello a gas di Fermi)



Heavy Ion collisions at intermediate energies

Inclusive observables correspond to averages, weighted on the impact parameter

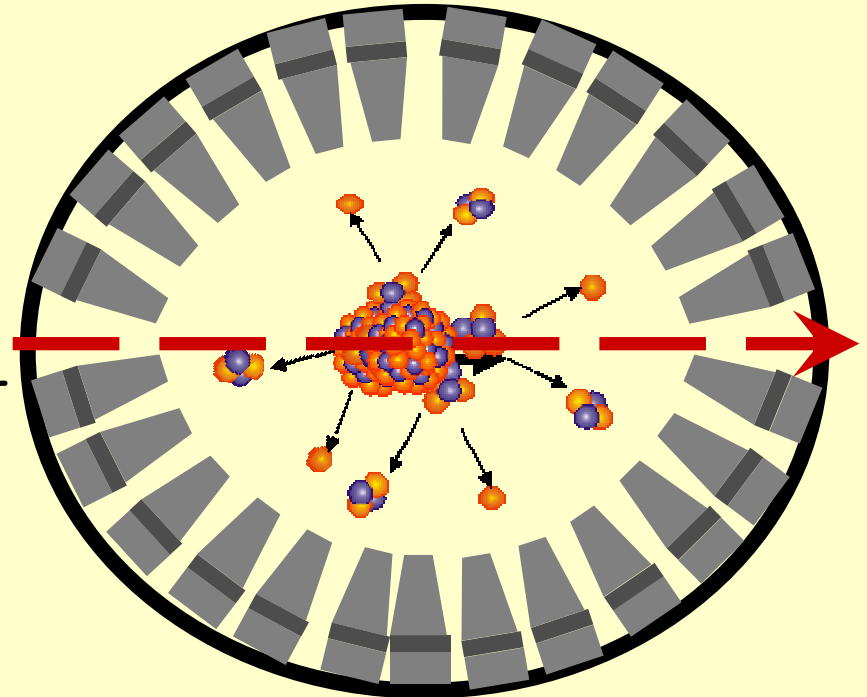


H.I. Collisions , intermediate energies (10-100 AMeV): 1-st generation 4 π devices

- $Z_i, k_i, \theta_i, \varphi_i$ are measured for almost all charged products, event by event, with high energy resolution (few %) and low energy thresholds (gas detectors)
- Fragments and particles are detected at $\sim 10^{14}$ fm/c, as they were at 10^3 fm/c, since the propagation in vacuum does not allow further interactions with matter.

- Statistical multidimensional analyses performed on global (event) observables allow to sort the events in classes of centrality.
- The decaying system can be identified and its calorimetric excitation energy can be estimated from the energy balance:

$$E^* + m_0 = \sum_{i=1}^M (m_i + k_i) + M_n (m_n + k_n)$$



- m_i are measured only for light products
- neutrons and γ are quite often not measured

How many detection cells are needed?

N = expected multiplicity,

X = number of detectors

ε = geometrical coverage/ 4π = probability to detect
1 particle $\varepsilon(1)$

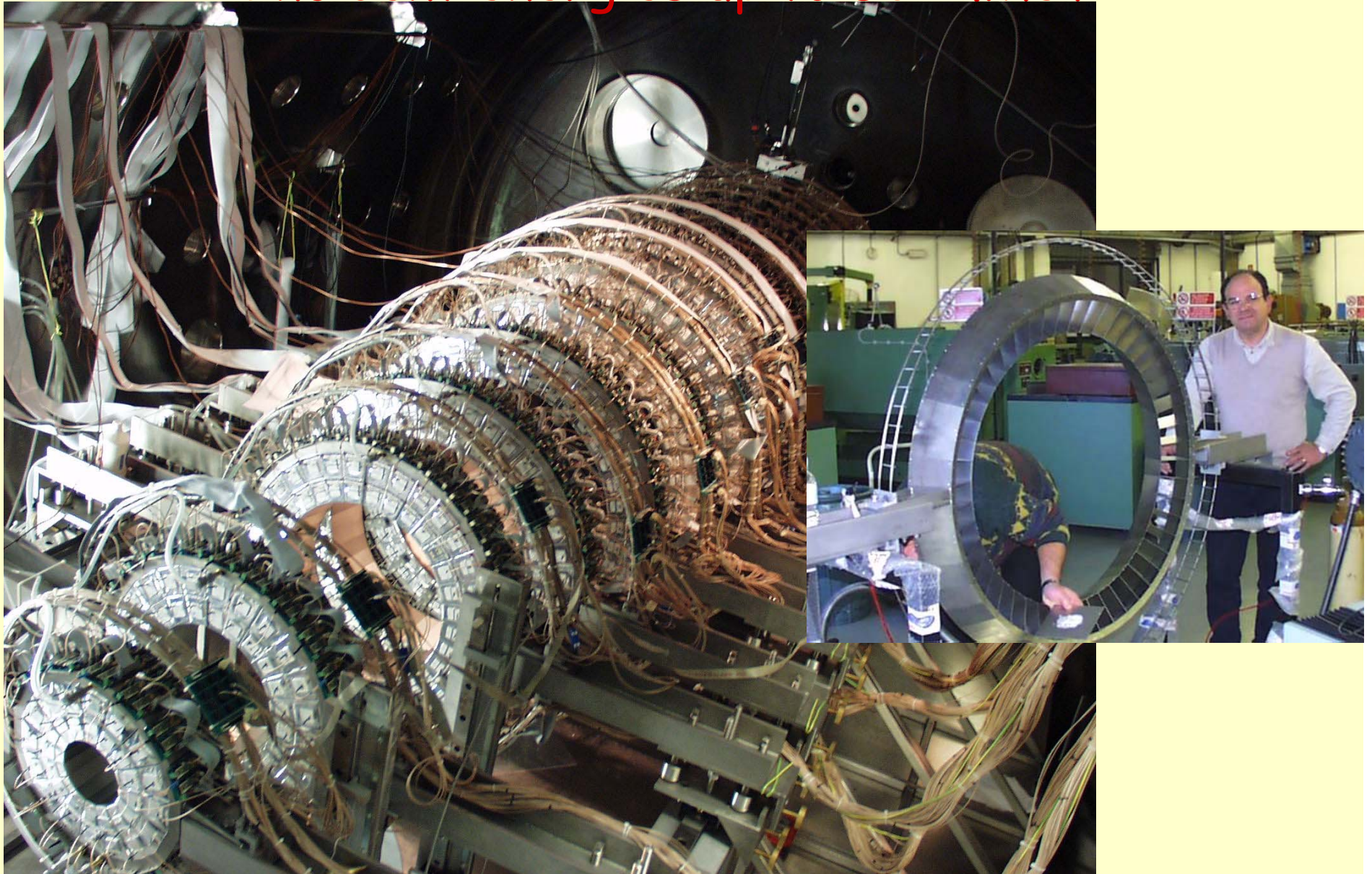
- $\varepsilon(N) = \varepsilon(1)^N$

- $P(\text{double}) = (N-1)/(2X)$

A reasonable compromise is $P(\text{double}) \sim \text{few}\%$

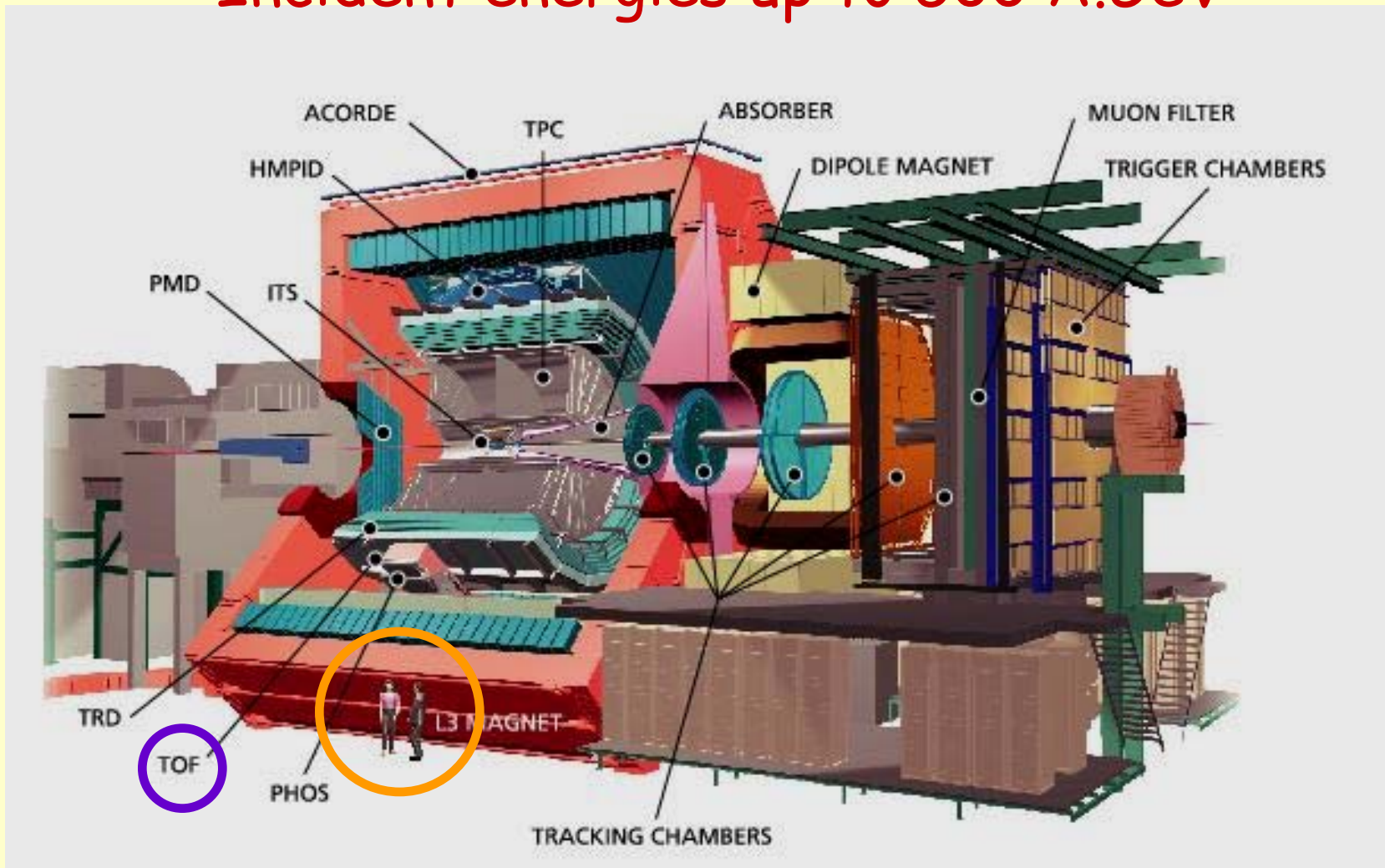
$P(\text{double}) = 1\% \rightarrow X=250$ for $N=50$,
 $\rightarrow X=2500$ for $N=500$
..... etc.

Chimera detector@LNS (CT) ($\sim 10^3$ detection cells)
Incident energies up to 50 A.MeV



Alice detector@CERN (CH)

Incident energies up to 500 A.GeV



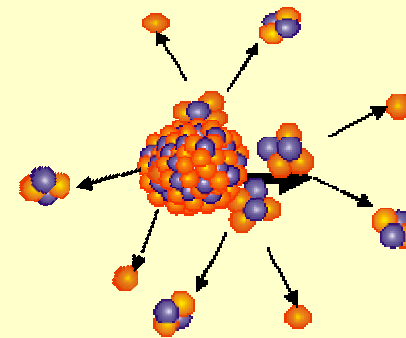
157 000 cells, 150 m² for TOF

Structure of the measured quantities

event Primary quantities:

1. $N_1 (Z_i, m_i, k_i, \theta_i, \varphi_i, i=1, N_1)$
2. $N_2 (Z_i, m_i, k_i, \theta_i, \varphi_i, i=1, N_2)$
3. $N_3 (Z_i, m_i, k_i, \theta_i, \varphi_i, i=1, N_3)$
4. $N_4 (Z_i, m_i, k_i, \theta_i, \varphi_i, i=1, N_4)$
5. $N_5 (Z_i, m_i, k_i, \theta_i, \varphi_i, i=1, N_5)$
- ...
- X $N_X (Z_i, m_i, k_i, \theta_i, \varphi_i, i=1, N_X)$

X = about 10^8
(some GigaBytes → some
TeraBytes after analysis)



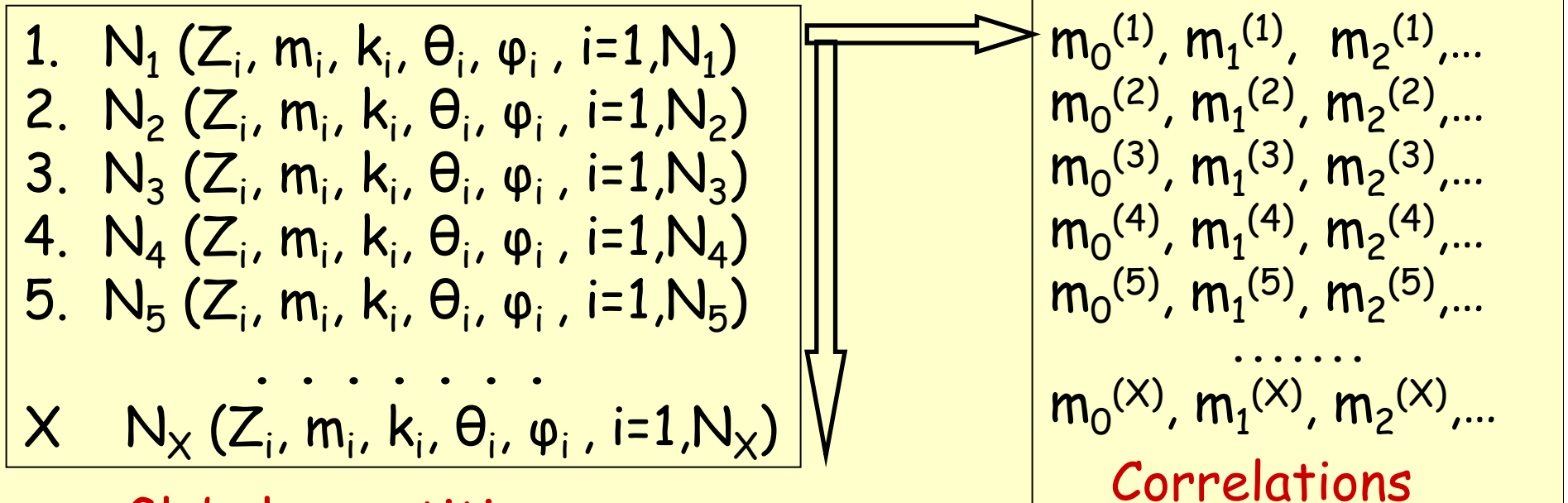
Each event = N (charged)
products,

- charge Z_i , mass m_i
- kinetic energy k_i ,
- flying direction θ_i, φ_i
($i=1, N_1$)

Structure of the calculated quantities

e.g.

Moments $m_k = \sum Z_i^k$,
 $m_0 = N$, $m_1 = N \langle Z \rangle$, $m_2 = N (\sigma^2 - \langle Z \rangle^2)$, etc.



Global quantities:

Energy balance:

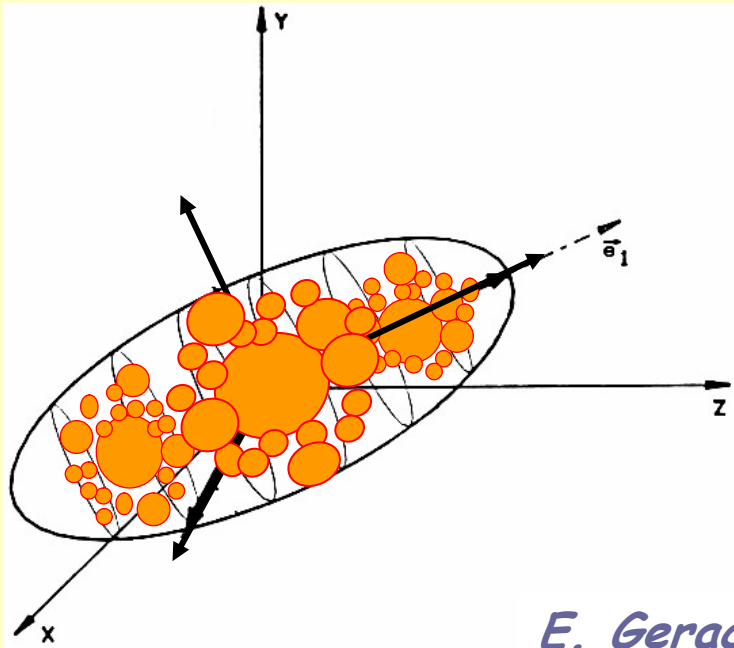
$$E^* + m_0 = \sum_{i=1}^M (m_i + k_i) + M_n (m_n + k_n)$$

"Flow" tensor:

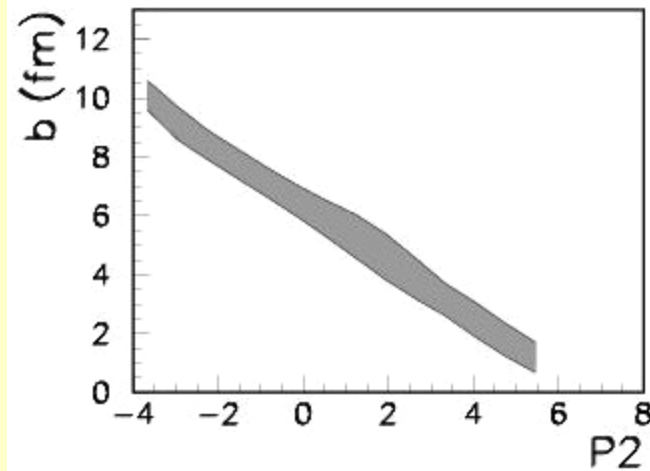
$$T_{ij} = \sum_{k=1}^M p_i^{(k)} p_j^{(k)} w^{(k)} \quad (i, j = 1, 3)$$

Sorting the events with a multidimensional analysis

Principal components/Neural networks



Filtered CMD model

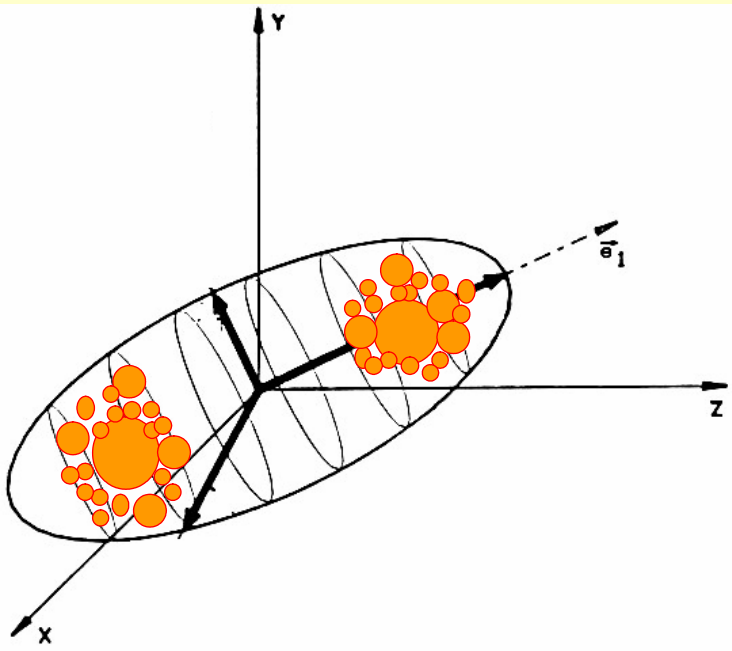


E. Geraci et al., NPA732(2004)173, NPA734 (2004)524

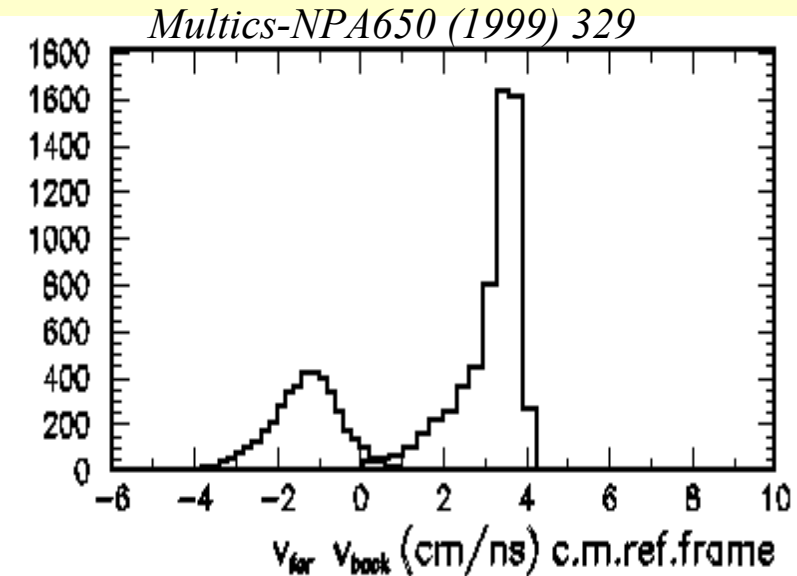
“Flow” tensor:

$$T_{ij} = \sum_{k=1}^M p_i^{(k)} p_j^{(k)} w^{(k)} \quad (i,j=1,3)$$

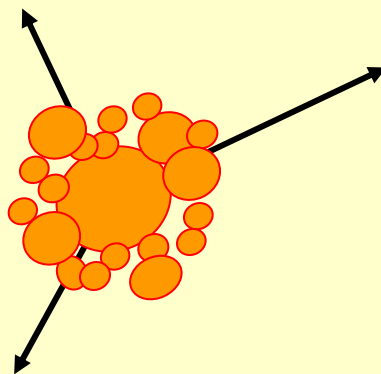
Sorting the events: multidimensional analysis



Peripheral
collisions:
many source.

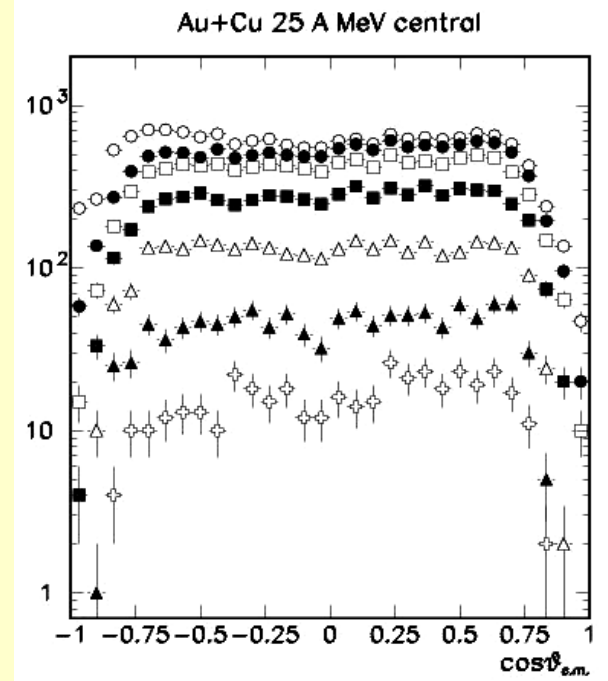


Central
collisions:
one source

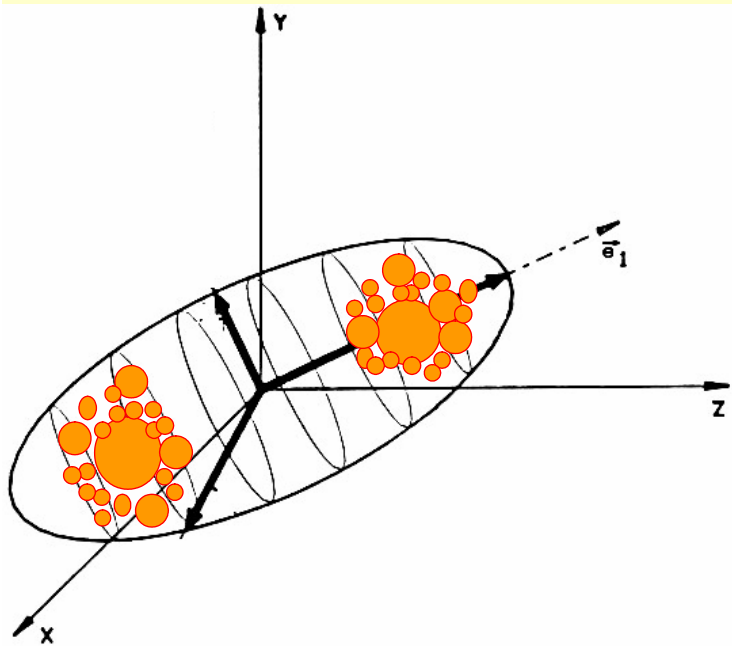


- $Z > 8$ open circles
- > 18 full points
- > 28 open squares
- > 38 full squares
- > 48 open triangles
- > 58 full triangles
- > 68 open crosses

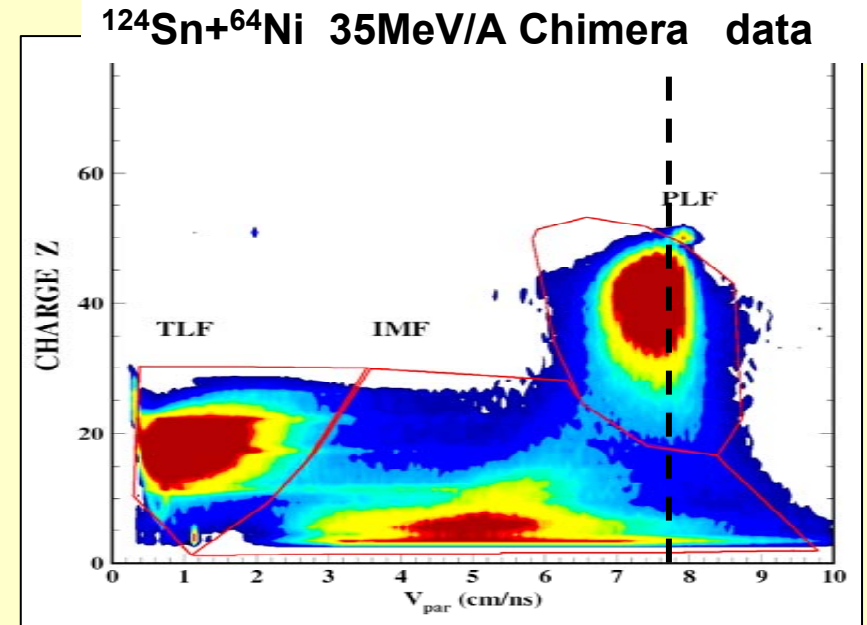
MulticsNPA734(2004)487



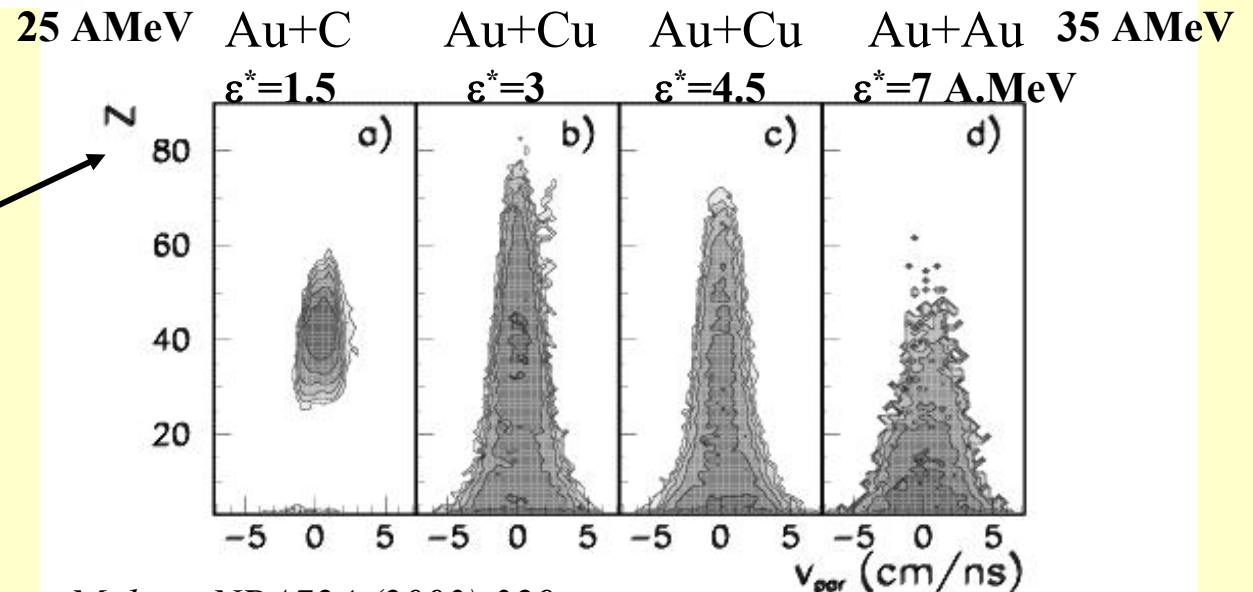
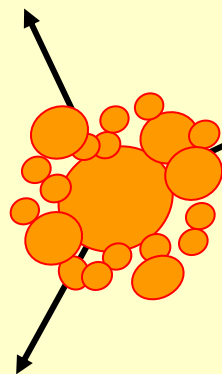
Sorting the events: multidimensional analysis



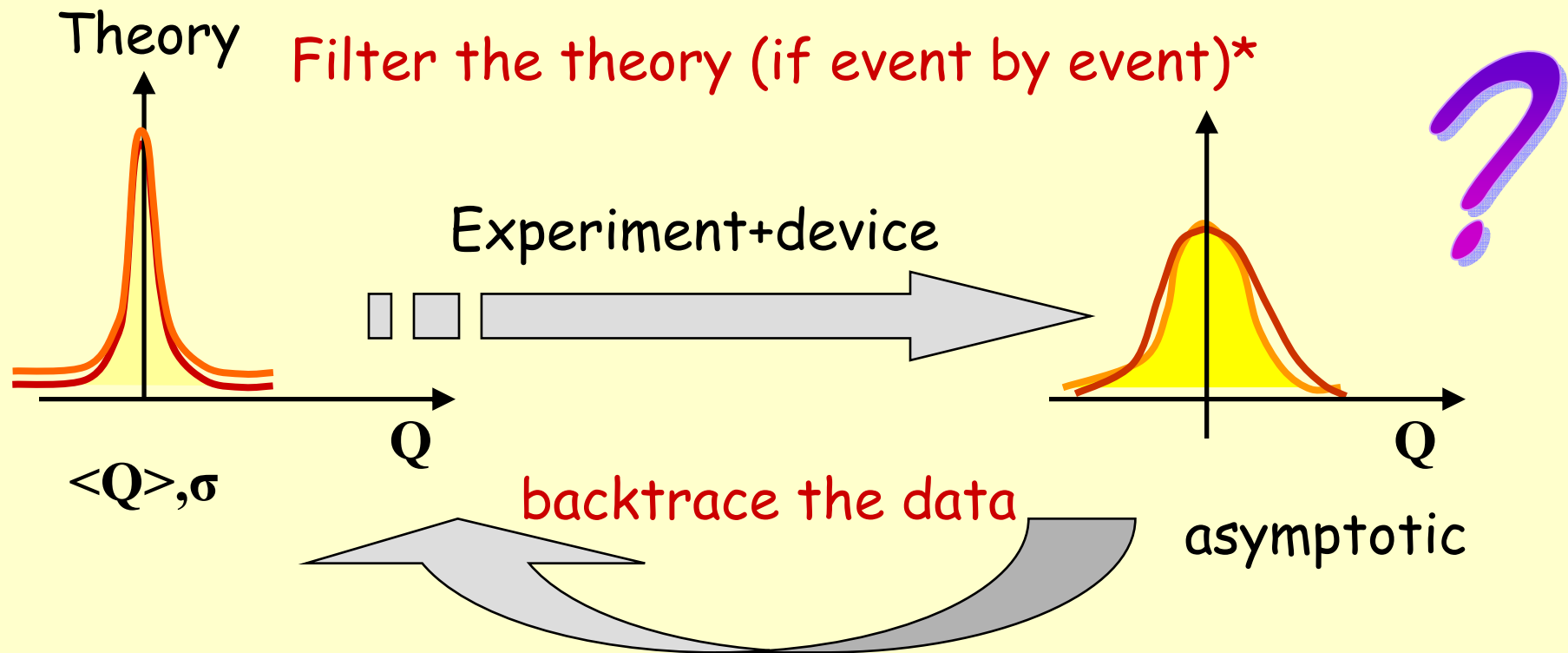
Peripheral collisions:
many sources



Central collisions:
one source



Dalla previsione del fenomeno alla rivelazione



* To be compared with data, Coulomb trajectories are needed.

True for:

- Statistical models
- Classical Molecular Dynamics

Not still true for: Lattice gas model

False for: percolation

Modelli statistici: esplorazione dello spazio delle fasi

Tutte le partizioni sono equiprobabili

SMM: J.Bondorf et al. NPA 443 (1985) 321, NPA 444 (1985) 460

MMMC: D.E. Gross Phys.Rep. 1993,

MMC: Al.&Ad. Raduta, NPA 1999

Nel caso piu' semplice, gli input sono: A_0 , E_0 =Energia termica di A_0

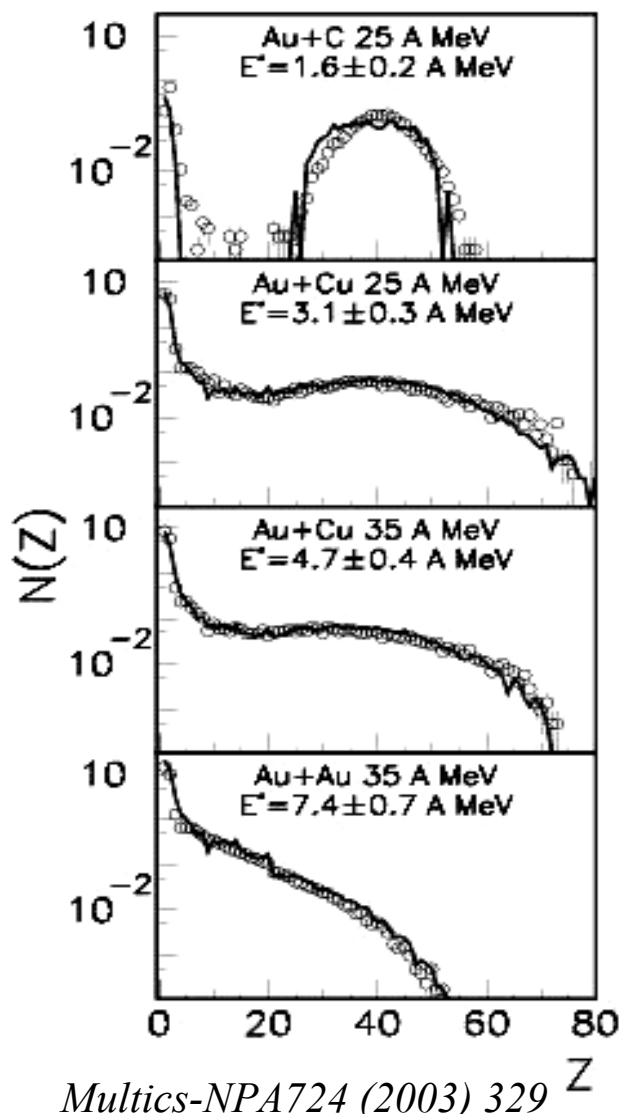
- Da considerazioni geometriche viene calcolata la probabilita' $p(A_0, M)$ che un sistema di A_0 nucleoni venga diviso in M parti intere,
servono formule ricorsive, infatti un sistema con $A_0 = 100$ si puo' rompere in 10^8 modi.
- Si estrae una molteplicita' M^* , secondo la distribuzione di probabilita' $p(A_0, M)$
- Si determina (random) la massa degli M^* frammenti (partizione di un intero in M^* parti),
- Si determina la energia associata alla partizione scelta e la si accetta se:
 $Q\text{-valore} + \text{Coulombiano} \leq E_0$
 l'energia rimanente $\Rightarrow$ moto termico & energia interna delle M parti

 seguono momenti termici iniziali,
 traiettorie coulombiane
 decadimenti secondari, etc.

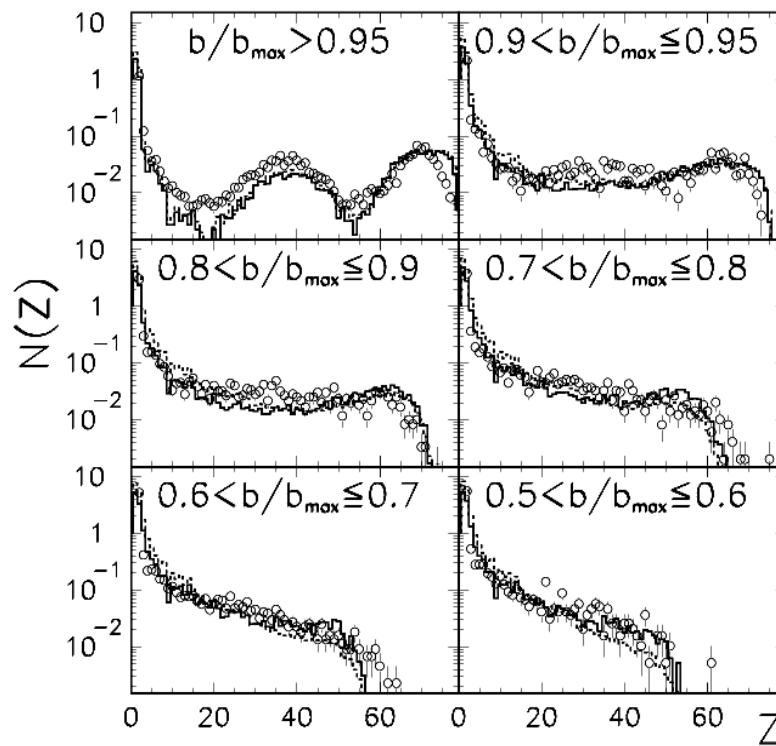
Failure: residual nuclear interaction

Checking equilibrium: uniform population of the available phase space

Central collisions



Multics-NPA724 (2003) 329



Au source:

peripheral
collisions

symbols: data
lines: thermal
model (SMM)

$\langle \epsilon^* \rangle = 1.5, 2.5,$
 $3.5, 4.5, 5, 6$
 A MeV

Multics-NPA650 (1999) 329

Static observables from
 liquid+vapor to droplets are
 reproduced by thermal models

Transizioni di fase: modello di Fisher

(M. E. Fisher, Rep. Prog. Phys. 30 (1967) 615)

La variazione di energia libera quando si forma una goccia di liquido di massa A da un gas di $A+B$ nucleoni (A e B in equilibrio, cioè alla stessa temperatura e pressione) è data dalla differenza fra:

$$G_{\text{con goccia}} = \mu_\ell A + \mu_g B + 4\pi R^2 \sigma(T) + T \tau \ln A$$

$$G_{\text{no goccia}} = \mu_g(A+B), \quad \text{dove}$$

$4\pi R^2 \sigma$ = tensione superficiale della goccia,

$T \tau \ln A$ = termine che tiene conto che la goccia è finita e la sua superficie è chiusa

Fisher inoltre scrive il termine di superficie $4\pi R^2 \sigma(T) = 4\pi r_0^2 A^{2/3} \sigma(T)$ come $a_s (T_c - T) A^\sigma$ con σ esponente che descrive il rapporto superficie/volume della goccia e tiene conto del fatto che al punto critico liquido e vapore sono indistinguibili.

si può anche includere δC = variazione della energia di Coulomb,
(J.Lopez and C.Dorso, World Scientific -2000)

Transizioni di fase: modello di Fisher

M. E. Fisher, Rep. Prog. Phys. 30 (1967) 615

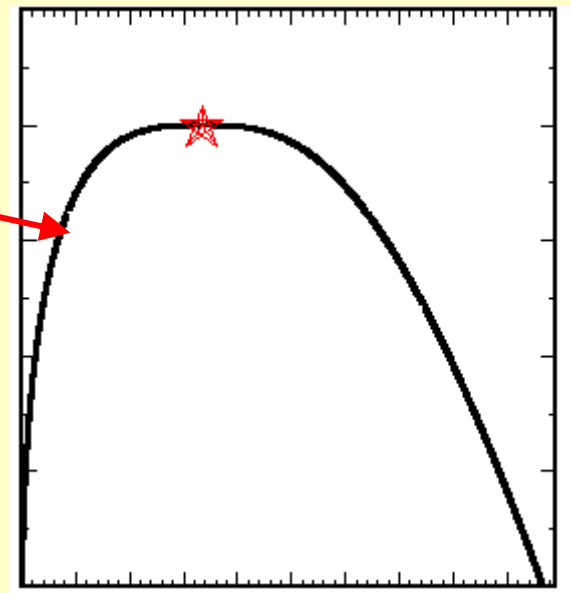
La probabilit  (insieme gran canonico) di formazione di una goccia di massa A , a partire da un gas di $A+B$ nucleoni  :

$$P(A) = Y_0 \exp\left[-\frac{\Delta G}{T}\right] = \\ = Y_0 A^{-\tau} \exp\left[\frac{\mu_g - \mu_\ell}{T} A - \frac{a_s (T_c - T) A^\sigma}{T}\right]$$

Alla coesistenza $\mu_g = \mu_\ell$

Al punto critico $\mu_g = \mu_\ell$
 $a_s(T - T_c) = 0$

$$\Rightarrow P(A) = Y_0 A^{-\tau}$$



Classical Molecular Dynamics (A. Bonasera)

In the classical molecular dynamics model it is assumed that the nucleus is made up of A nucleons that behave classically. These particles move under the influence of a pairwise spherical interaction potential V given by [4]

$$V_{np}(r) = V_r [\exp(-\mu_r r)/r - \exp(-\mu_r r_c)/r_c] \\ - V_a [\exp(-\mu_a r)/r - \exp(-\mu_a r_a)/r_a],$$

$$V_{nn}(r) = V_{pp}(r) = V_0 [\exp(-\mu_0 r)/r - \exp(-\mu_0 r_c)/r_c], \quad (1)$$

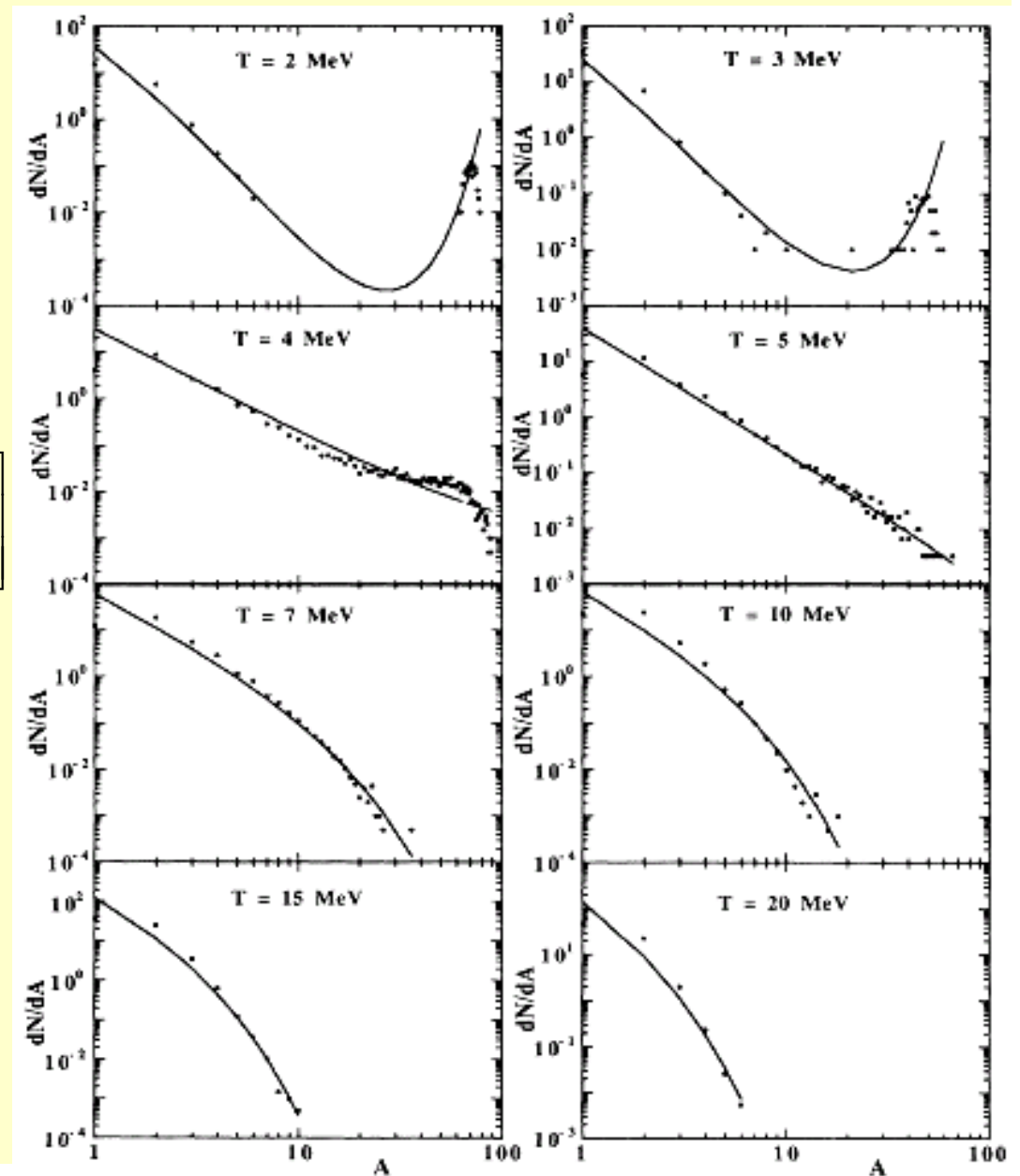
where $r_c = 5.4$ fm is a cutoff radius and V_{np} is the potential acting between a neutron and a proton while V_{nn} is the potential acting between two identical nucleons. The first interaction is attractive at large r and repulsive at small r , while the latter is purely repulsive. In this way, no bound state of identical nucleons can exist.

M. Belkacem, V. Latora, A. Bonasera PRC 52(1995)271

Classical Molecular Dynamics (A. Bonasera)

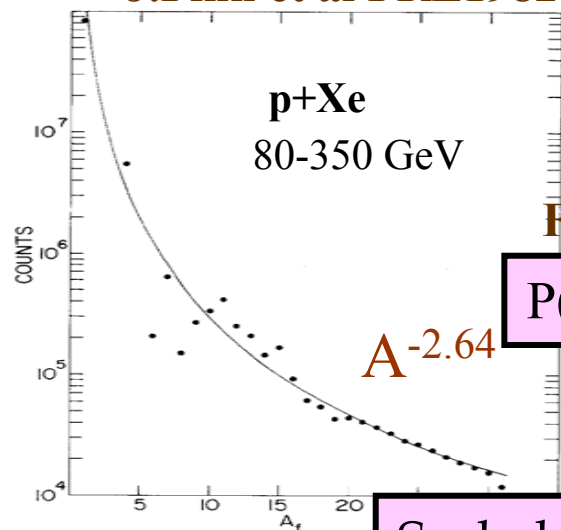
$$P(A) = Y_0 \exp\left[-\frac{\Delta G}{T}\right] =$$

$$= Y_0 A^{-\tau} \exp\left[\frac{\mu_g - \mu_l}{T} A - \frac{a_s (T_c - T) A^\sigma}{T}\right]$$



M. Belkacem, V. Latora, A. Bonasera PRC 52(1995)271

J.Finn et al PRL1982



Data: Self similarity and scaling

Power-laws are free of scales

All the information falls on a single curve

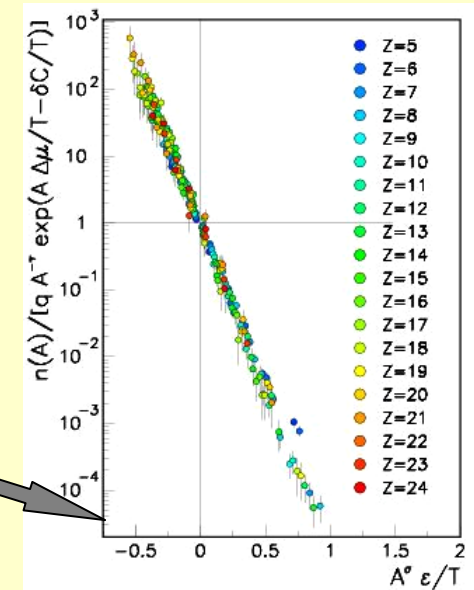
Multics NPA724 (2003) 45

Fisher 1967

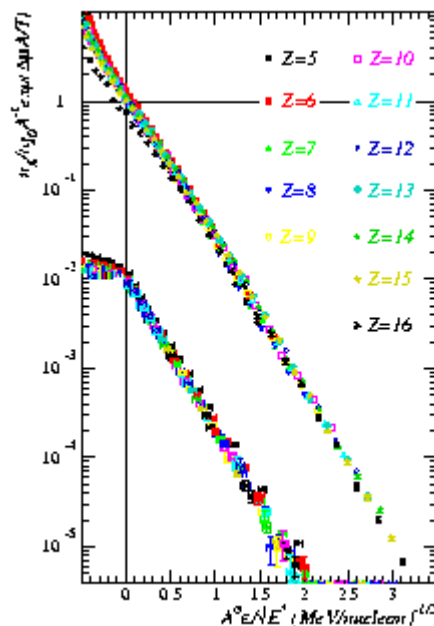
$$P(A) = Y_0 A^{-\tau} \exp(A \Delta\mu/T - c_0 \varepsilon A^\sigma / T)$$

Scaled yield: $P(A)/[Y_0 A^{-\tau} \exp(A \Delta\mu/T)]$

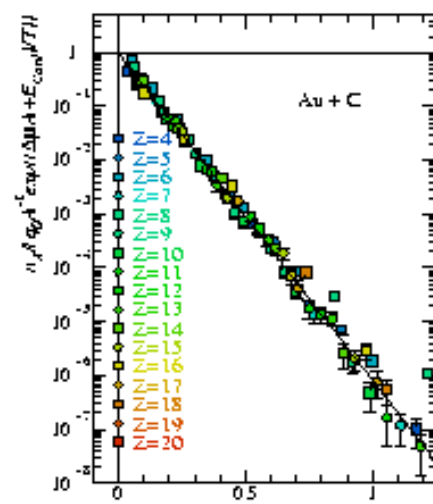
Scaled temperature: $\varepsilon A^\sigma / T$



IsIs PRL2



EoS PRC2003



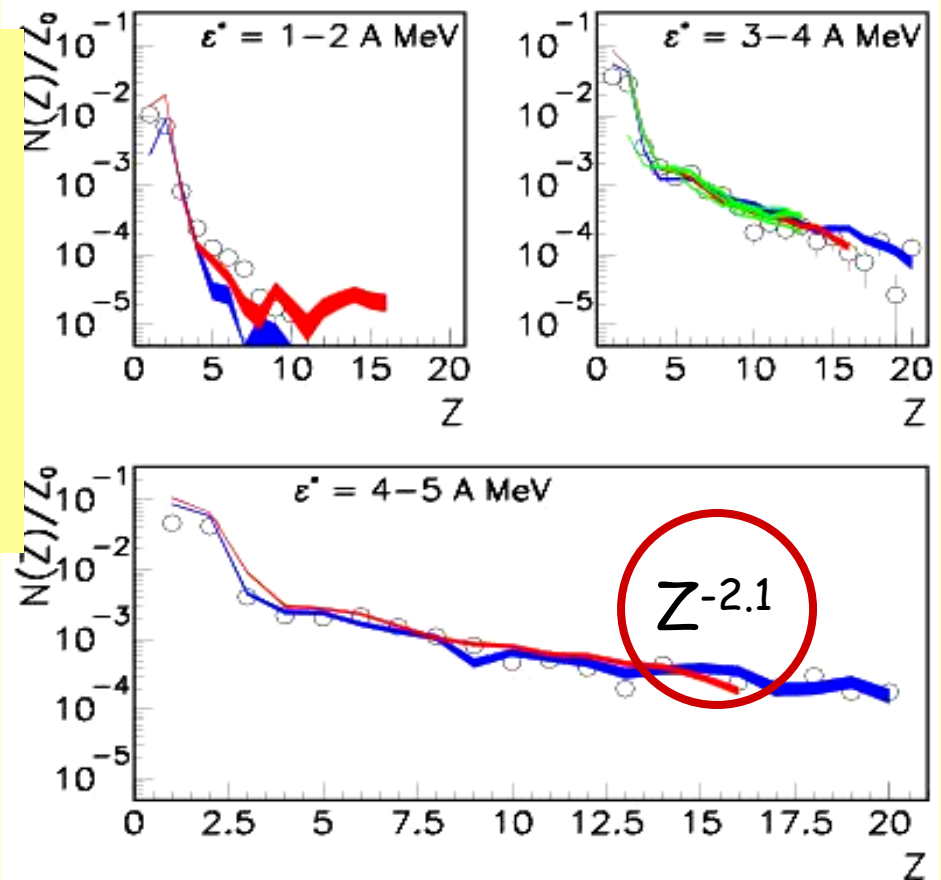
Au Liquid-Gas

τ	2.1 ± 0.1	2.196 ± 0.024
σ	0.66 ± 0.02	0.647 ± 0.006
ε_c^*	$4.4 \pm 0.1 \text{ A MeV}$	

1-st generation 4π devices & stable beams

- More information on
WCI 2003-2005 in:
<http://cyclotron.tamu.edu/wci3/>

*world-wide review of the field of
dynamics and thermodynamics with
nucleonic degrees of freedom*
- For all multifragmentation
experiments, **independently on the
entrance channel**, the region in
which power-laws are observed in
reaction observables corresponds
to **$E^*/A = 5 \pm 1$ A.MeV**
Within a phase-transition scenario,
this value represents the transition
energy.



Multics: Central from $Z_0=85$ to $Z_0=100$ (lines)

Multics: Au peripheral $Z_0=79$ (symbols)

Isis: π +Au 8 GeV/c NPA734(2004)487

Fasa: p,α+Au 4-14 GeV NPA709(2002)392

Critical exponents from moment analysis

$$P(A) = Y_0 \exp\left[-\frac{\Delta G}{T}\right] =$$

$$= Y_0 A^{-\tau} \exp\left[\frac{\mu_g - \mu_\ell}{T} A - \frac{a_s (T_c - T) A^\sigma}{T}\right]$$

$P(A)$ ha un massimo ad una temperatura $T_{\max}(A)$ diverso per ogni A .

Dall'andamento di $T_{\max}(A)$ vs. A si può ricavare σ .

Dai due esponenti σ, τ si possono ricavare altri esponenti critici, tramite le relazioni:

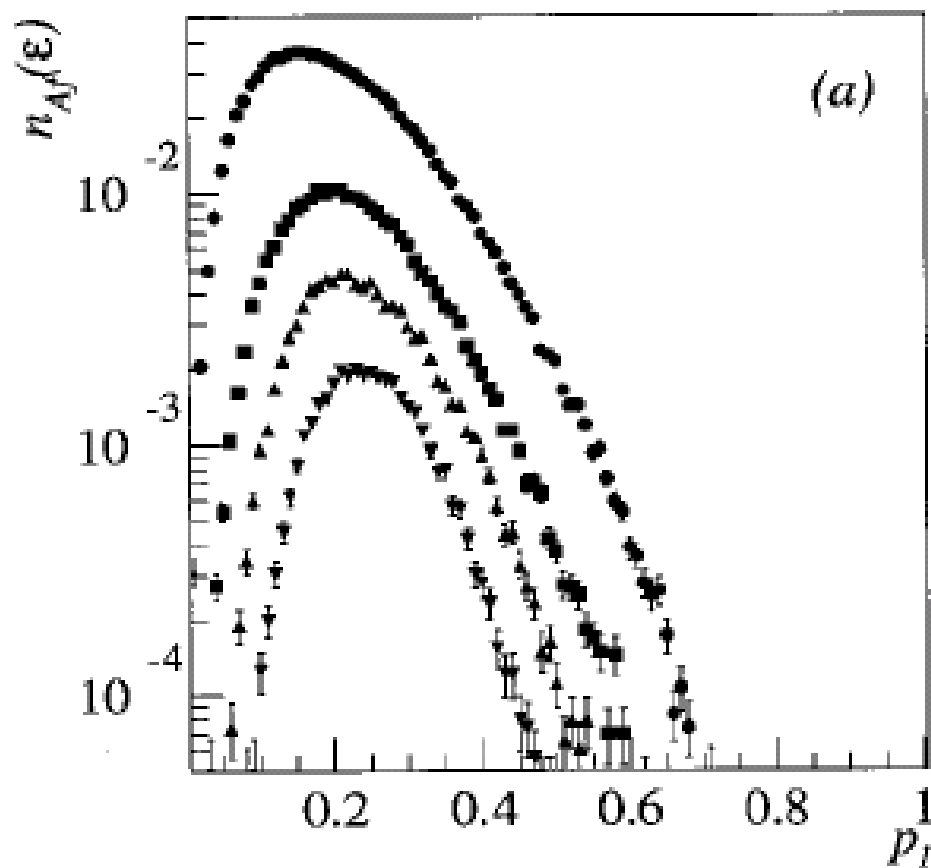
$\beta = (\tau - 2)/\sigma$, $\gamma = (3 - \tau)/\sigma$
e studiare (per dati e modelli) i momenti:

$$m_1 = \sum n_s s \sim |\epsilon|^\beta$$

$$m_2 = \sum n_s s^2 \sim |\epsilon|^{-\gamma}$$

$$m_k = \sum n_s s^k \sim |\epsilon|^{(\tau-1-k)/\sigma}$$

per verificare la compatibilità delle loro distribuzioni **e le loro correlazioni** con un comportamento critico



Critical « correlations » of static moments

Classical Molecular Dynamics (A. Bonasera)

$$\sigma = (\tau - 2)/\beta, \quad \gamma = (3 - \tau)/\sigma$$

$$m_1 = \sum n_s s \sim |\epsilon|^\beta$$

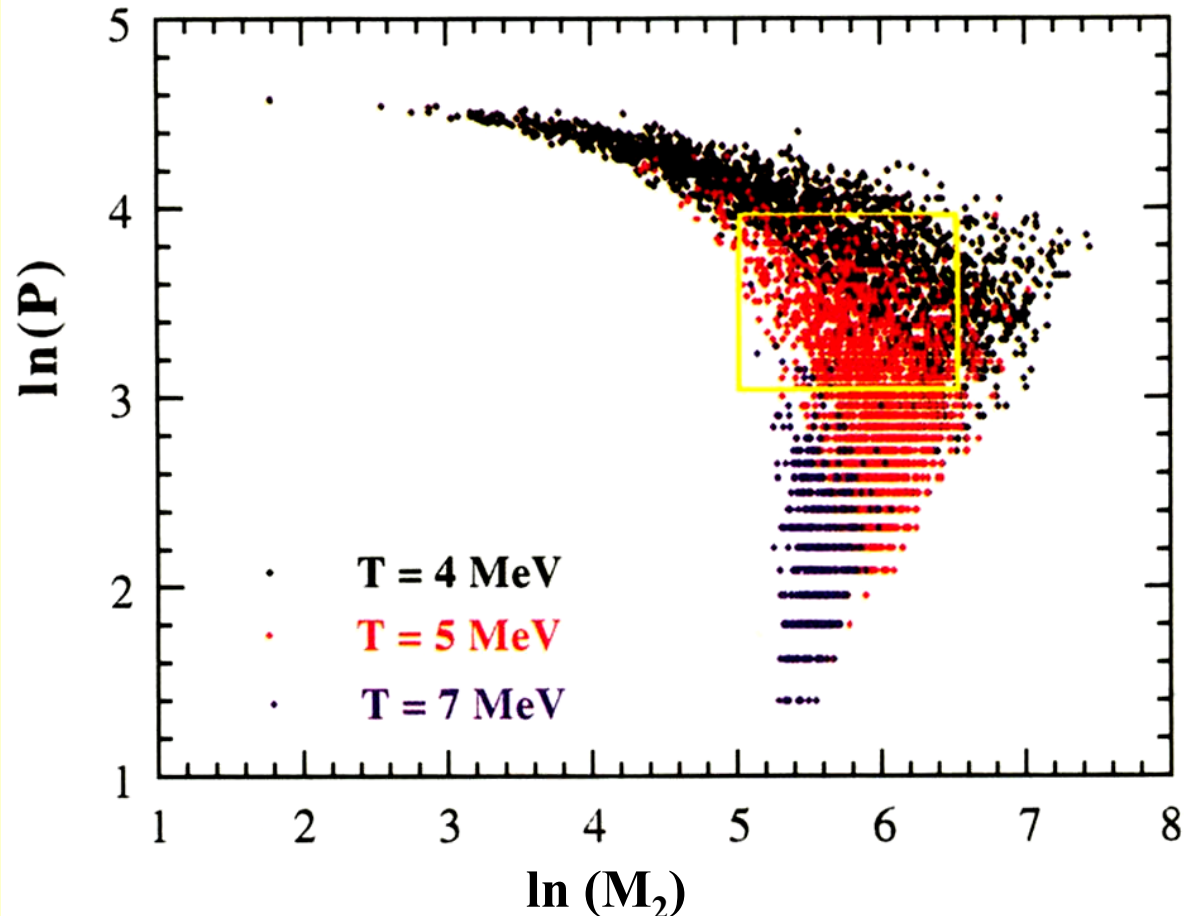
$$m_2 = \sum n_s s^2 \sim |\epsilon|^{-\gamma}$$

$$m_k = \sum n_s s^k \sim |\epsilon|^{(\tau-1-k)/\sigma}$$

liquid branch and
gas branch
meet at the
critical region

the logarithm of the largest cluster P in each event versus the logarithm of the second moment S_2 of the cluster size distribution of the remaining fragments in the event. This plot is generally called Campi scatter plot.

M. Belkacem, V. Latora, A. Bonasera PRC 52(1995)271



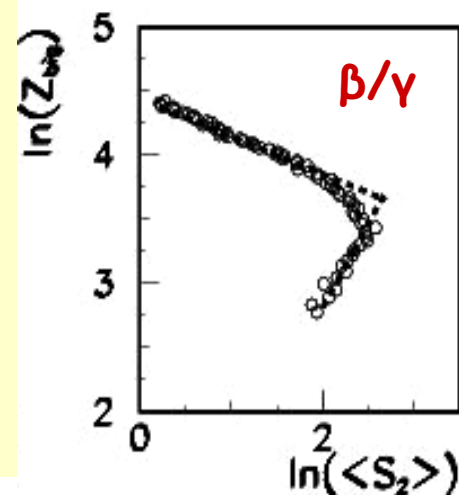
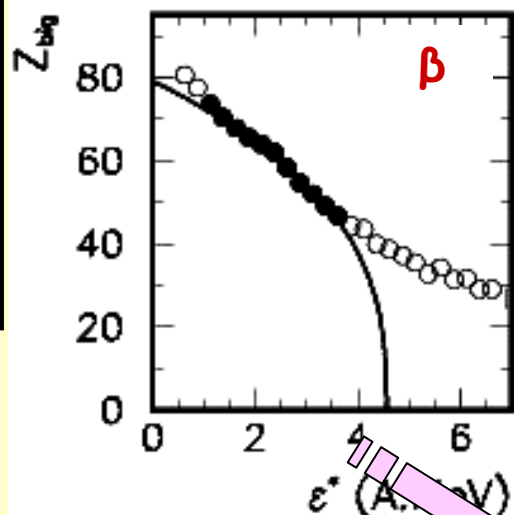
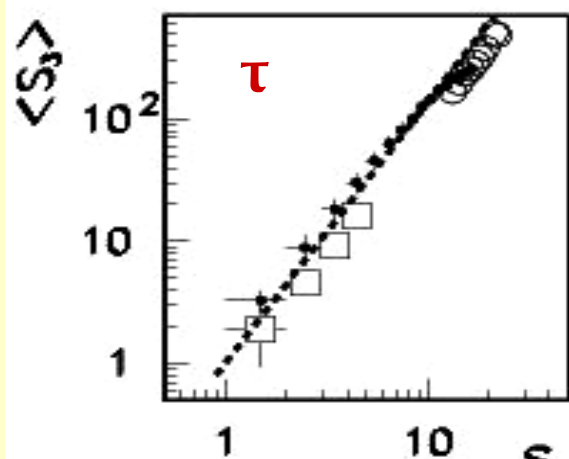
Critical exponents from moment analysis (data)

$$m_1 = \sum n_s s \sim |\epsilon|^\beta$$

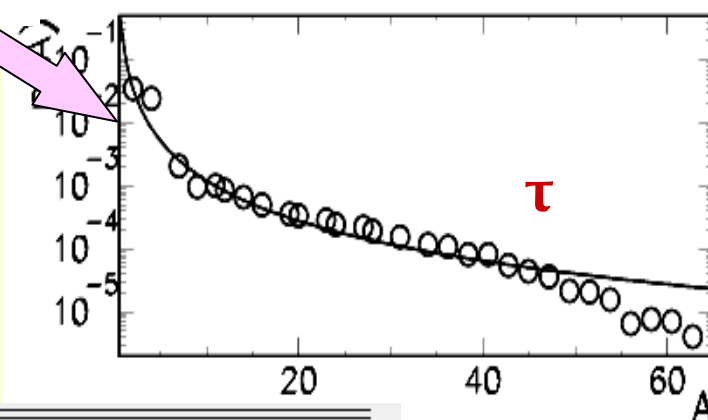
$$m_2 = \sum n_s s^2 \sim |\epsilon|^{-\gamma}$$

$$m_k = \sum n_s s^k \sim |\epsilon|^{(\tau-1-k)/\sigma}$$

$$\sigma = (\tau-2)/\beta$$



S_2 *EoS: PRC 62 (2000) 064603*



	Au	Liquid-Gas
τ	2.13 ± 0.04	2.196 ± 0.024
γ	1.29 ± 0.01	1.24 ± 0.01
β	0.31 ± 0.04	0.305 ± 0.005
$\epsilon^*_{crit} = 4.5 \pm 0.2 \text{ A MeV}$		

Multics-NPA650 (1999) 329

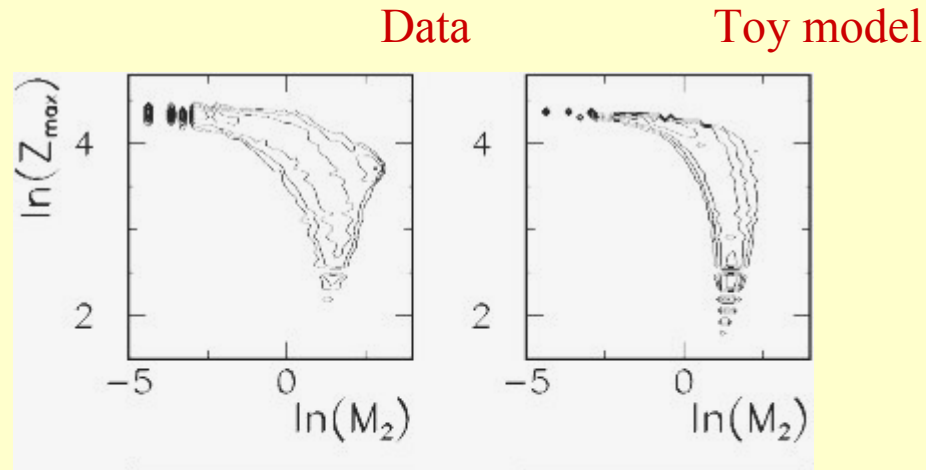
Exponent/system	Au + C
τ	2.2 ± 0.1
σ	0.64 ± 0.05
$\gamma = \frac{3-\tau}{\sigma}$	1.3 ± 0.2
γ_+ (matching)	1.4 ± 0.3
γ_- (matching)	1.4 ± 0.3
$\langle \gamma \rangle$ (matching)	1.4 ± 0.3
$\Delta \gamma$ (matching)	0.0 ± 0.4
ν (hyperscaling)	0.63 ± 0.07

Exponents	This work
τ	2.22 ± 0.46 [Eq. (30)]
	2.31 ± 0.03 [Eq. (12)]
	2.13 ± 0.10 [Eq. (26)]
β	0.33 ± 0.01
γ	1.15 ± 0.06
σ	0.68 ± 0.04

Nimrod PRC71(2005)054606

Perche' studiare le
correlazioni?

Segnali necessari
e/o (?)
sufficienti?



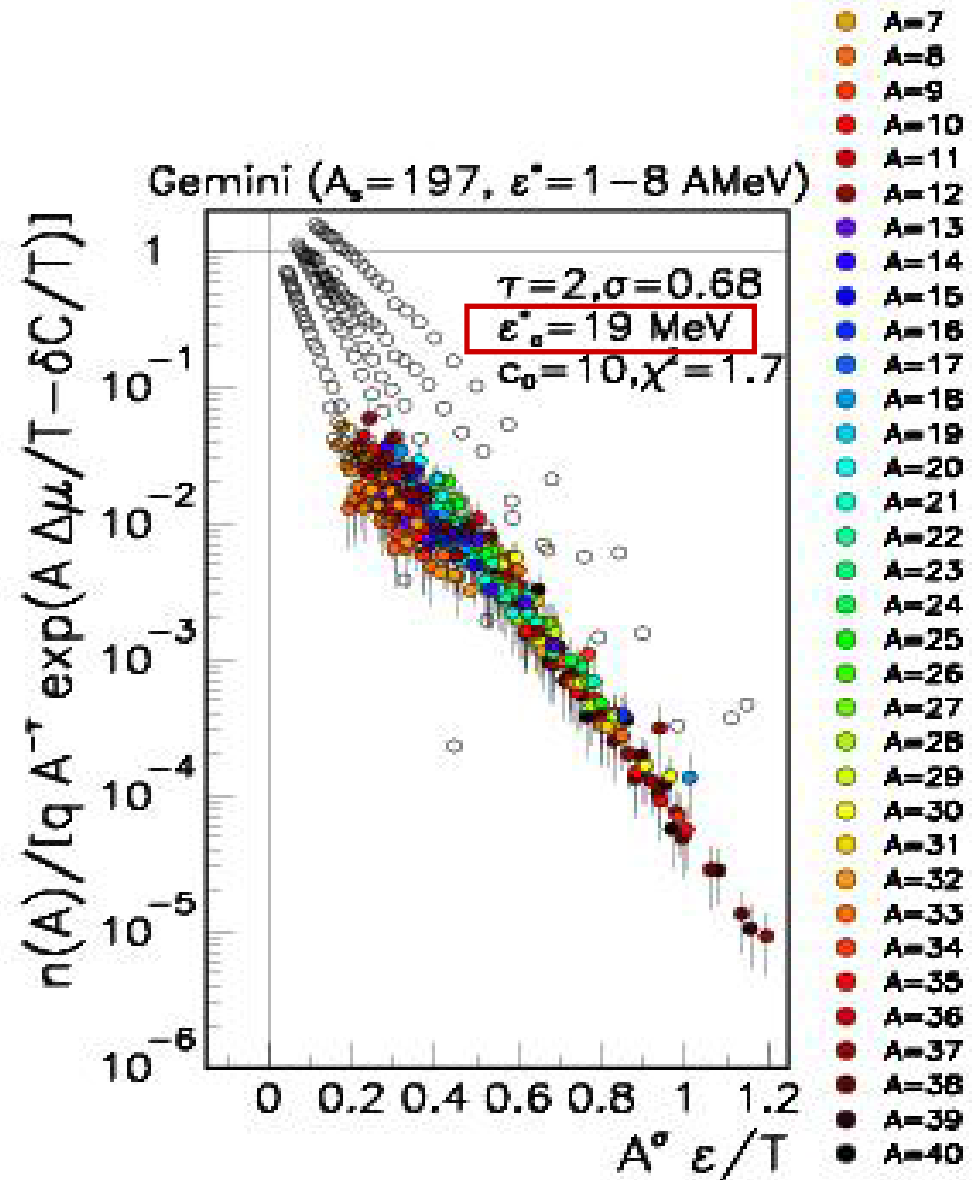
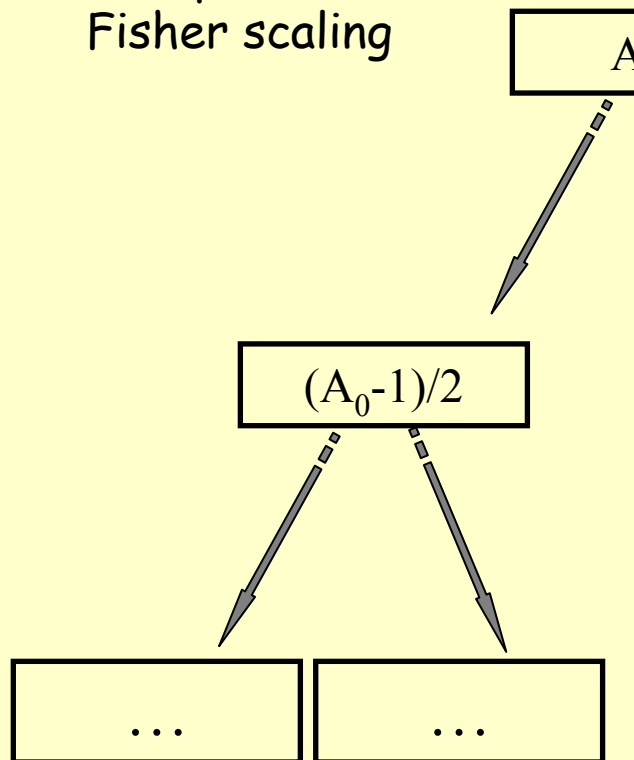
Review paper:

A. Bonasera, M. Bruno, C. O. Dorso and P. F. Mastinu, *Riv. Nuovo Cimento* 23, 1 (2000)

No power law !!!

Sequential binary decays: evaporation&fission (Gemini)

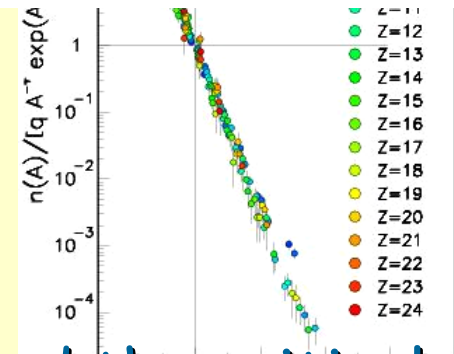
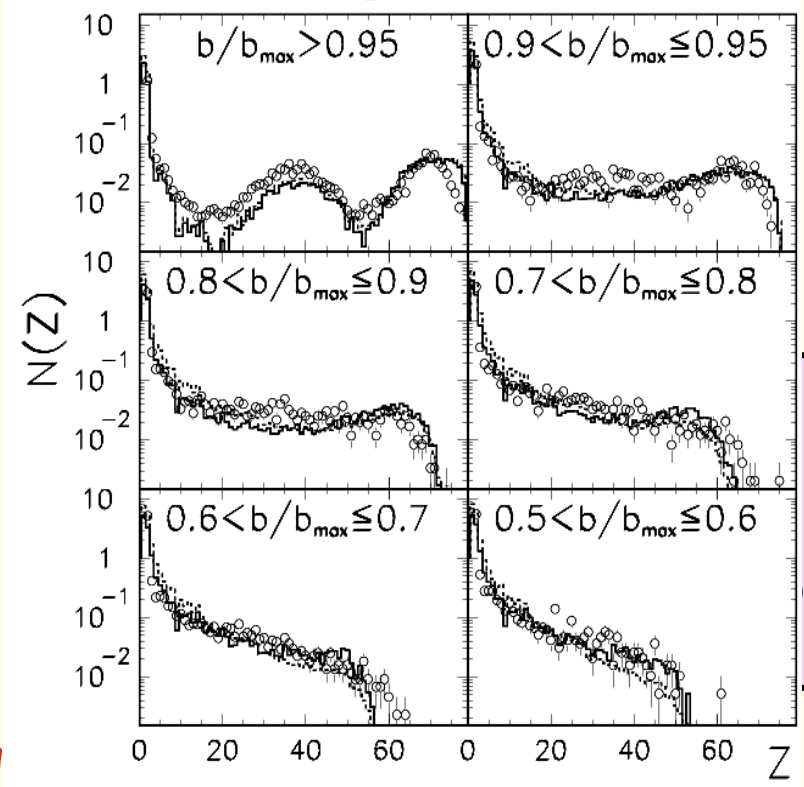
Size distributions
exponential,
not free of scales,
not compatible with
Fisher scaling



Liquid-gas phase transition: is the game over?

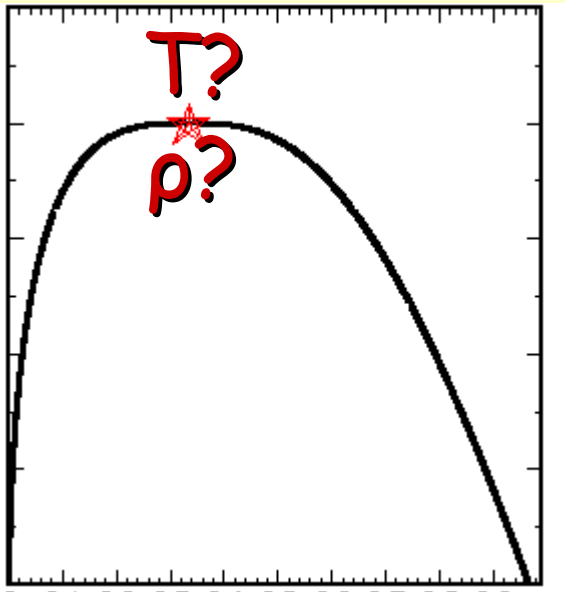
We observed:

- suppression/enhancement of branching ratios (liquid-gas)
- fragment yields described by critical exponents (liquid-gas universality class)
- equilibrated collisions behaving as a universal process, independent of the entrance channel



Can we conclude that the system reached the critical point?

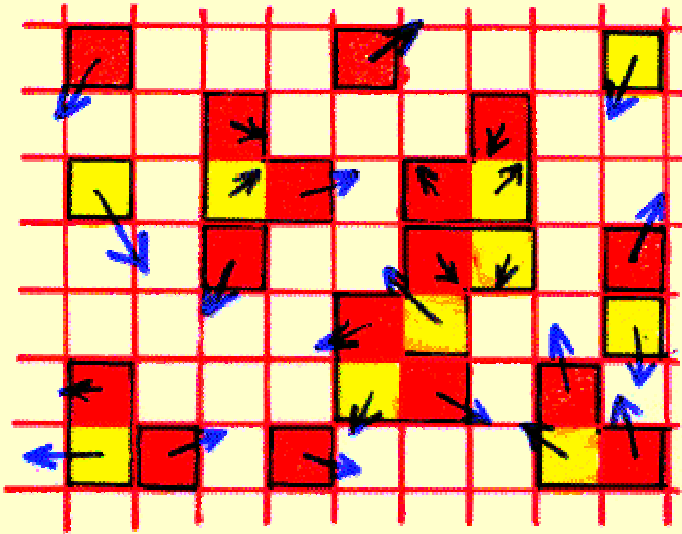
Liquid-gas phase transition: is the game over?



We have to

- **thermodynamically** characterize the system (T, ρ)
- look at more observables
- look at **thermodynamical** models

Lattice gas model



- Cubic lattice (L^3),
- A_0 nucleons occupy A_0 sites, $\rho = A_0/L^3$
- occupancy ($\tau = 0,1$) from the partition sum:

$$Z(A_0, \beta) = \sum W(E) \exp(-\beta E)$$

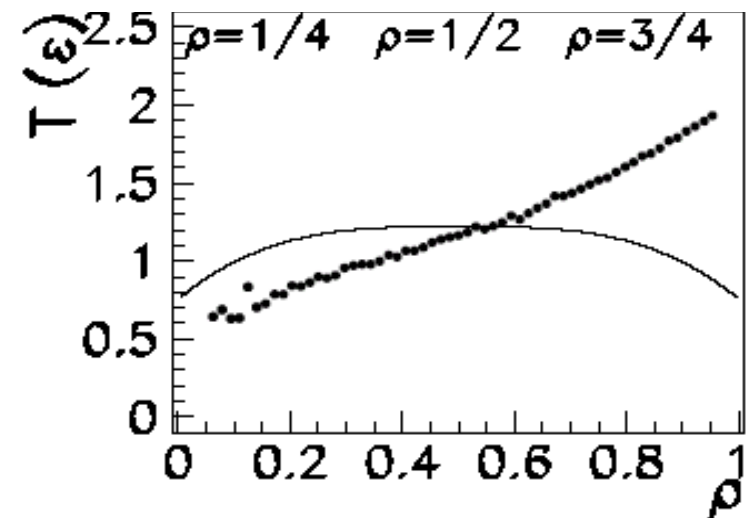
- interaction among nearest neighboring

$$H = \sum_{i=1}^n \frac{p_i^2}{2m} \tau_i^2 + \sum_{i \neq j} \epsilon \tau_i \tau_j$$

The system is **finite**:

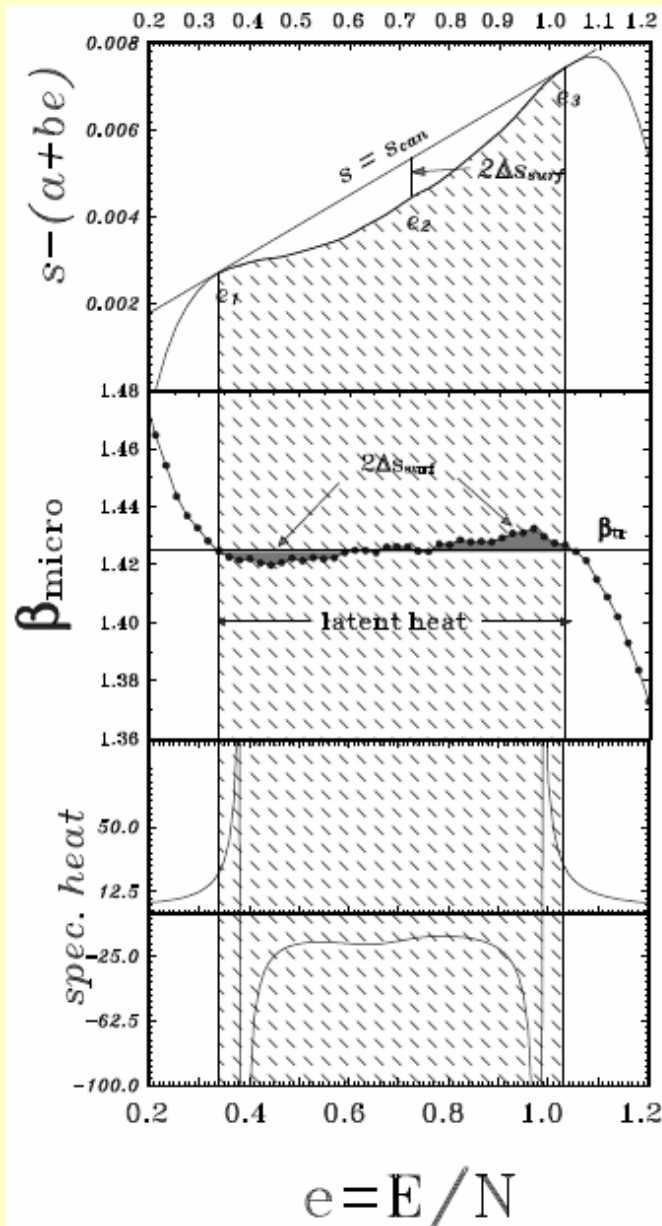
power-laws are found at all densities

Also inside the coexistence region



*F. Gulminelli, V. Duflot, Ph. Chomaz
PRL 1999*

Finite systems-Thermodynamical anomalies



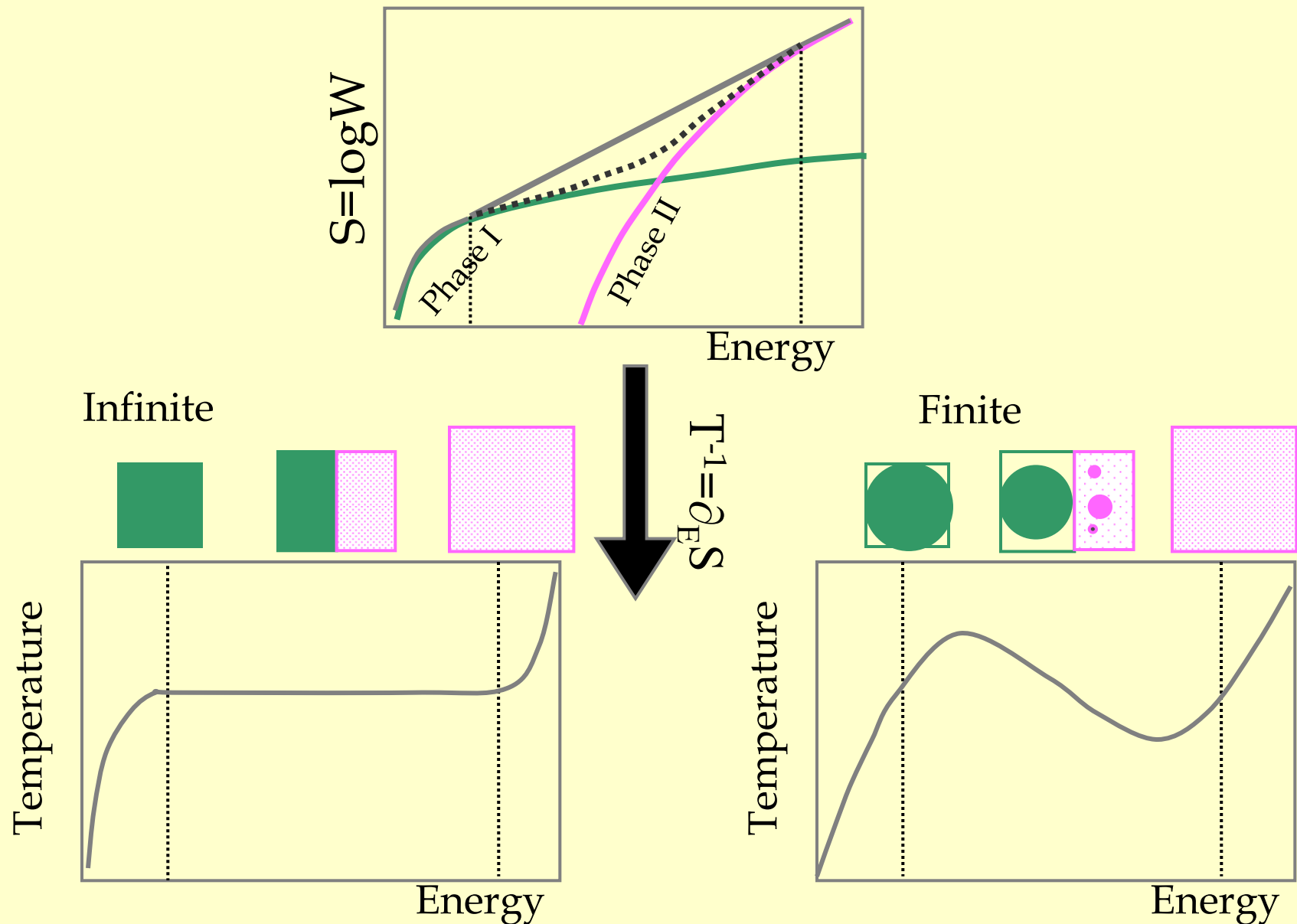
D.H.E. Gross, A. Ecker, and X.Z. Zhang. Microcanonical thermodynamics of first order phase transitions studied in the potts model. *Ann. Physik*, 5:446–452, 1996, and <http://xxx.lanl.gov/abs/cond-mat/9607150>.

The surface gives a negative contribution to S

$$\beta = T^{-1} = \partial_E S$$

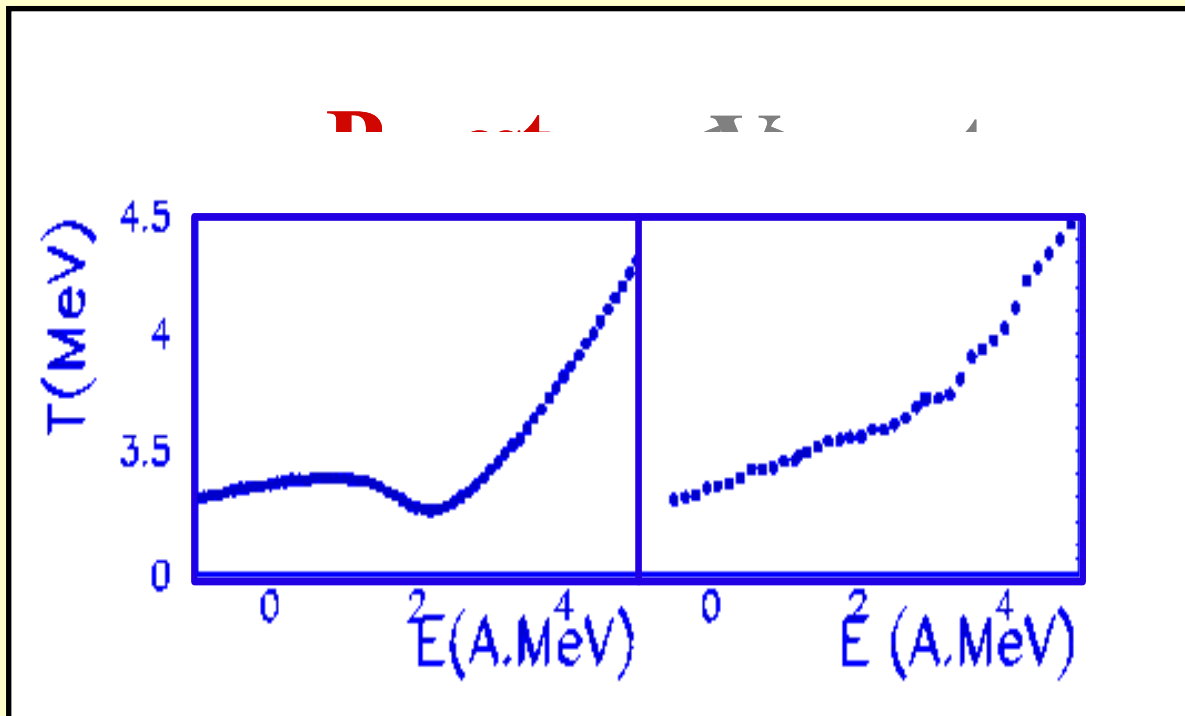
$$C = \partial_T E$$

Finite systems-Thermodynamical anomalies



Lattice gas model

The caloric curve
depends on the transformation



(ideal gas)

$$\partial T = \partial E / C_V$$

$$\partial T = \partial E / C_p$$

Microcanonical heat capacity of finite systems

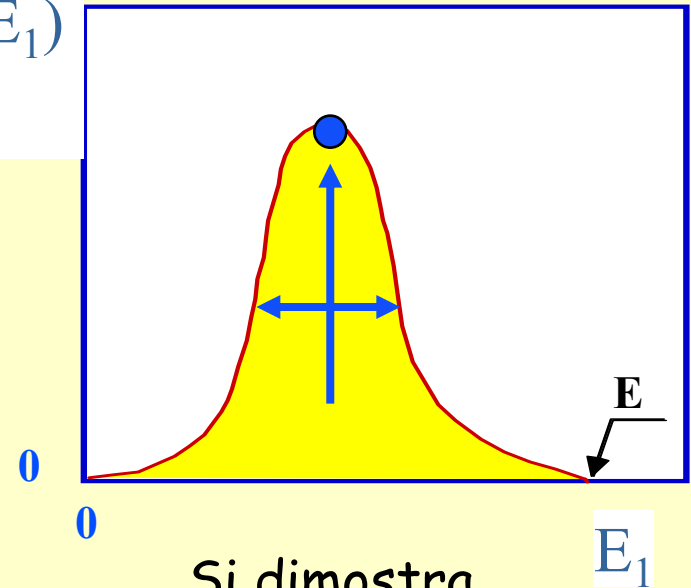
Ph. Chomaz, F. Gulminelli NPA 647 (1999) 153

Suddividiamo un sistema con energia E in due sottosistemi 1 e 2, tali che

$$E = E_1 + E_2$$

$$P(E_1) = \frac{W_1(E_1)W_2(E_2)}{W(E)}$$

$P(E_1)$



Si dimostra analiticamente che:

(S.K. Ma Statistical mechanics- Chap.6)

- In corrispondenza del valore piu' probabile:

$$1/T_1 = \partial S_1 / \partial E_1 = \partial S_2 / \partial E_2 = 1/T_2 = 1/T$$

- Le fluttuazioni di E_1

$$\sigma_1^2 = \frac{C_1 C_2}{T^{-2} (C_1 + C_2)}$$

$$C_1 = \partial E_1 / \partial T$$

$$C_2 = \partial E_2 / \partial T$$

- Il calore specifico del sistema $C \approx C_1 + C_2$

$$C = \frac{C_1^2}{(C_1 - \sigma_1^2 / T^2)}$$

Microcanonical heat capacity of finite systems

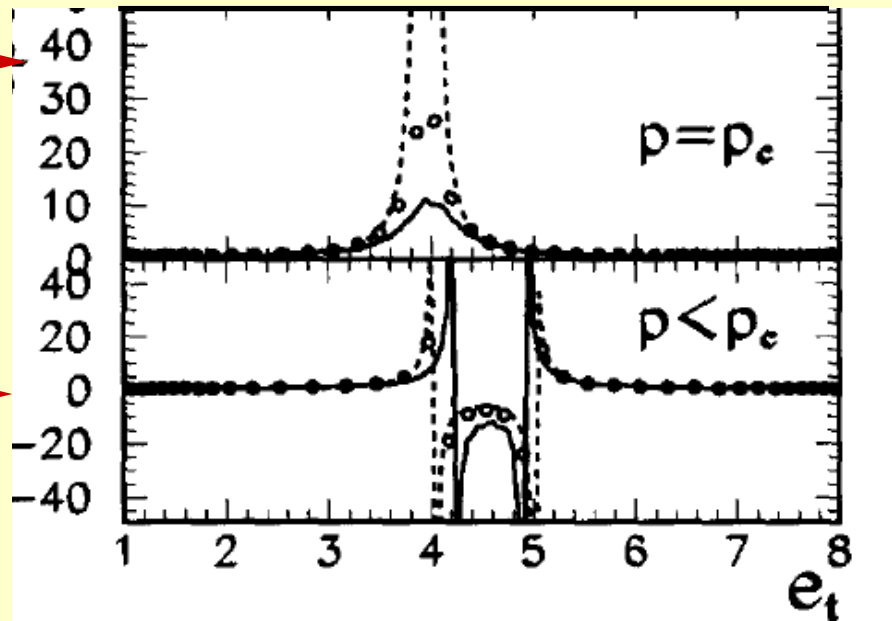
Ph. Chomaz, F. Gulminelli NPA 647 (1999) 153

$$C = C_1^2 / (C_1 - \sigma_1^2 / T^2)$$

2nd order



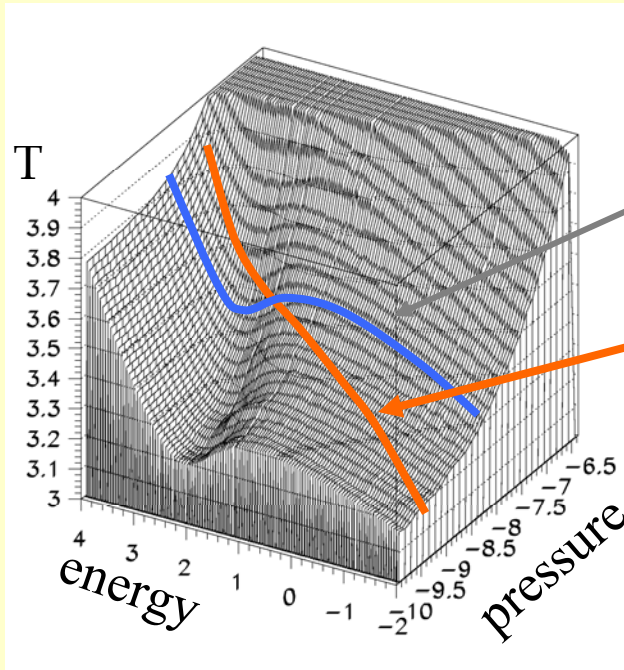
1st order



² It should be noticed that in the canonical case the energy fluctuations of the subsystem 1 are independent of the total fluctuations and only give information on the heat capacity of the subsystem itself. This will not be the case in a microcanonical ensemble because then the fluctuations in the subsystem 1 are correlated to the fluctuations in subsystem 2 through the total energy conservation constraint.

Advantages of studying abnormal fluctuations

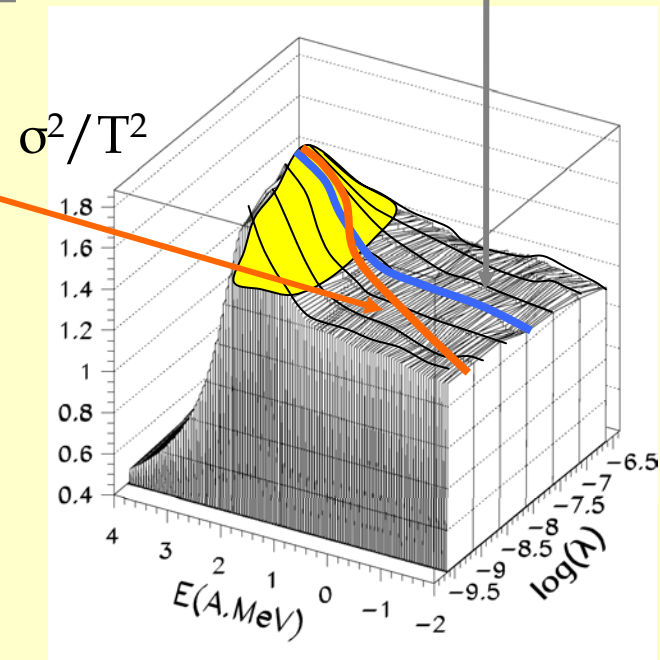
Lattice Gas Model



The caloric curve depends on the transformation

$$p = \text{cte}$$

$$V = \text{cte}$$

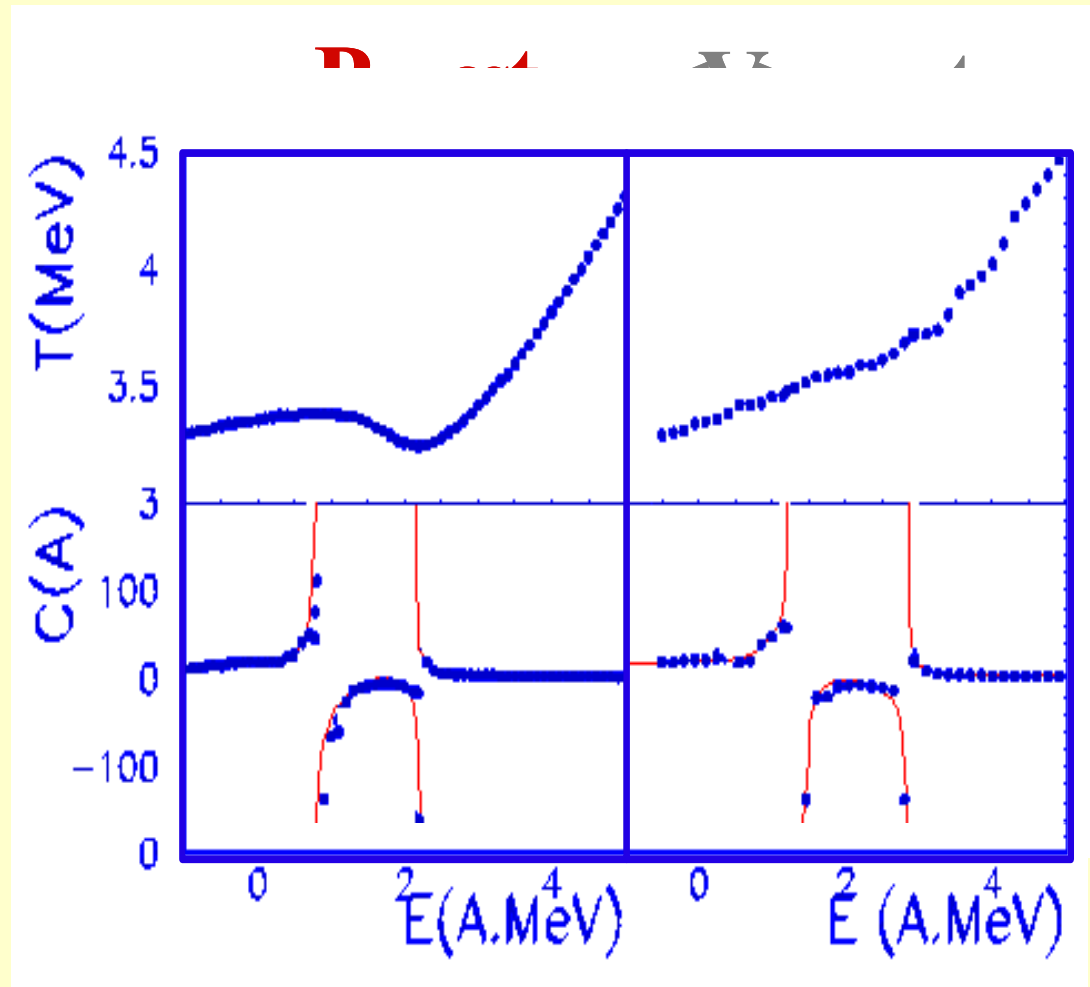


Fluctuations are unique

Microcanonical thermodynamics of finite systems

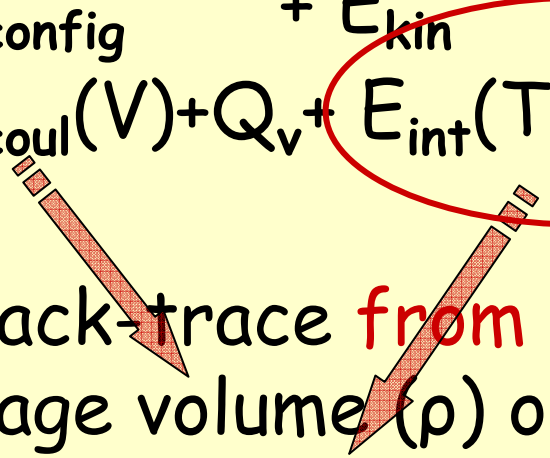
Lattice Gas Model

- The caloric curve depends on the transformation
- Fluctuations are independent on the transformation, they are state variables



Microcanonical thermodynamics of finite systems

Events sorted as a function of E^* (calorimetry)

$$E^* = E_{\text{config}} + E_{\text{kin}}$$
$$E^* = E_{\text{coul}}(V) + Q_v + E_{\text{int}}(T) + E_{\text{tr}}(T)$$


We can back-trace **from data**

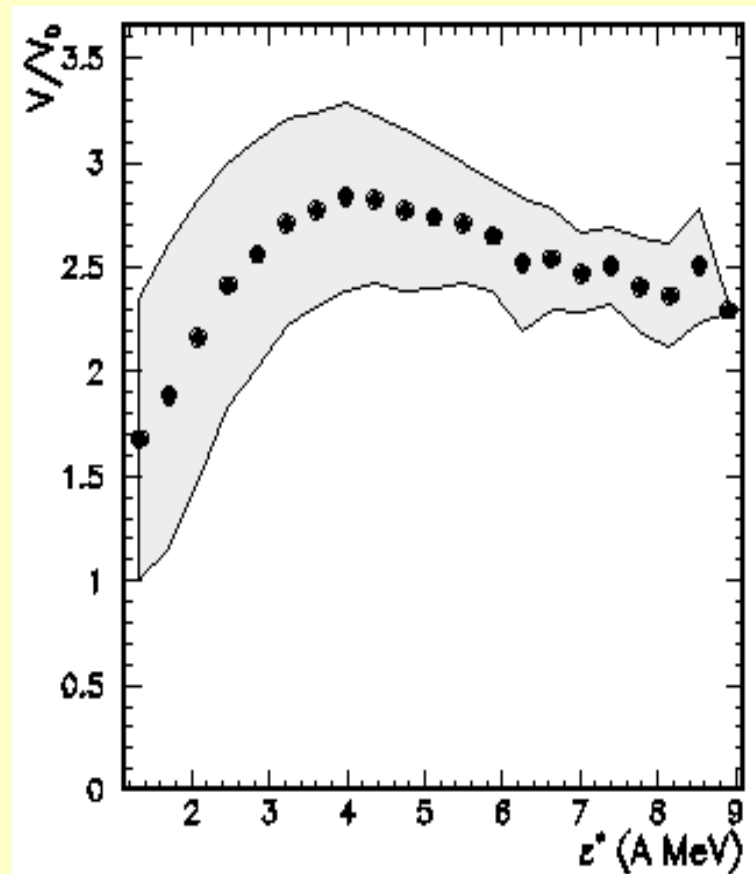
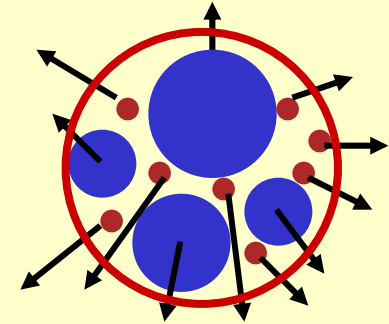
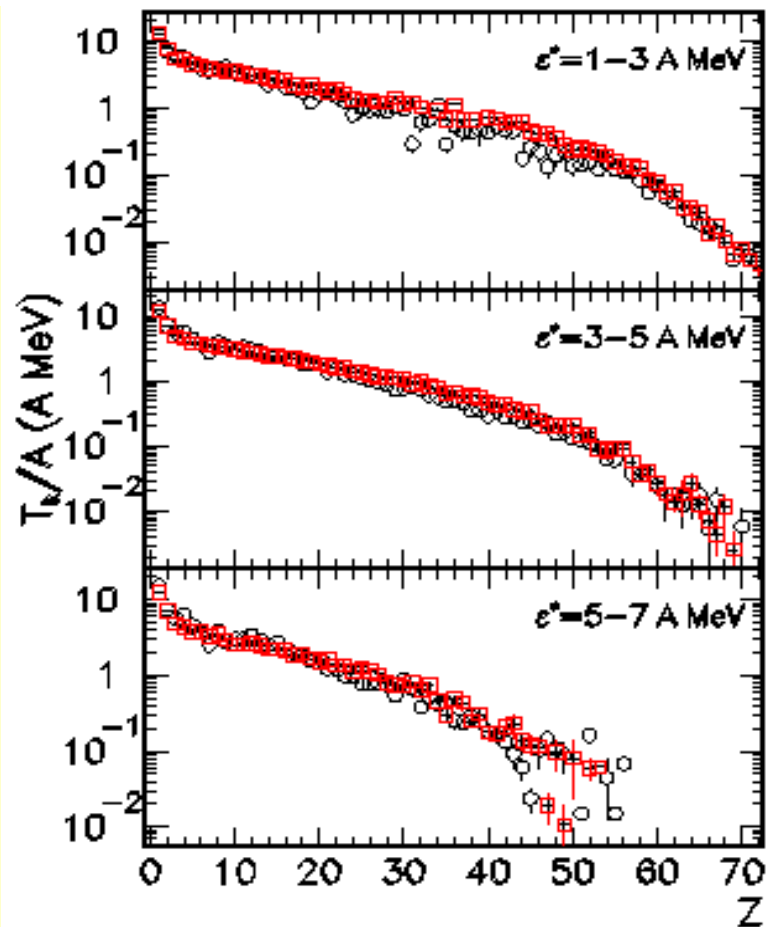
- the average volume $\langle v \rangle$ of the system
- the temperature T

under the constraint of energy conservation

Multics-Nucl.Phys.A699(2002)795

Early information from measured observables: average volume

Circles=Multics data
Squares=Coulomb trajectories



Early information from measured observables : Temperature

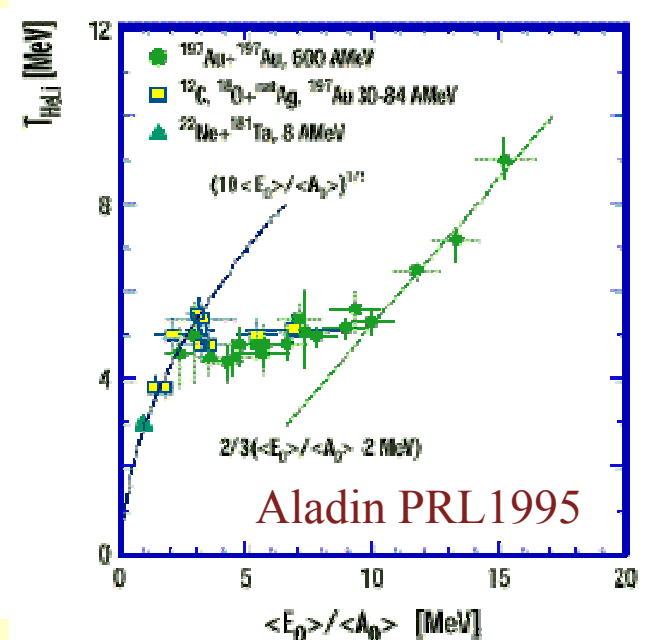
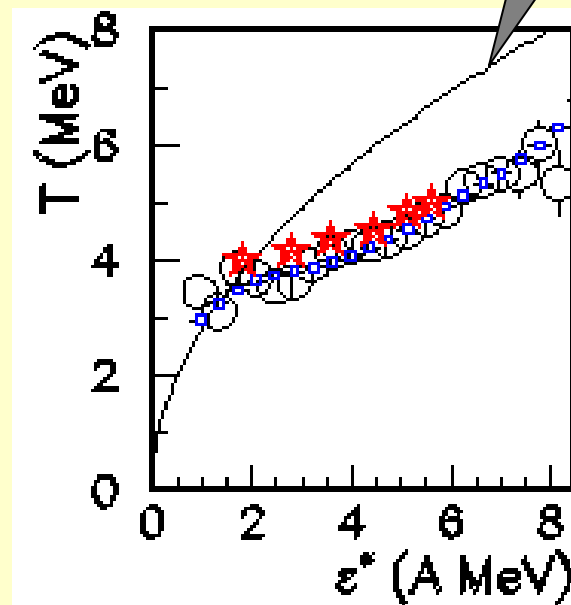
■ $\langle E_{kin} \rangle = (3/2) \langle m-1 \rangle T + \langle \Sigma a A_{IMF} \rangle T^2$

Multics-NPA699(2002)795

○ $T = \left\langle \frac{E_{tr}}{(3/2)m - 1} \right\rangle$

★ Isotope thermometer
P.M.Milazzo, PRC58(1998) 953

● Indra correlation data
N.Marie, PRC58(1998)256



T, E_{int} from independent measurements/methods

Microcanonical heat capacity from fluctuations

$$E^* = E_{\text{config}} + E_{\text{kin}} \quad (\sigma^2_{\text{config}} = \sigma^2_{\text{kin}})$$

$$\begin{cases} E_{\text{config}} = Q_v + E_{\text{coul}}(V) \\ E_{\text{kin}} = E_{\text{trasl}}(T) + E_{\text{internal}}(T) \end{cases}$$

The system being thermodynamically characterized:

Ph. Chomaz, F. Gulminelli, NPA 647(1999) 153

$$C = C_{\text{kin}}^2 / (C_{\text{kin}} - \sigma_{\text{kin}}^2 / T^2)$$

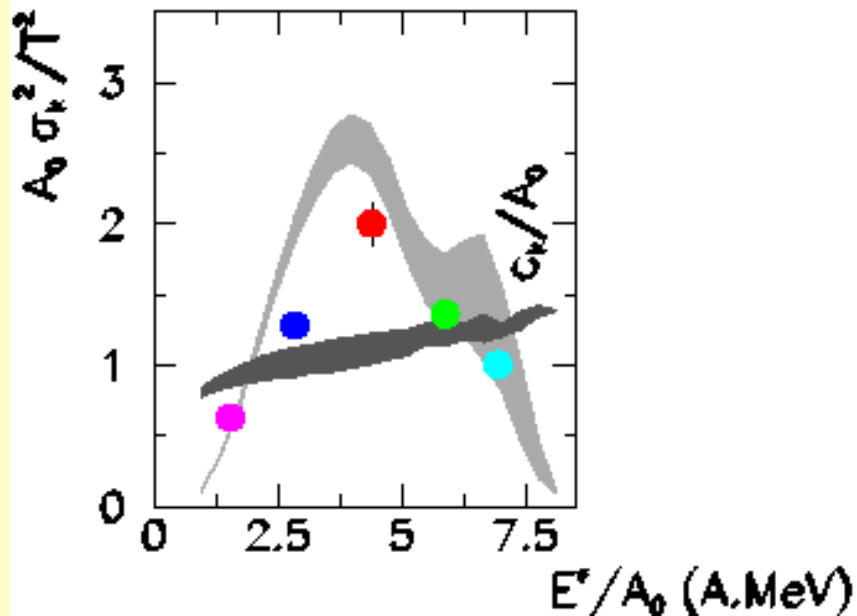
Microcanonical fluctuations
larger than the canonical
expectation?

where:

$$C_{\text{kin}} = dE_{\text{kin}} / dT$$

Multics-PLB473 (2000) 219; NPA699 (2002) 795; NPA734 (2004) 512

Heat capacity from fluctuations



Multics:

PLB473 (2000) 219

NPA699 (2002) 795

NPA734 (2004) 512

NPA699(2002)795

Grey area: peripheral collisions

Points: central collisions:

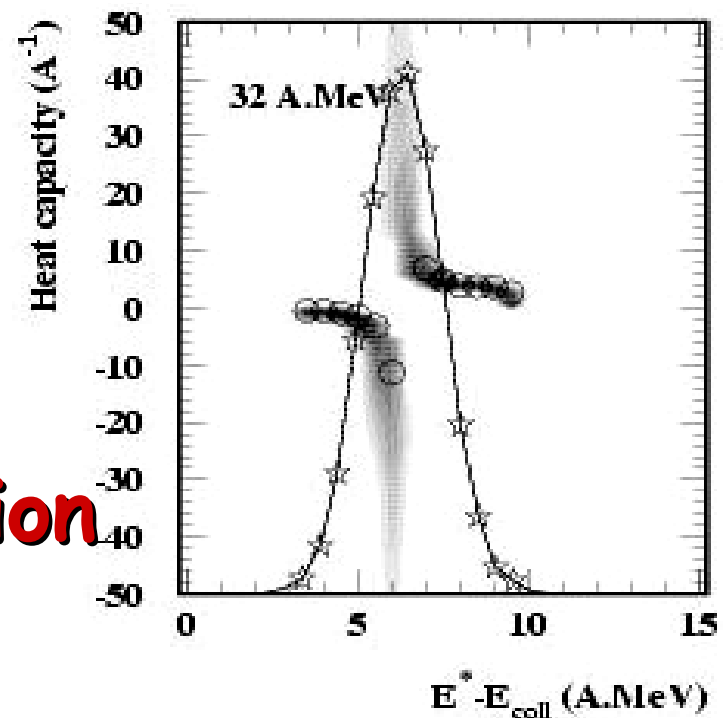
Au+C

Au+Cu

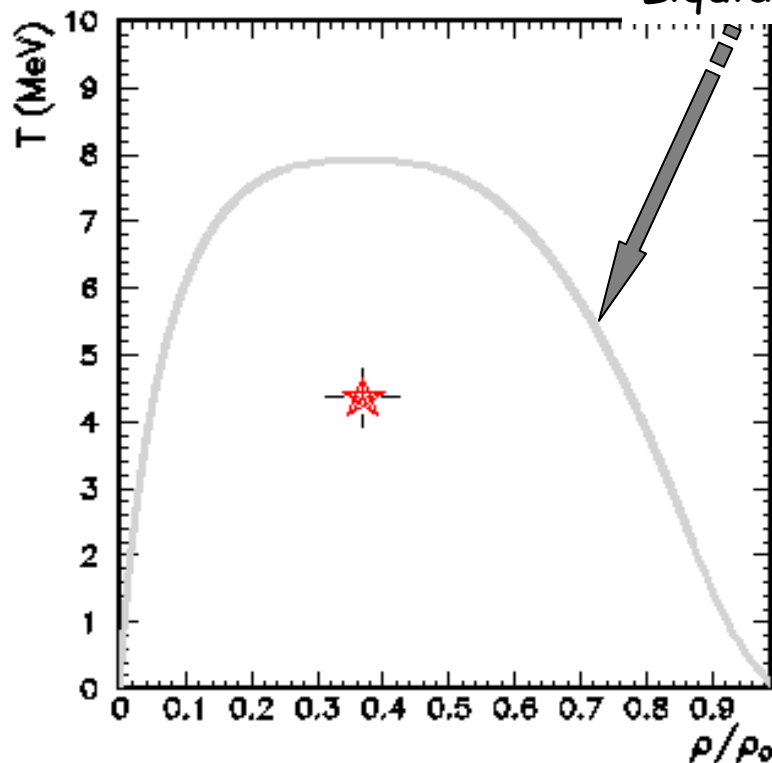
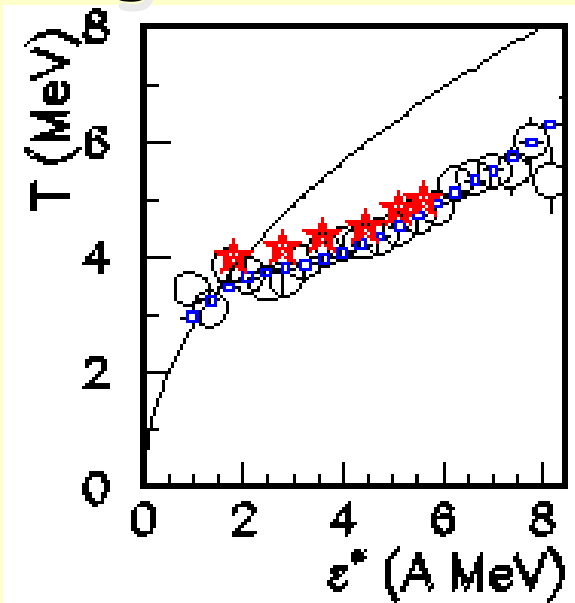
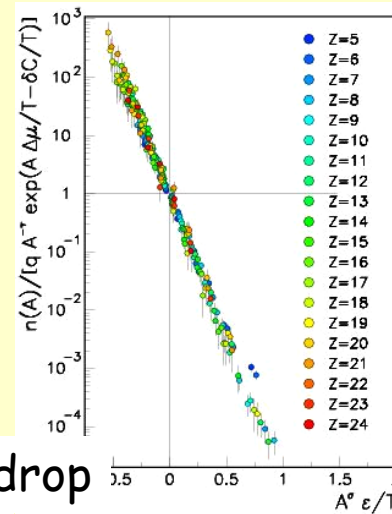
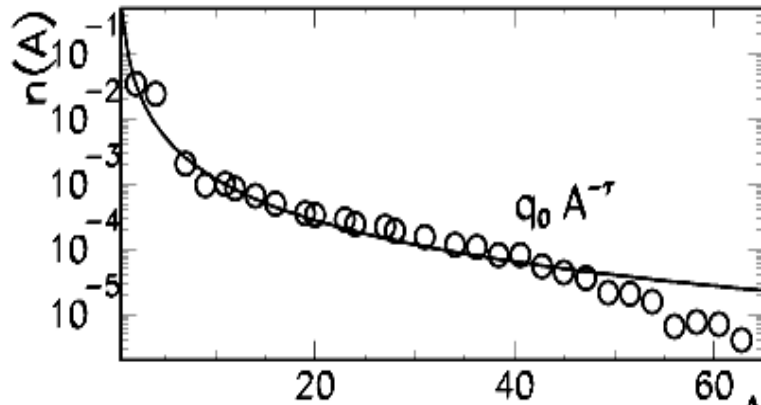
Au+Au



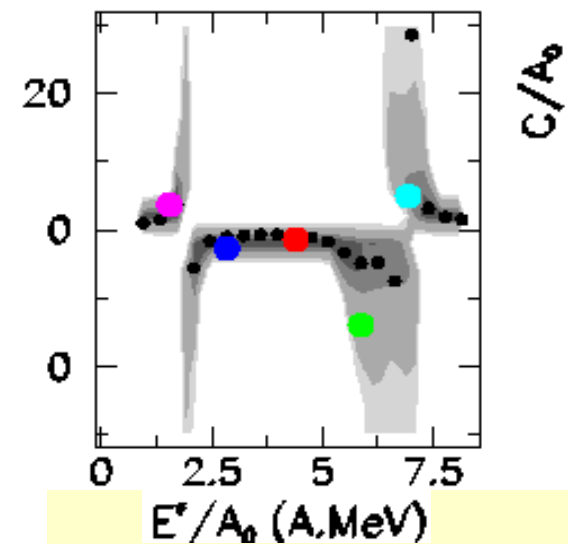
1-st order phase transition



Liquid-gas phase transition: is the game over?



Critical behavior inside the coexistence region

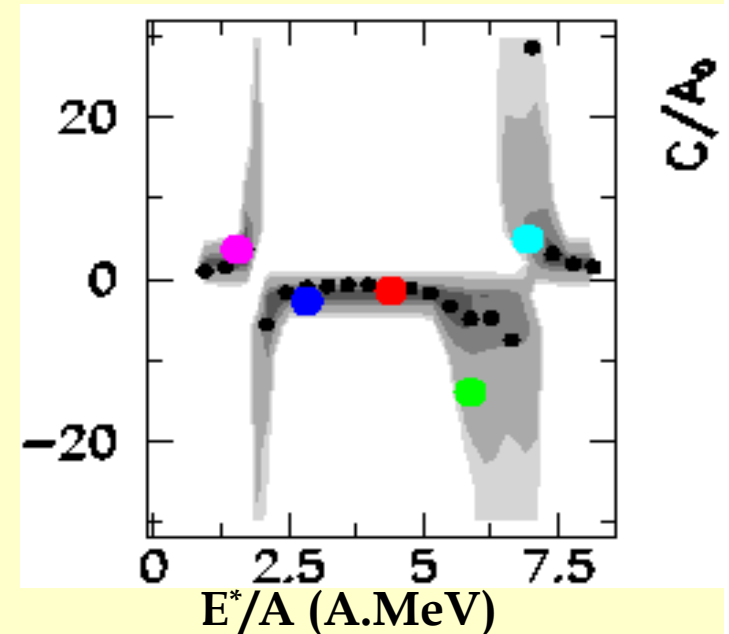


What is left for future measurements?

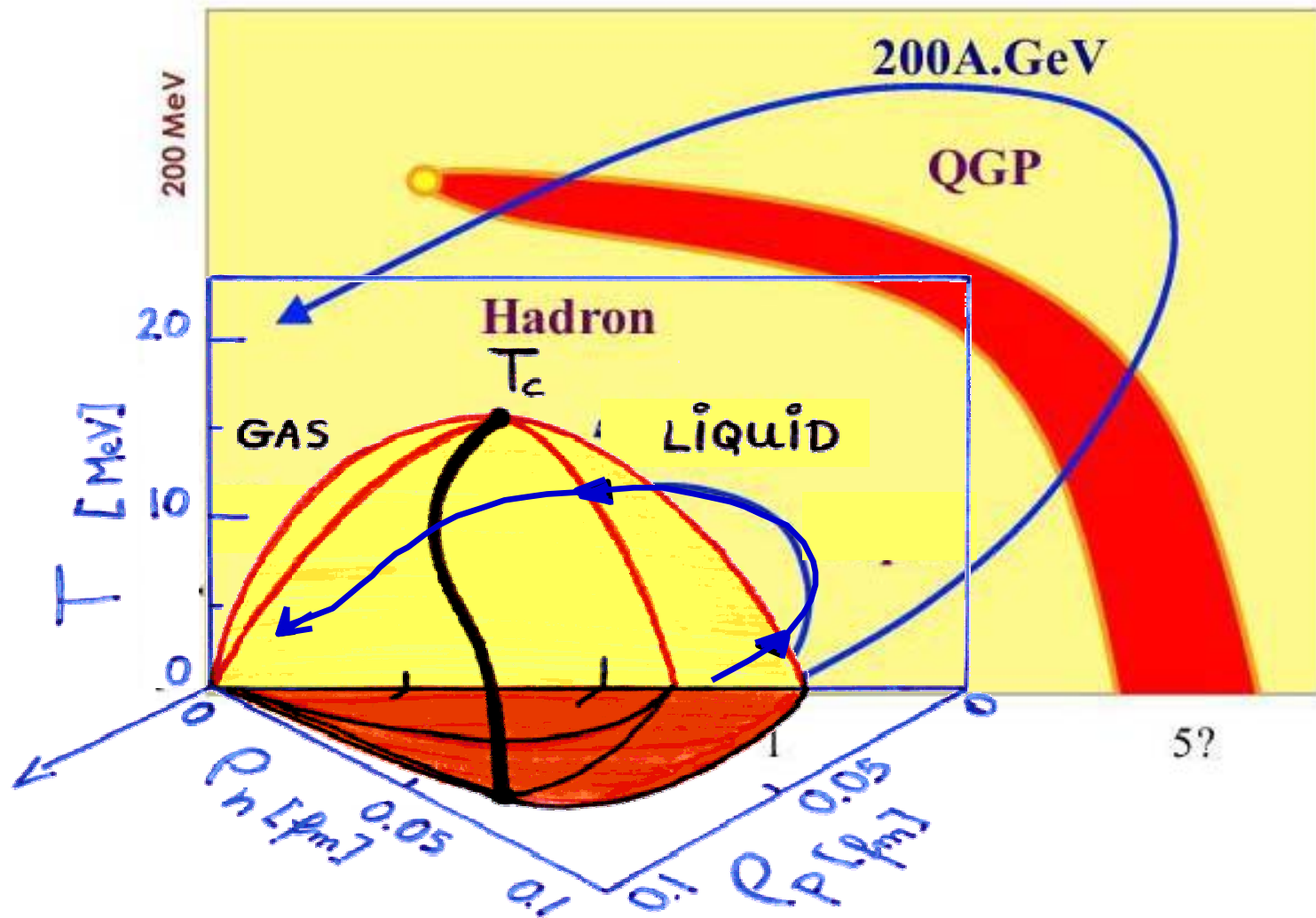
- A better quantitative nuclear metrology of hot nuclei
- Coincident experimental information are needed on:
 - critical partitioning of the system, fluctuations
 - calorimetric excitation energy
 - isotopic temperature
 - proximity of the decay products

4 π mass and charge detection !!

Multics NPA 2004



Multics	$E_1=2\pm0.3$	$E_2=6.5\pm0.7$
Isis	$E_1=2.5$	$E_2=7.$
Indra		$E_2=6.\pm0.5$



What is left for future measurements?

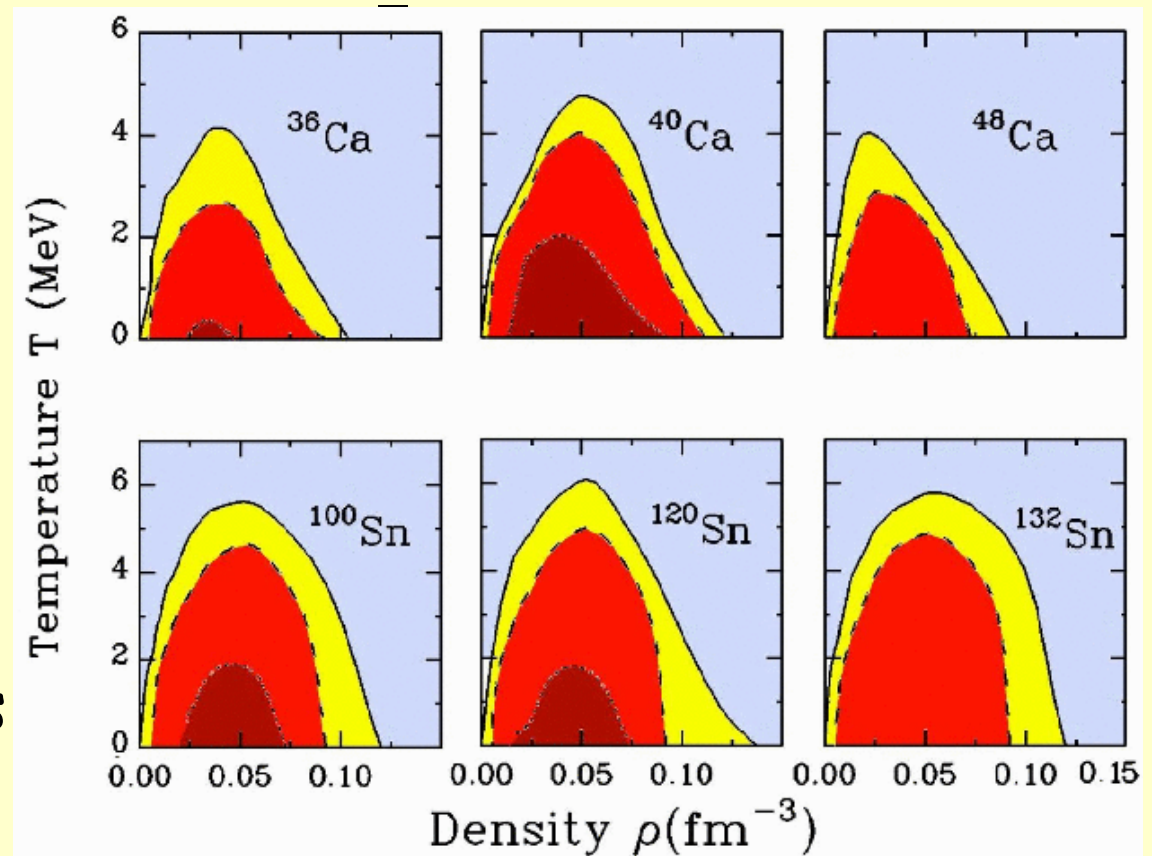
the 3-rd dimension of the EoS

2-nd generation devices and exotic beams are needed, to fully investigate the phase transition

by changing:

- the Coulomb properties
- the isospin content

of the fragmenting source

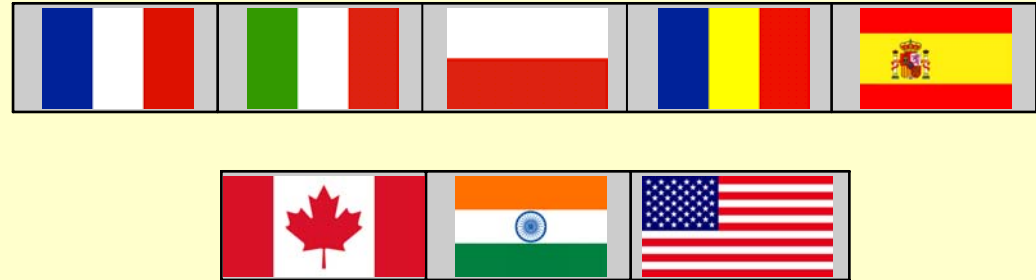
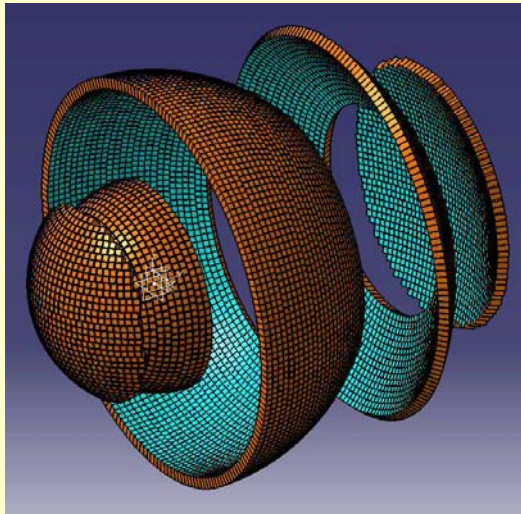


M.Colonna et al., PRL 88(2002) 122701

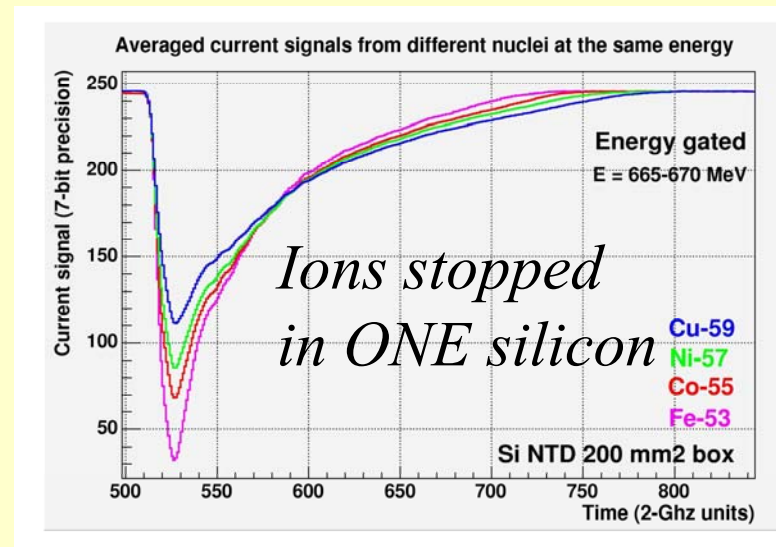
Instability growth time
100 fm/c (dashed/orange)
50 fm/c (dotted/red)

More asymmetric systems are less unstable

FAZIA : Four π A-Z Identification Array



- ~6000 telescopes
- Compactness of the device
- Ebeam from Barrier up to 100 A.MeV
- Telescopes: Si-ntd/Si-ntd/CsI
- Possibility of coupling with other detectors
- Complete Z (~70) and A (~50) id.
- Low-energy & identification threshold
- Digital electronics for pulse-shape id.





Systems and subsystems

S.K. Ma Statistical mechanics- Chap.6

Relazione fra la entropia di un sistema e le entropie dei suoi sottosistemi.

Dividiamo il sistema in due parti, 1 e 2 non interagenti

$$\begin{aligned}W_1(E_1) &= \exp(S_1(E_1)) \\W_2(E_2) &= \exp(S_2(E_2)) \\E &= E_1 + E_2\end{aligned}\tag{1}$$

$$\begin{aligned}W(E) &= \int_0^E \int_0^E dE_1 dE_2 W_1(E_1) W_2(E_2) \delta(E - E_1 - E_2) \\&= \int_0^E dE_1 W_1(E_1) W_2(E - E_1) \\&= \int_0^E dE_1 \exp(S_1(E_1) + S_2(E - E_1))\end{aligned}\tag{2}$$

$$P_1^E(E_1) = \frac{W_1(E_1)W_2(E-E_1)}{W(E)} \quad (3)$$

Lavoriamo attorno al massimo di $S_1(E_1) + S_2(E-E_1)$, che corrisponde all'equilibrio termico $T_1 = T_2 = T$, con

$$T_i^{-1}(E_i) = \frac{\partial S_i}{\partial E_i}(E_i). \quad (4)$$

L'espansione in serie di $S_1 + S_2$ nell'intorno del massimo porta a ²:

$$\begin{aligned} S_1(E_1) + S_2(E_2) &= S_1(\bar{E}_1) + S_2(\bar{E}_2) + \frac{1}{T}(E_1 - \bar{E}_1 + E_2 - \bar{E}_2) \\ &- \frac{1}{2} \left[\frac{1}{K_1}(E_1 - \bar{E}_1)^2 + \frac{1}{K_2}(E_2 - \bar{E}_2)^2 \right] + \dots \end{aligned} \quad (5)$$

con:

$$\frac{1}{K_i} = -\left(\frac{\partial^2 S_i}{\partial E_i^2}\right)_{E_i=\bar{E}_i} = \frac{1}{T^2 C_i} \quad (6)$$

dove

$$C_i = \frac{\partial \bar{E}_i}{\partial T} \quad (7)$$

Il massimo sia in $E_1 = \bar{E}_1$, $E_2 = \bar{E}_2 = E - \bar{E}_1$. Il termine corrispondente alla derivata prima di $S_1 + S_2$ va a 0, quindi:

$$\begin{aligned}
 P_1^E(E_1) &= \frac{\exp\left(S_1(\bar{E}_1) + S_2(E - \bar{E}_1) - \frac{1}{2} \left[\frac{(E_1 - \bar{E}_1)^2}{K_1} + \frac{(E_1 - \bar{E}_1)^2}{K_2} \right]\right)}{\int_0^E dE_1 \exp(S_1(E_1) + S_2(E - E_1))} \\
 &= \frac{\exp\left(S_1(\bar{E}_1) + S_2(E - \bar{E}_1) - \frac{1}{2} \left[\frac{(E_1 - \bar{E}_1)^2}{K} \right]\right)}{\exp\left(S_1(\bar{E}_1) + S_2(E - \bar{E}_1)\right) \int_0^E dE_1 \exp\left(\frac{-(E_1 - \bar{E}_1)^2}{2K}\right)}
 \end{aligned} \tag{8}$$

con

$$K = \frac{K_1 K_2}{K_1 + K_2} = \frac{1}{\frac{1}{K_1} + \frac{1}{K_2}} \tag{9}$$

$K > 0$ perche' la derivata seconda di $S_1 + S_2$ sia negativa (massimo), $K_i \neq 0$, ma non necessariamente $K_i > 0$. Sostituendo³:

$$z = E_1 - \bar{E}_1, \quad dz = dE_1, \quad \int_0^E \rightarrow \int_{-z_1^*}^{z_1^*} \rightarrow \int_{-\infty}^{\infty} = 2 \int_0^{\infty} \tag{10}$$

$$P_1^E(E_1) = \frac{\exp\left(\frac{-(E_1 - \bar{E}_1)^2}{2K}\right)}{\int_{-\infty}^{\infty} dz \exp\left(\frac{-z^2}{2K}\right)} = \frac{1}{\sqrt{2\pi K}} \exp\left(\frac{-(E_1 - \bar{E}_1)^2}{2K}\right) \tag{11}$$

Segue:

- $P_1^E(E_1)$ gaussiana normalizzata (area =1) $(\bar{E}_1, \sigma_1^2)$, con valore massimo $\frac{1}{\sqrt{2\pi K}}$,
- $\sigma_1^2 = K = \frac{1}{\frac{1}{K_1} + \frac{1}{K_2}} = \frac{1}{\frac{1}{C_1 T^2} + \frac{1}{C_2 T^2}} = T^2 \frac{C_1 C_2}{C_1 + C_2}$
- $\bar{C} = C_1 + C_2 = \frac{C_1^2}{(C_1 - \frac{1}{T^2})}$

Isotope analysis

Grand-canonical approach : $Y(N,Z) \propto \exp[(-F(N,Z) + \mu_n N + \mu_p Z)/T]$

T from **double ratios**: $\frac{Y(\text{He}^3)/Y(\text{He}^4)}{Y(\text{Li}^6)/Y(\text{Li}^7)}$

Isoscaling :

Two reactions :
 $\checkmark (Z/A)_1 \neq (Z/A)_2$
 $\checkmark T_1 = T_2 \quad V_1 = V_2$
 $\checkmark \alpha_q = (\mu_{q2} - \mu_{q1})/T$



$$R_{iso}^{(1,2)}(N,Z) = \frac{Y_2(N,Z)}{Y_1(N,Z)} \propto \exp[\alpha_n N + \alpha_p Z]$$

Approximate link to
symmetry energy C_{sym} :

$$\alpha_n = \frac{\Delta\mu_n}{T} = \frac{4C_{sym}}{T} \left([Z/A]_{(1)}^2 - [Z/A]_{(2)}^2 \right)$$

Isobaric ratio (for mirror nuclei) :

$$\frac{Y(N_1, Z_1)}{Y(N_2, Z_2)} \approx \frac{\rho_n}{\rho_p} e^{\Delta B/T}$$

Temperatura

Ipotesi: equilibrio

slope: effetti dinamici

doppio rapporto isotopico
si elimina la dipendenza
dalle proprietà chimiche

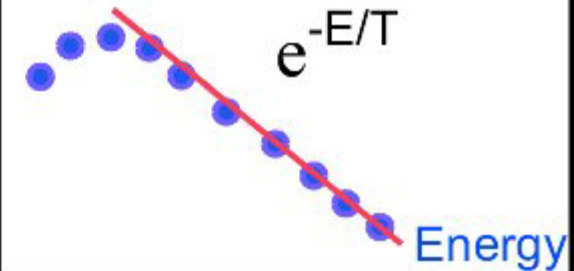
popolazione di stati eccitati

slope parameter

$p, d, t, \alpha, \dots$

$\pi, K, p, \dots$

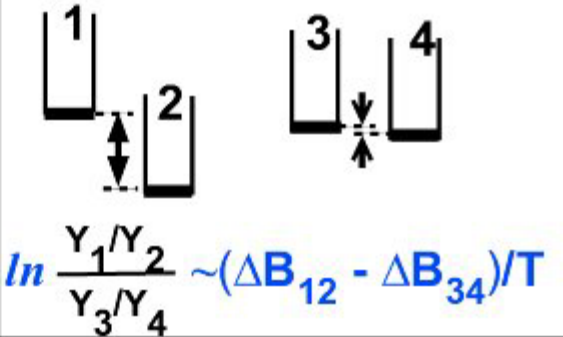
$\ln(dY/dE)$



isotopic composition

$(^3\text{He}/^4\text{He})/(^6\text{Li}/^7\text{Li})$

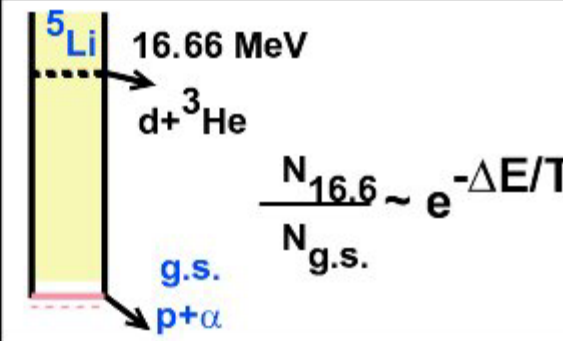
$(\pi^-/K^-)/(\Sigma^-/\Xi^-)$



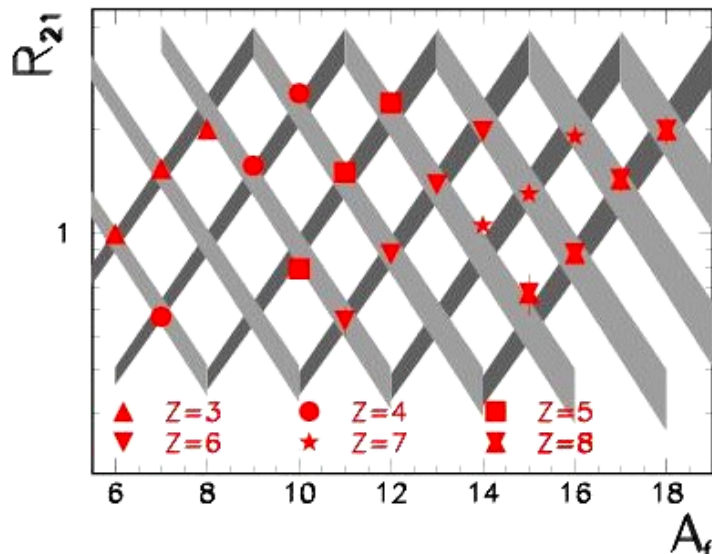
internal excitation

$^5\text{Li}_{16.66}/^5\text{Li}_{\text{gs}}$

Δ/N



Symmetry energy and free nucleon densities



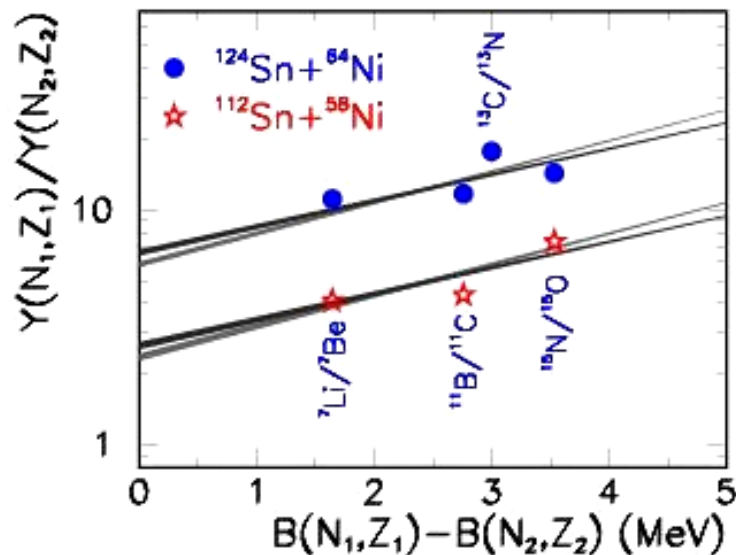
Isotopic ratio :

$$R_{21}(N, Z) = \frac{Y_{124\text{Sn}+^{64}\text{Ni}}(N, Z)}{Y_{112\text{Sn}+^{58}\text{Ni}}(N, Z)} \propto e^{\alpha N + \beta Z}$$

$$C_{\text{sym}} = \frac{\alpha T}{\left(\frac{Z}{A}\right)_1^2 - \left(\frac{Z}{A}\right)_2^2}$$

$$\alpha = 0.44 \pm 0.01$$

Symmetry Energy ~18-20 MeV

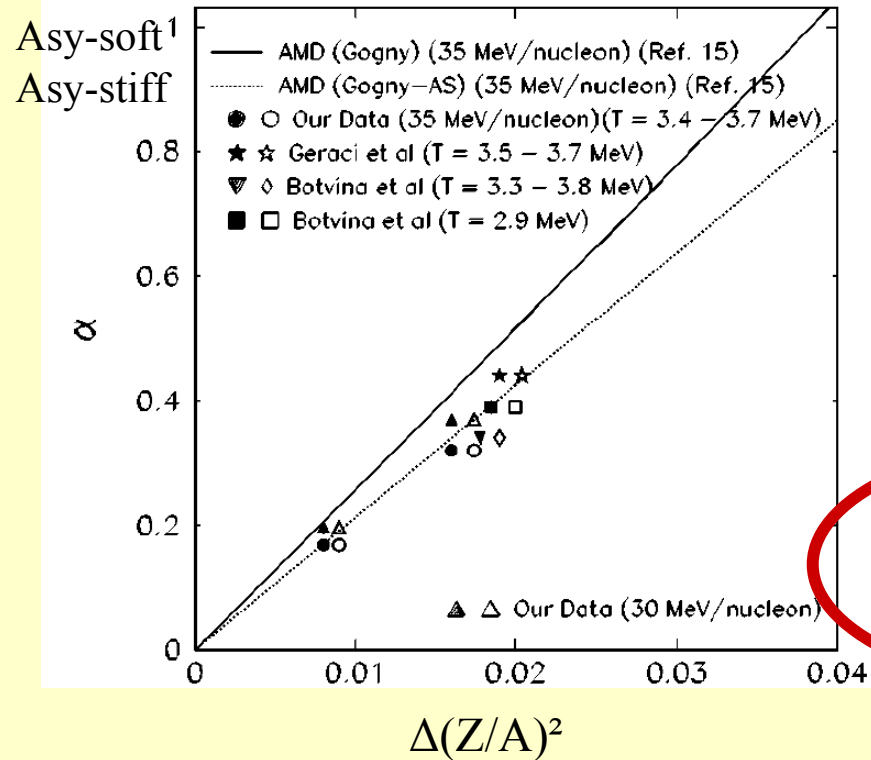


Isobaric ratio (for mirror nuclei):

$$\frac{Y(N_1, Z_1)}{Y(N_2, Z_2)} \approx \frac{\rho_n}{\rho_p} e^{\Delta B/T}$$

112,124Sn+58,64Ni 35 AMeV central collisions
CHIMERA-REVERSE Experiment

Extraction of symmetry energy



D. Shetty et al., P. R.C 70 (2004) 011601

E. Geraci et al., NPA 732 (2004)

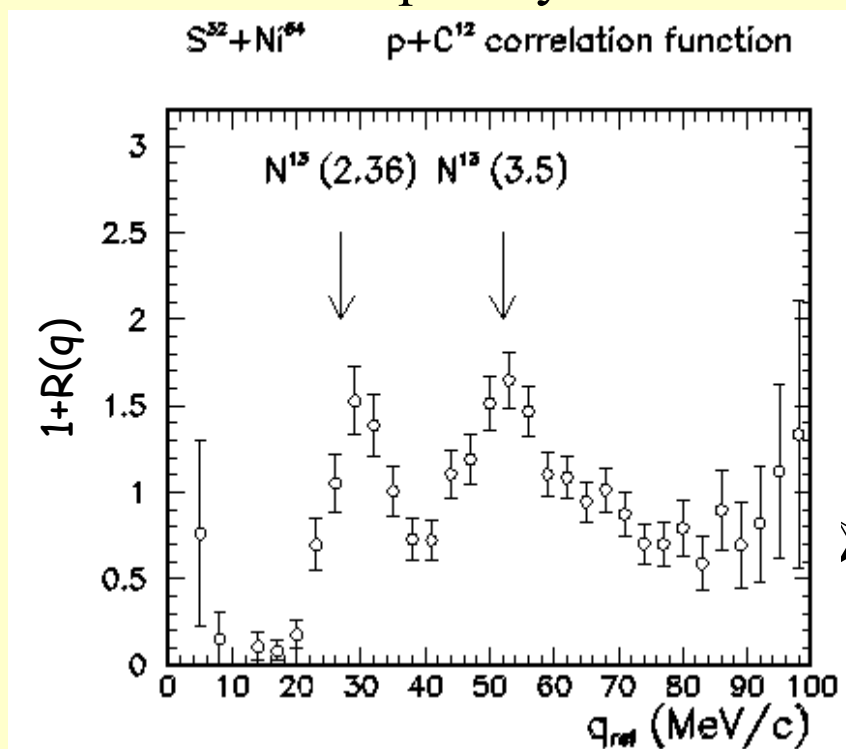
A. Botvina et al., PRC 65 (2002):

Sequential feeding?

Conclusions

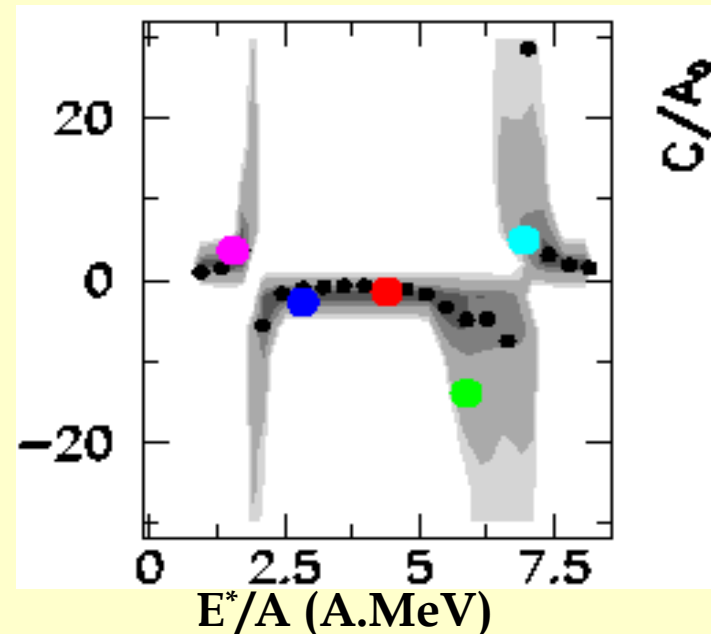
➤ The physics of hot nuclei: a unique laboratory

- for the thermodynamics of finite, charged, 2-component systems
- for a quantitative nuclear metrology
- for interdisciplinary connections



nucl-ex collaboration & garfield

Multics NPA 2004



Multics	$E_1 = 2 \pm 0.3$	$E_2 = 6.5 \pm 0.7$
Isis	$E_1 = 2.5$	$E_2 = 7.$
Indra		$E_2 = 6. \pm 0.5$

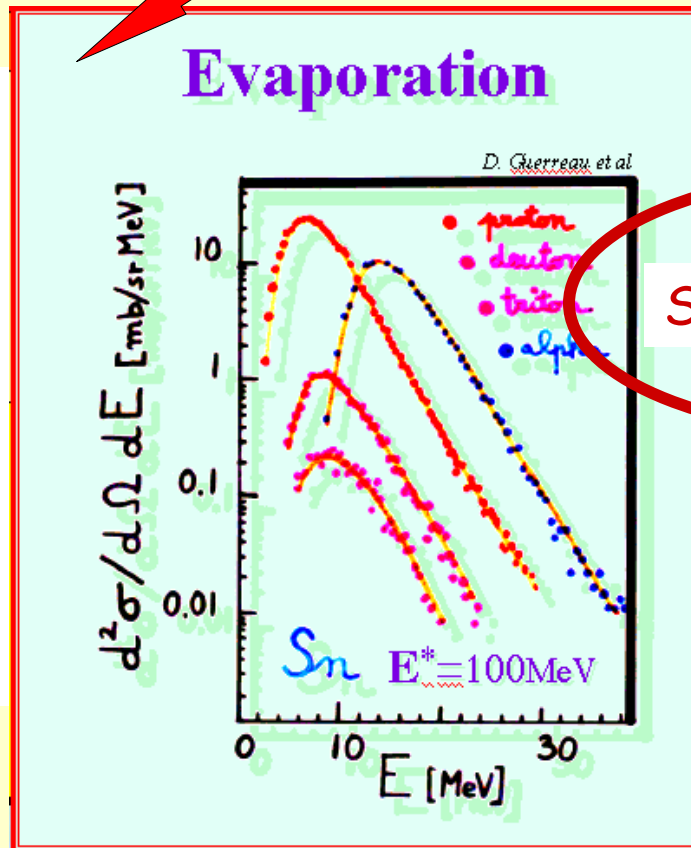
➤ We need:

- 4π mass and charge detection
- 20-50 A.MeV radioactive beams

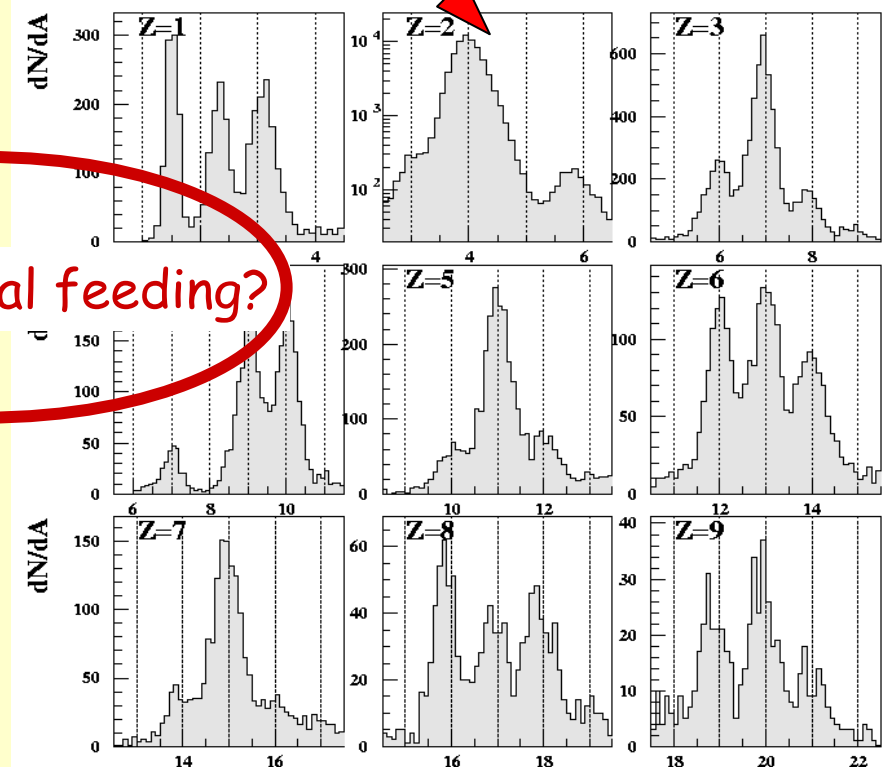
Temperature and caloric curve

For the *caloric curve* one needs to measure:

- Heavy residue (or QP)
- Slopes of 1-st chance l.c.p. energy spectra
- Isotopes (for double ratios)



Sequential feeding?



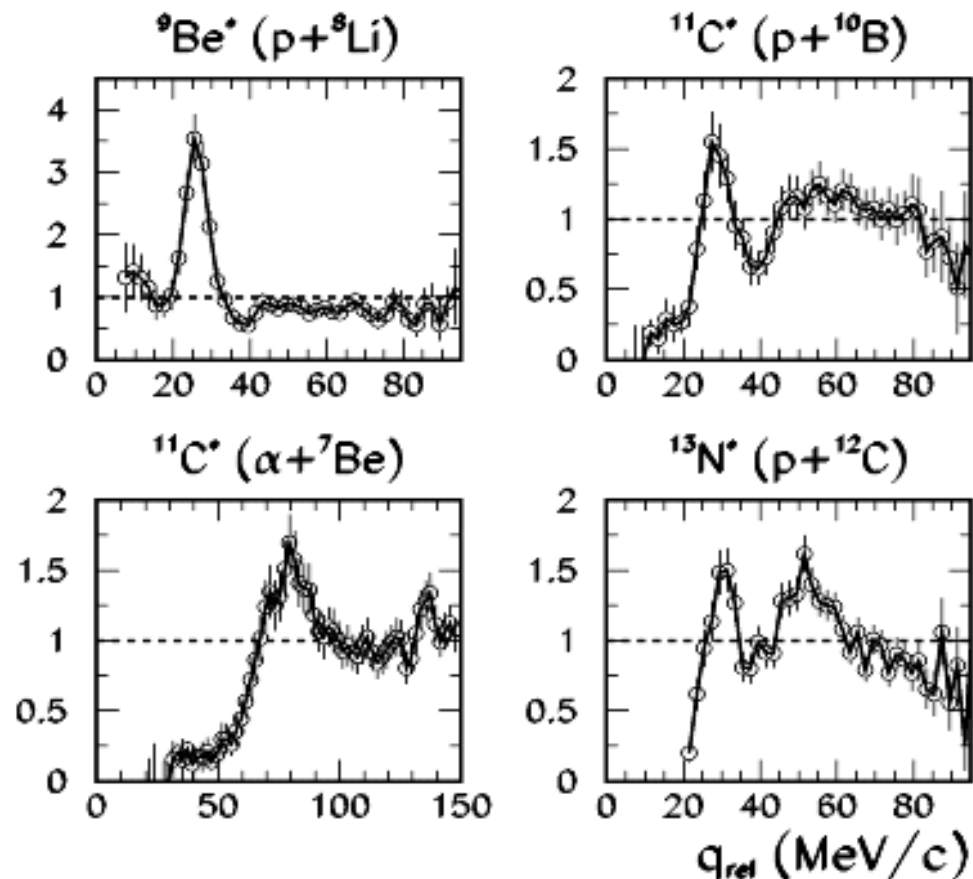
N. Le Neindre et al, NIM A490 (2002) 251

Experiments with n-rich/poor systems

$^{32}\text{S}+^{58,64}\text{Ni}$ 14.5 AMeV 3-IMF events

Observed 35 resonances, from He^4 (d+d) to Ne^{20} ($\alpha+\text{O}^{16}$)

A rough calculation of "feeding correction" through correlation functions suggests an increase of T by 0.5 MeV for few % of decrease in the He^4 yield



Before drawing conclusions on temperature, densities:

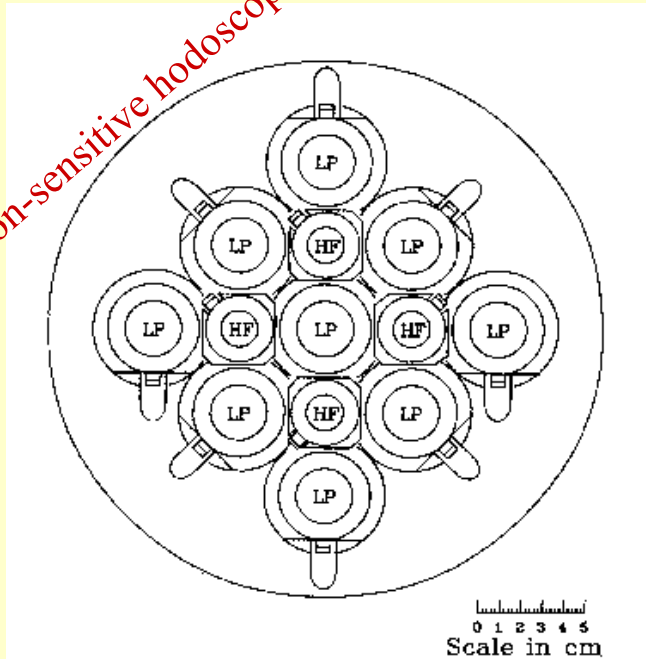
➤ Isotope emission time scales have to be checked through correlation functions

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Resonance spectroscopy

t- α correlation function (Li^{7*})

Position-sensitive hodoscope



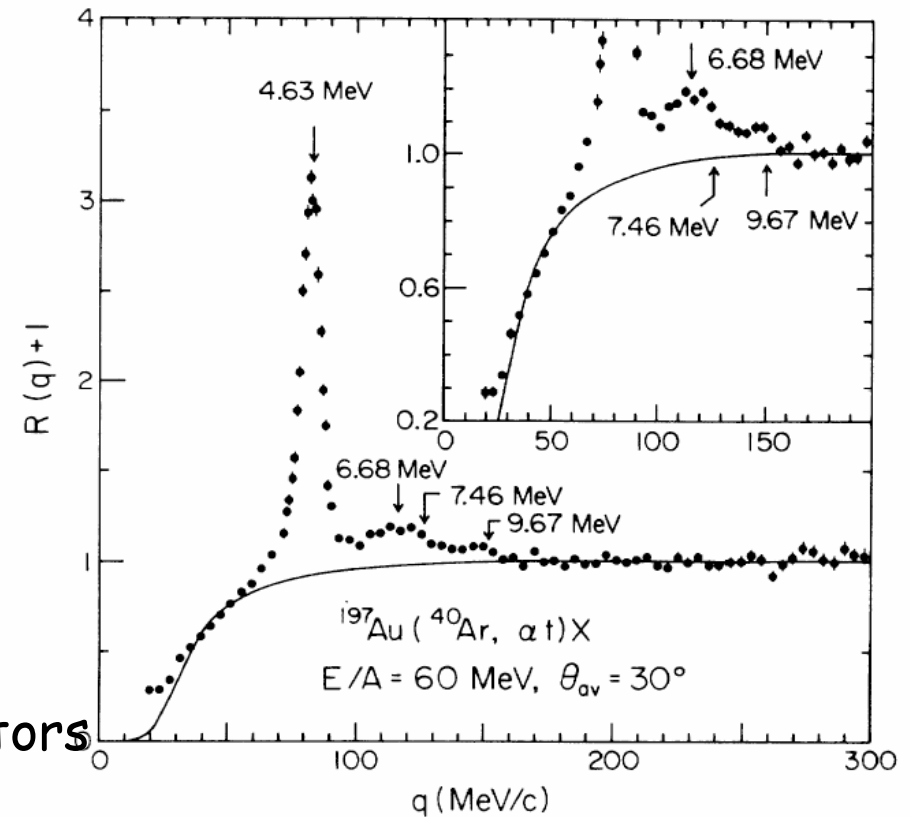
m =multiplicity, N =number of detectors

$$\cdot \varepsilon(m) = \varepsilon(1)^m$$

$$\cdot P(\text{double}) = (m-1)/(2N)$$

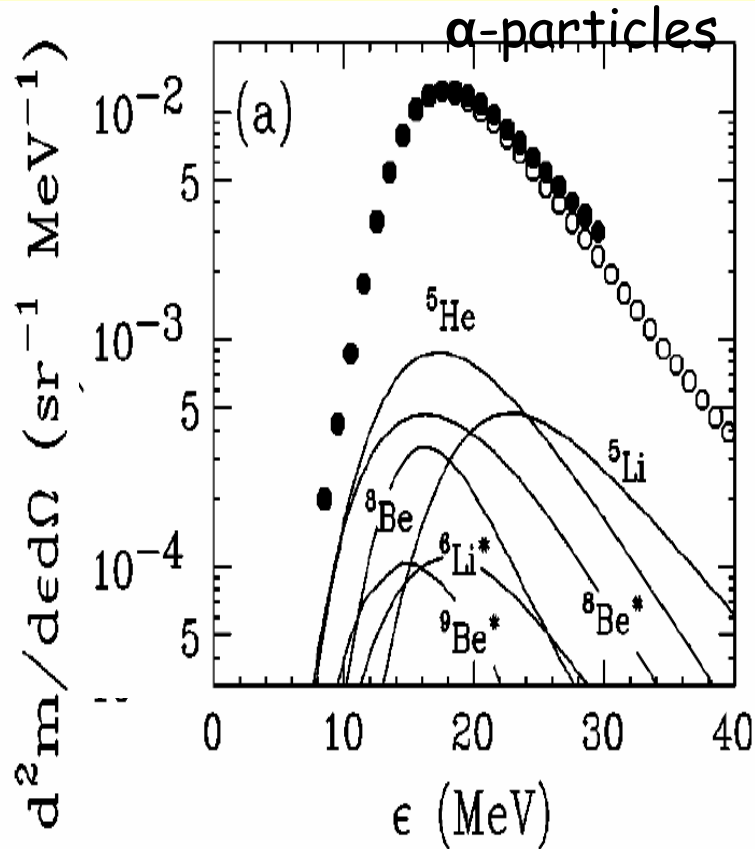
A reasonable compromise is $P(\text{double}) < 5\%$

For $m=3$ $N=10$



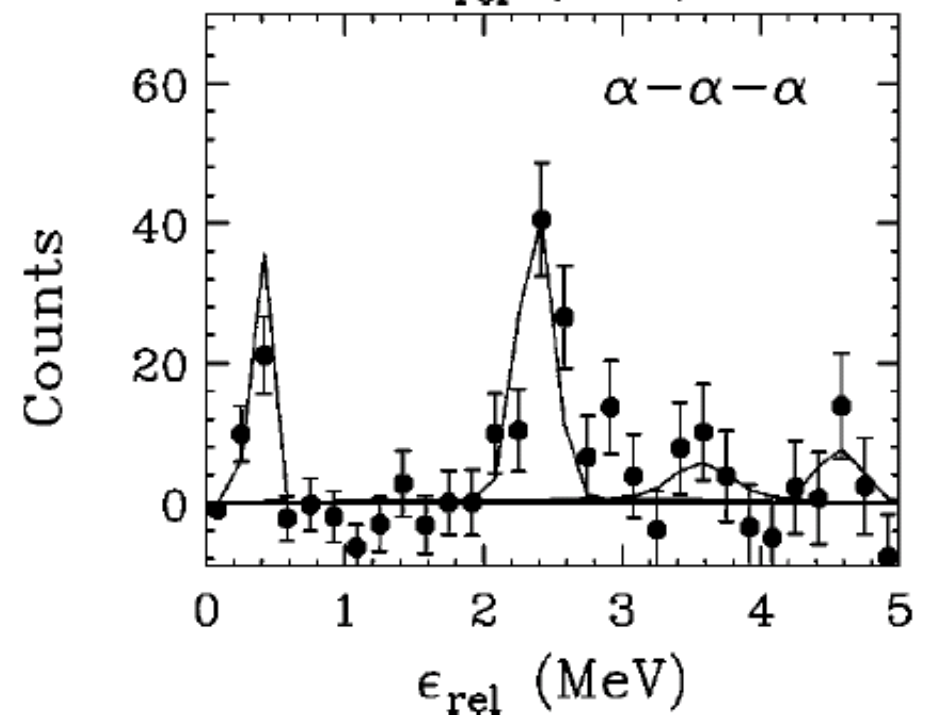
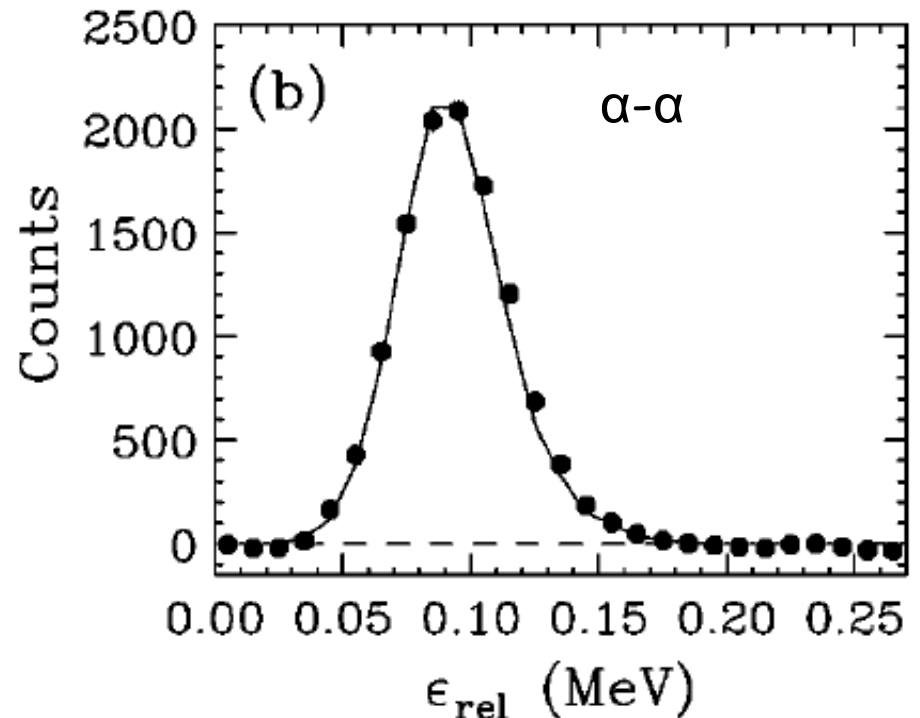
Pochodzalla et al., PRC35 (1987)1695

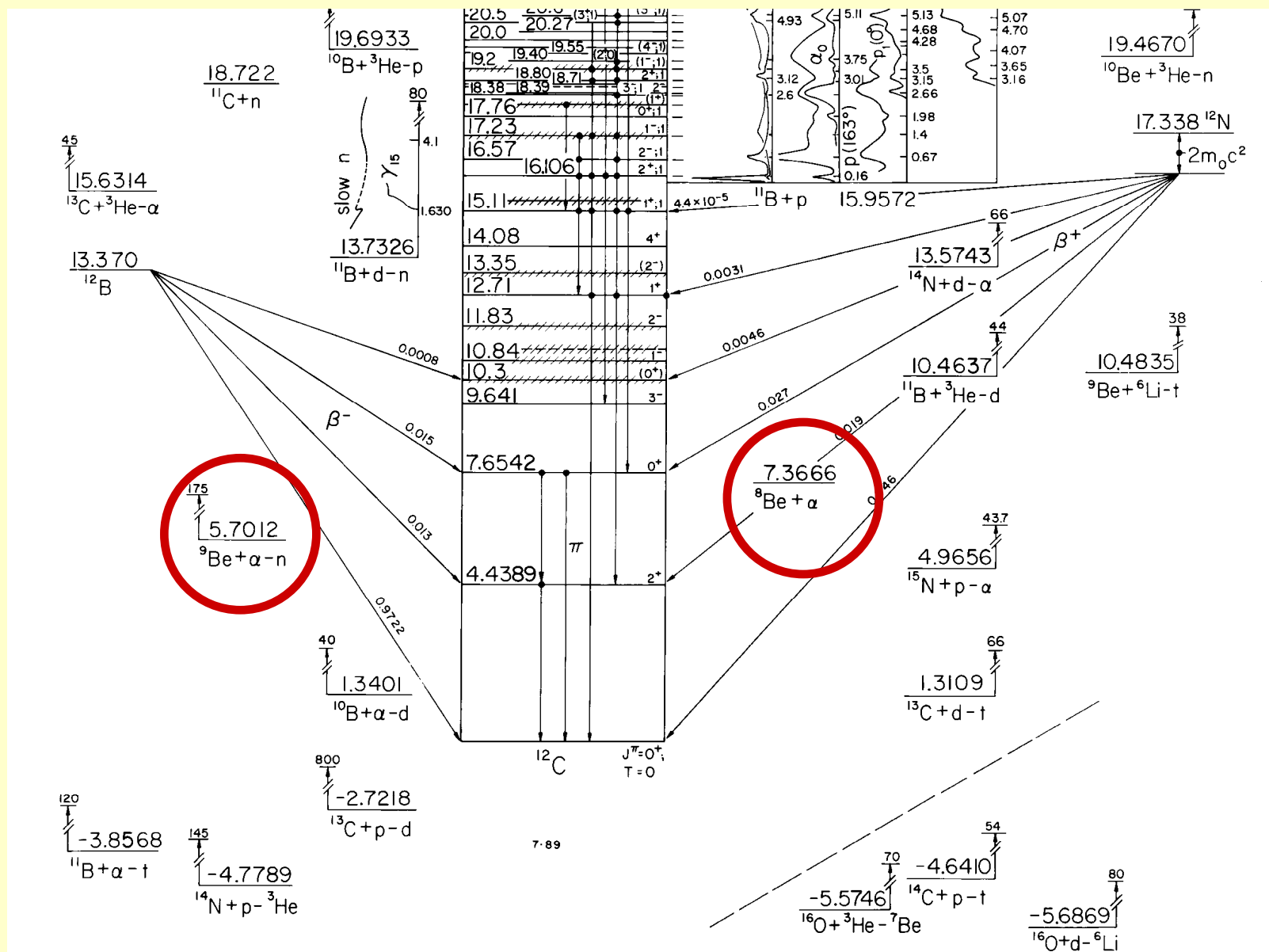
Why many-body correlations?

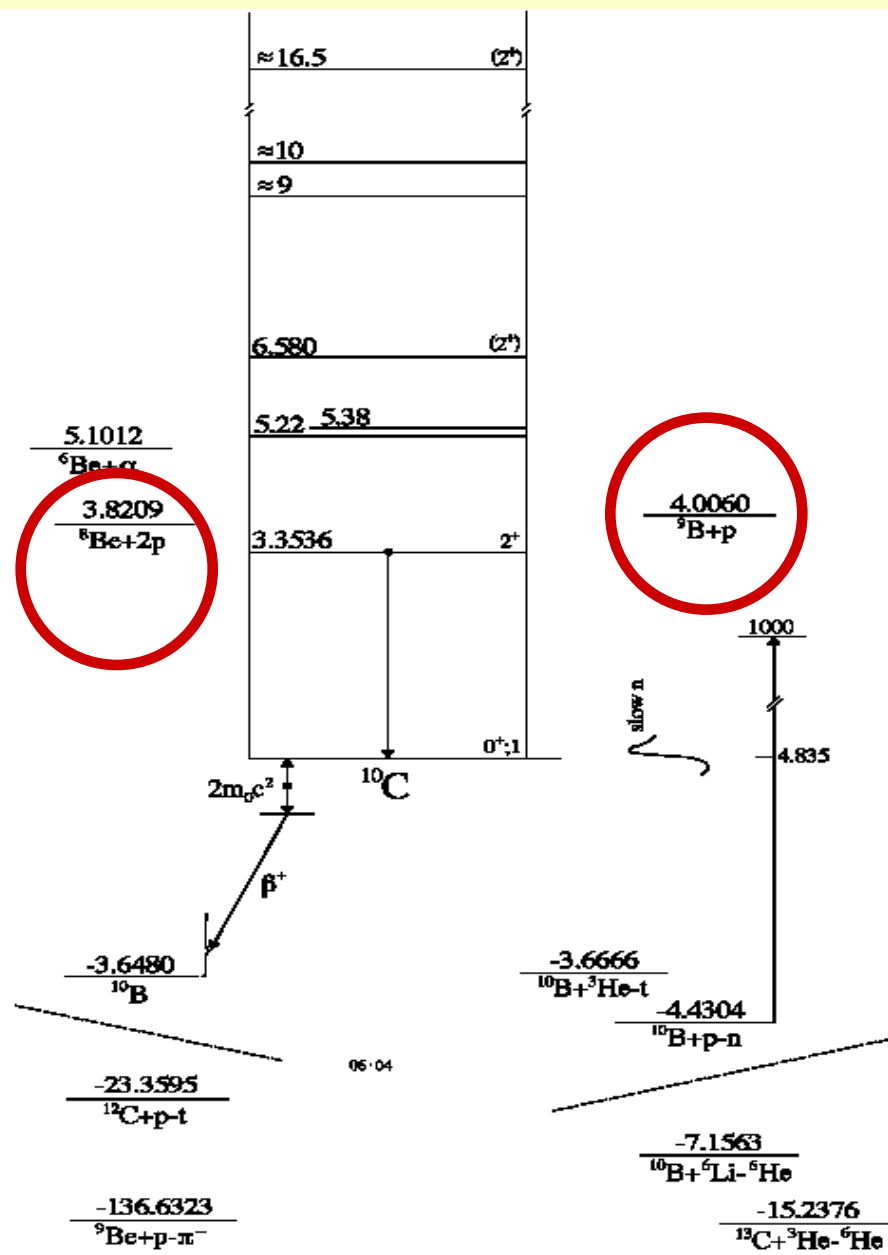


R.J. Charity et al., PRC63 024611
 $^{60}\text{Ni}+^{100}\text{Mo}$ 11 A.MeV

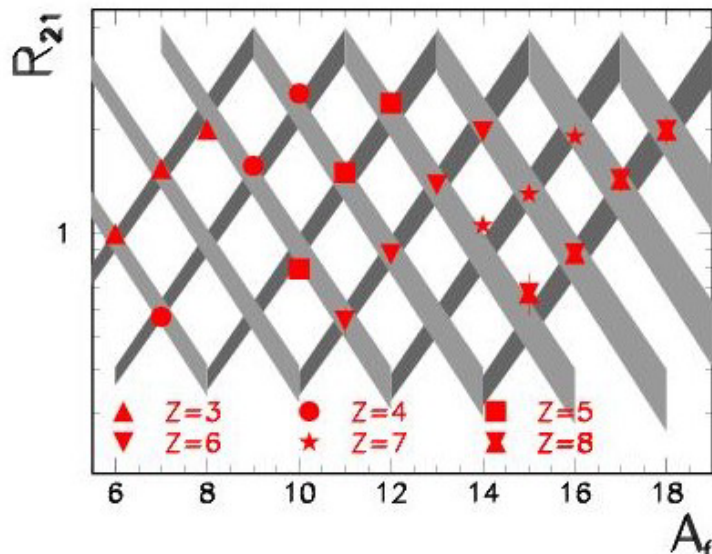
$\Delta\theta \approx 0.6^\circ \rightarrow$ high granularity
 but in a limited angular coverage
 & not HR full identification







$^{112}\text{Sn}+^{58}\text{Ni}$ and $^{124}\text{Sn}+^{64}\text{Ni}$ 35 AMeV central collisions CHIMERA-REVERSE Experiment



Isotopic ratio :

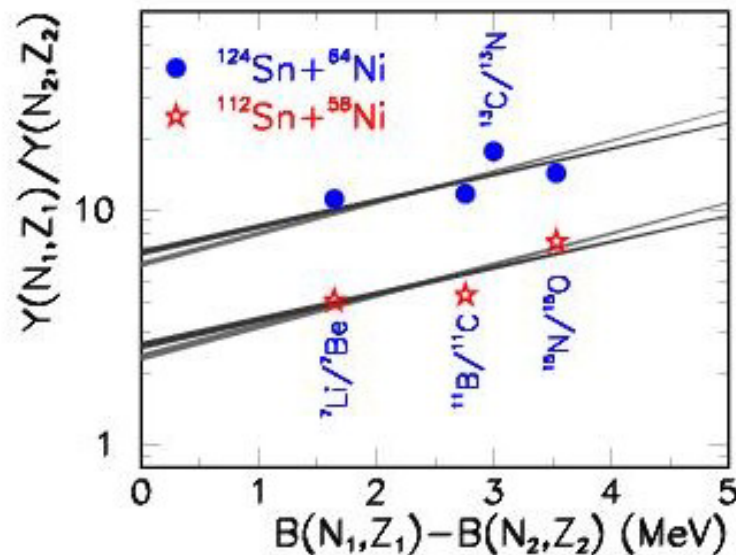
$$R_{21}(N, Z) = \frac{Y_{^{124}\text{Sn}+^{64}\text{Ni}}(N, Z)}{Y_{^{112}\text{Sn}+^{58}\text{Ni}}(N, Z)} \propto e^{\alpha N + \beta Z}$$

$$C_{\text{sym}} = \frac{\alpha T}{\left(\frac{Z}{A}\right)_1^2 - \left(\frac{Z}{A}\right)_2^2}$$

$$\alpha = 0.44 \pm 0.01$$

C_{sym} = Symmetry Energy ~18-20 MeV

D.Shetty et al., P. R.C 70 (2004) 011601



Isobaric ratio (for mirror nuclei) :

$$\frac{Y(N_1, Z_1)}{Y(N_2, Z_2)} \approx \frac{\rho_n}{\rho_p} e^{\Delta B/T}$$